

Auxiliary Material for

Hot upwelling conduit beneath the Atlas Mountains, Morocco

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Introduction

To generate the seismic waveform synthetics, we applied a 2-D, staggered grid Finite Difference scheme with implementation of accurate point source mechanism [Li *et al.*, 2014]. We used a grid size of 200 m with time step of 0.01 s, which makes the calculation accurate up to 2 Hz. In order to test the trade-off between the velocity perturbation and the depth extent of the low velocity anomaly, we conducted resolution tests to evaluate the different parameters.

The original tomography model [Bezada *et al.*, 2013] predicts travel times well (Fig. S1). However, it is too smooth to predict the sharp travel time jump and secondary arrivals. For a weaker LVA model ($\delta V_s = -4\%$) (Fig. S2), the travel times are shifted to earlier arrivals and the amplitudes are smaller. In contrast, a stronger model ($\delta V_s = -8\%$) has later arrivals and the second arrivals are late and stronger than those in the data. If the length of the LVA is 100 km (Fig. S3), a much stronger shear velocity reduction ($\sim -10\%$) is needed to explain the travel time change from station PM21 and PM22. However, such a model predicts much later second arrivals than observed in the data. In contrast, if we extend the LVA to 300 km depth (Fig. S4), the reduced shear velocity anomaly matches

the travel time but produces earlier and weaker second arrivals that do not compare to the observations.

The width of the LVA primarily affects the variation of the waveform amplitude near 4° distance. If the anomaly is wider (~ 110 km, Fig. S5), the region with large amplitude is broad. A narrower anomaly will generate a confined large amplitude zone in the synthetic waveforms that are not observed in the data (Fig. S5). In Fig. S6, we test a high velocity anomaly structure instead of the LVA. Although this model predicts the travel time jump from station PM21 to PM22, it generates very weak secondary late arrivals. By comparing the travel times, amplitudes, and the arrival times between the second and first pulses for all 15 stations along the profile for the deep focus event and the *S-SKS* differential times, we found that a 200 km long and 80 km wide LVA with shear wave velocity perturbation of -6% (Fig. 2) best explained the *S* data. Such an anomaly also explains the observed change in the differential travel time between *S* and *SKS* for teleseismic events. From station PM21 to PM22, while *SKS* travel times stay constant, the sharp *S* travel time delay results *S-SKS* differential travel time increase sharply (Fig. S7).

The differential *P* travel times for teleseismic events display variations of less than ~ 1 sec (Fig. S8). The *P* records of the deep Granada event do not have the second arrivals as in the *S* data (Fig. S9), which suggests the *P* anomaly is much weaker than the *S* anomaly. We tested models with the same geometry as the *S* model and different *P* velocity perturbations. To match the amplitude variation across the profile, a model with *P* velocity drop of 3-4% is preferred (Fig. S9).

In Equation 1, V_o is the ambient P or S velocity for AK135 at the depth of 140 km. α is a constant value of 0.15 [Goes *et al.*, 2012], and Q is the attenuation quality factor. We adopt the velocity derivatives for the shallow mantle from Goes *et al.* [2000]. We have $\partial V_s/\partial T = -0.3791 \text{ m}\cdot\text{s}^{-1}\cdot\text{K}^{-1}$ and $\partial V_p/\partial T = -0.5364 \text{ m}\cdot\text{s}^{-1}\cdot\text{K}^{-1}$. The melt contribution is added using $\partial V_s/\partial \phi = -2.04V_{s0}$ and $\partial \ln V_p/\partial \phi = -1.23V_{p0}$ assuming the unrelaxed state [Kreutzmann *et al.*, 2004]. We chose Q_s from the partly empirical model defined in Goes *et al.* [2012] as

$$Q_s = A_g \left(\frac{C_{OH}^{ref}}{C_{OH}} \right)^s f^{\alpha_g} \exp \left(\frac{\alpha_g \gamma T_m}{T_{ref} + \Delta T} \right)$$

The parameters are $A_g = 0.1$, $\alpha_g = 0.15$, $\gamma = 38$, $s = 0.37$. T_m is the melting temperature. In the calculation, we use the dry solidus temperature following Hirschmann [2000] with $T_m = 1890 \text{ K}$ at 4.5 GPa (140 km depth). T_{ref} is the ambient temperature. C_{OH}^{ref} is the reference hydration of the mantle as 1000 H/10⁶Si. Because the effects of high water content on velocity are similar as those of partial melting [Karato, 2004], it is difficult to distinguish them from velocities alone. Thus, we assume a dry mantle with $C_{OH} = 50 \text{ H}/10^6\text{Si}$. Q_s is also affected by the seismic frequency f . Here, the frequency is $\sim 0.25 \text{ Hz}$ for S waves. For simplicity,

$$Q_p = Q_s \frac{3}{4} \left(\frac{V_p}{V_s} \right)^2 \quad [\text{Goes } et al., 2000].$$

1. fs01.pdf (Figure S1) Comparison between the *S* data and the synthetics for *P* tomography scaled by 1.8.
2. fs02.pdf (Figure S2) Comparison between the *S* data and the synthetics for the Granada earthquake. The model is the same as that in Fig. 3 but with different δV s. From left to right: -6% , -4% , -8% .
3. fs03.pdf (Figure S3) (A) Comparison between the *S* data and the synthetics for the Granada earthquake. The LVA model has the length of 100 km and different δV s. From left to right: -8% , -10% , -12% . (B) travel time and amplitude of the data (black triangles) and the synthetics (lines). The black, red, and blue lines are predictions from the low velocity anomaly with velocity perturbation of -8% , -10% , and -12% , respectively.
4. fs04.pdf (Figure S4) (A) Comparison between the *S* data and the synthetics for the Granada earthquake. The LVA model has the length of 300 km and different δV s. From left to right: -3% , -4% , -5% . (B) travel time and amplitude of the data (black triangles) and the synthetics (lines). The black, red, and blue lines are predictions from the low velocity anomaly with velocity perturbation of -3% , -4% , and -5% , respectively.
5. fs05.pdf (Figure S5) (A) Comparison between the *S* data and the synthetics for the Granada earthquake. The LVA model has the length of 200 km, δV s of -6% and different width. From left to right: 80 km, 50 km, 110 km. (B) travel time and amplitude of the data (black triangles) and the synthetics (lines).
6. fs06.pdf (Figure S6) Prediction from a high velocity model, indicated by blue box in (A). The preferred LVA model is outlined by the orange line. (B) Comparison between the *S* data and the synthetics for the high velocity model.
7. fs07.pdf (Figure S7) Prediction for the teleseismic *S*-*SKS* differential travel time for an event from the northeastern direction. The red triangles are data. The black line is the prediction from our preferred LVA model in Fig. 3.

8. fs08.pdf (Figure S8) P differential travel time between stations PM22 and PM21 against the back-azimuths of the events from station PM21. The star represents the upgoing P phase of the deep Granada earthquake. The radius is the incident angle from horizontal at the depth of 120 km assuming the 1-D AK135 reference .
9. fs09.pdf (Figure S9) (A) Comparison between the P data and the synthetics for the Granada earthquake. The LVA model has identical shape as that in Fig. 3 but different δV_p . From left to right: -3% , -4% , -5% . (B) travel time and amplitude of the data (black triangles) and the synthetics (lines).
10. fs10.pdf (Figure S10) Histograms of estimates of T_{ref} , ΔT , ϕ , dV_p/V_{p_0} , and dV_s/V_{s_0} from the Monte Carlo simulations assuming a hydrated mantle with $C_{OH} = 300 \text{ H}/10^6 \text{ Si}$ and 120 K lower melting temperature than that for the dry solidus.

Reference

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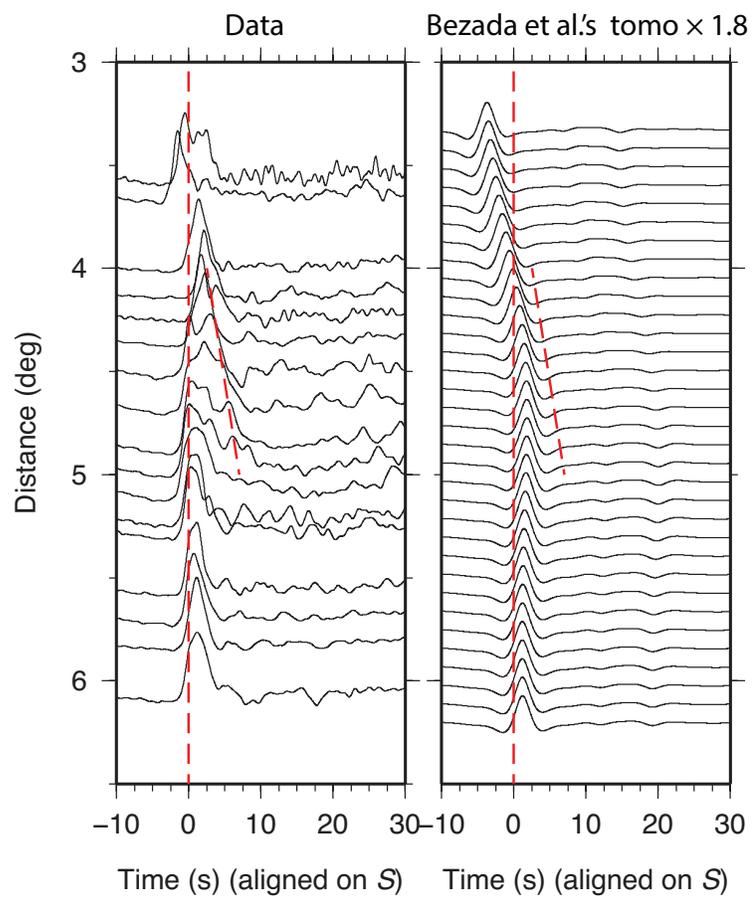


Figure S1

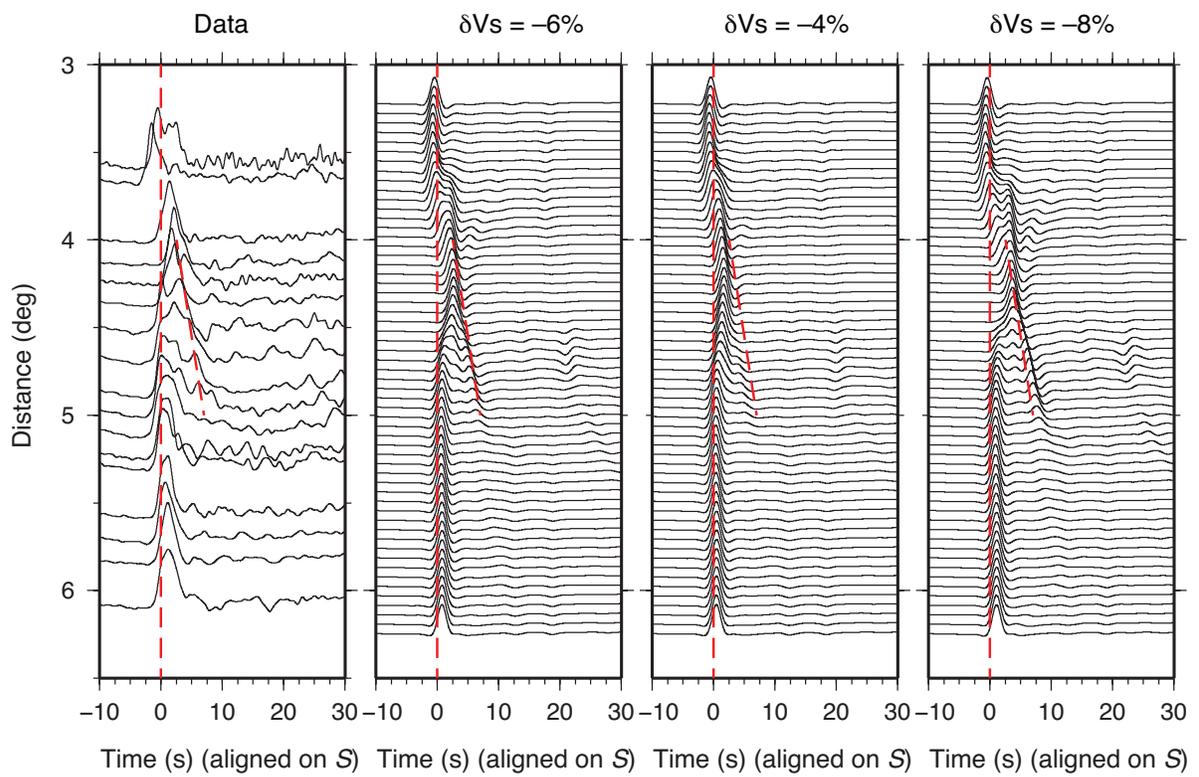


Figure S2

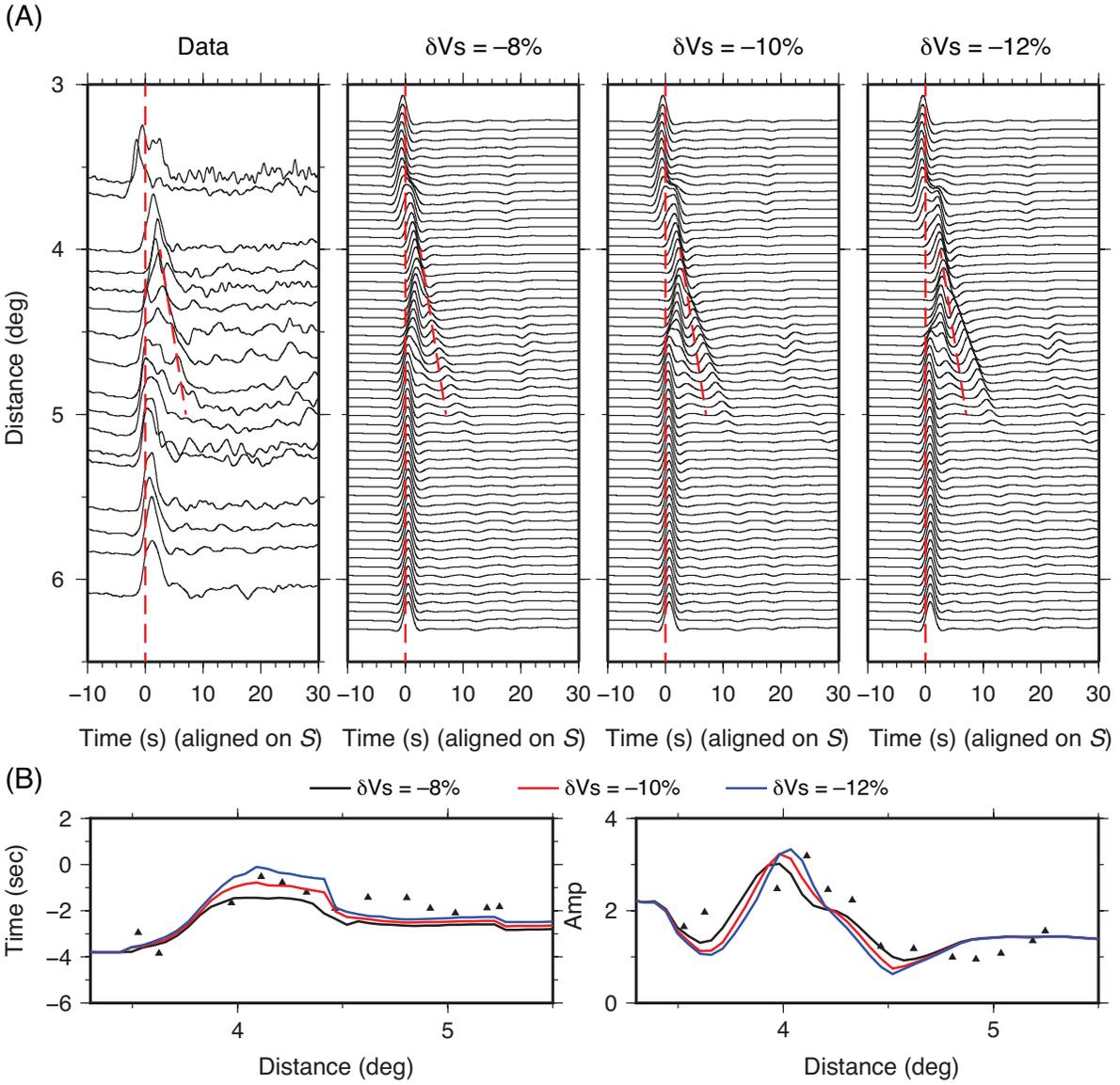


Figure S3

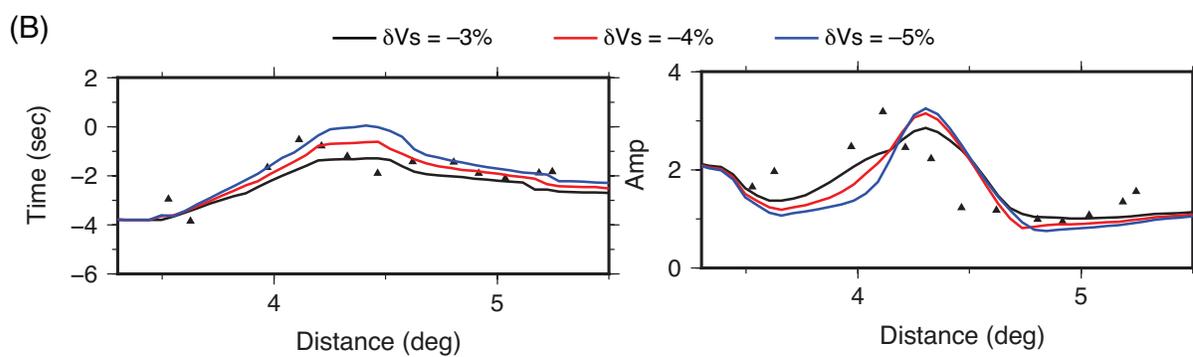
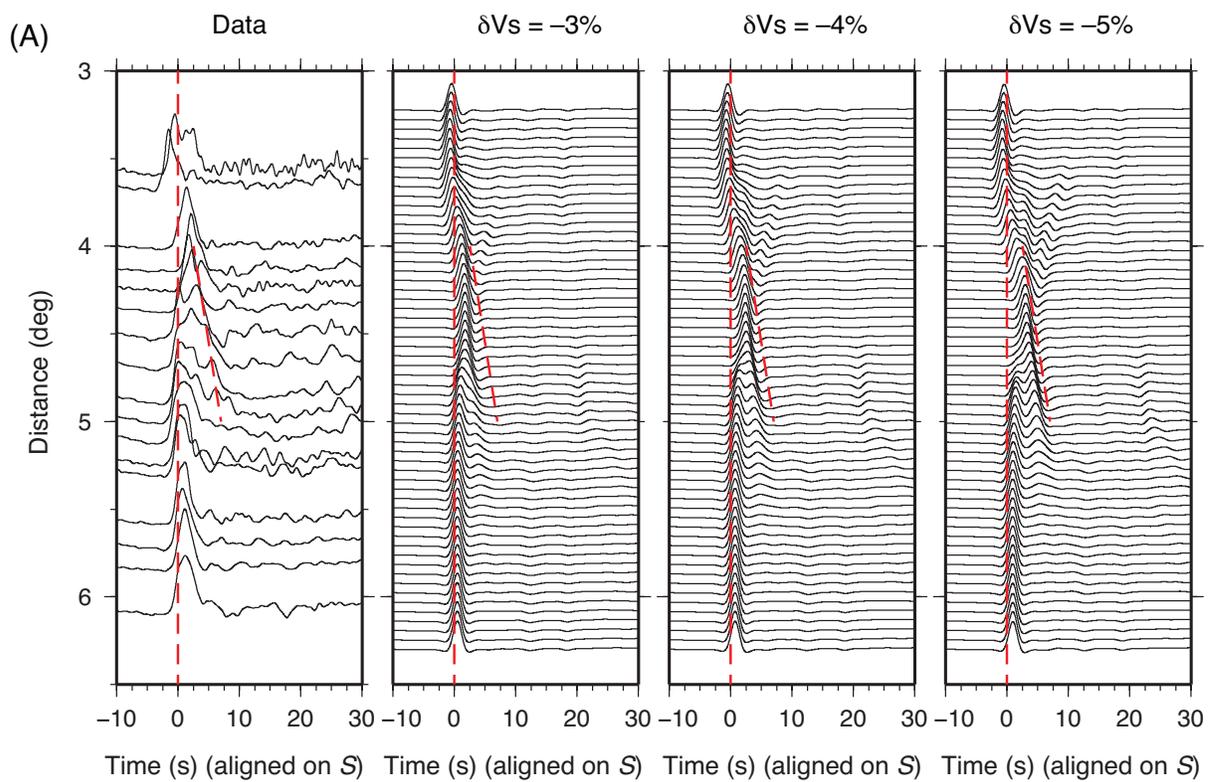


Figure S4

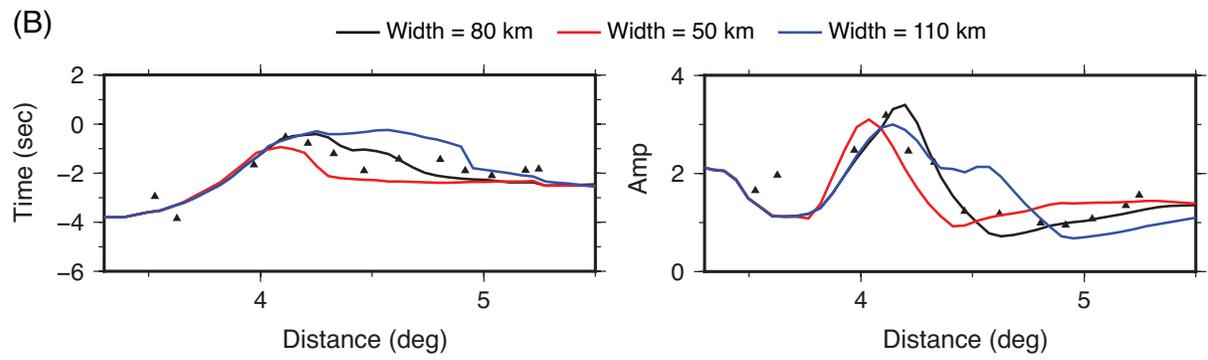
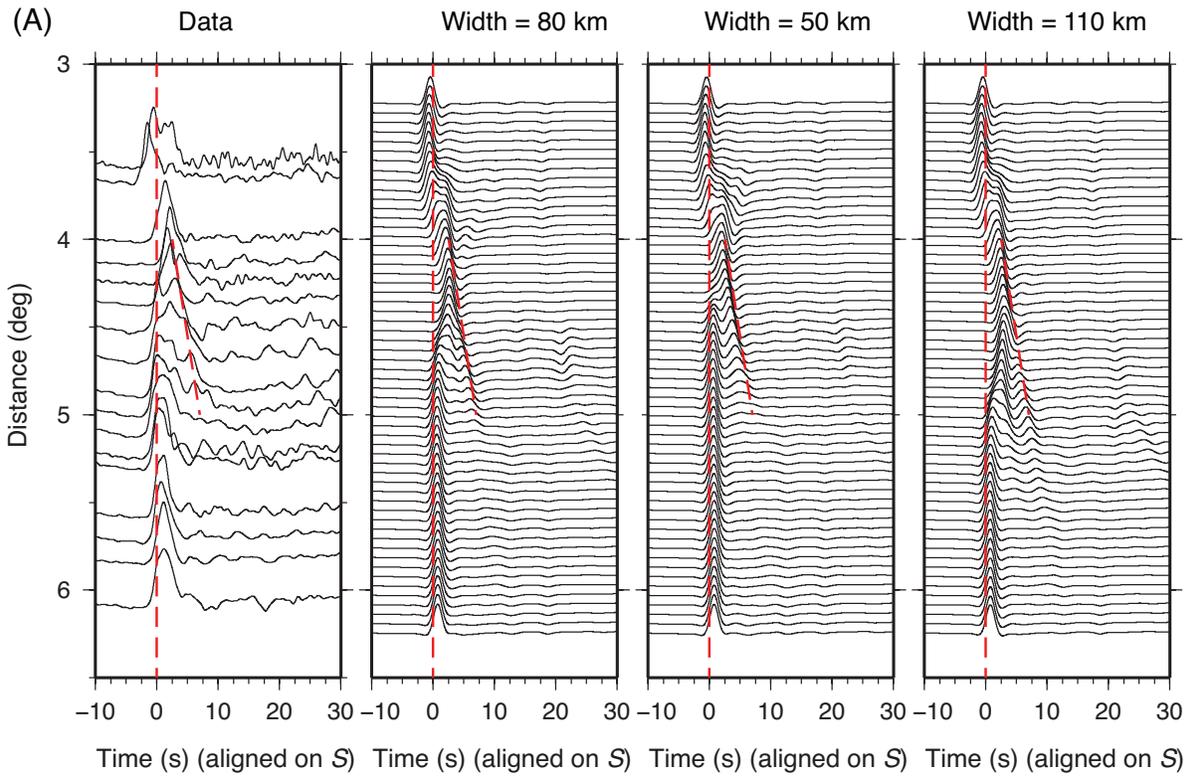


Figure S5

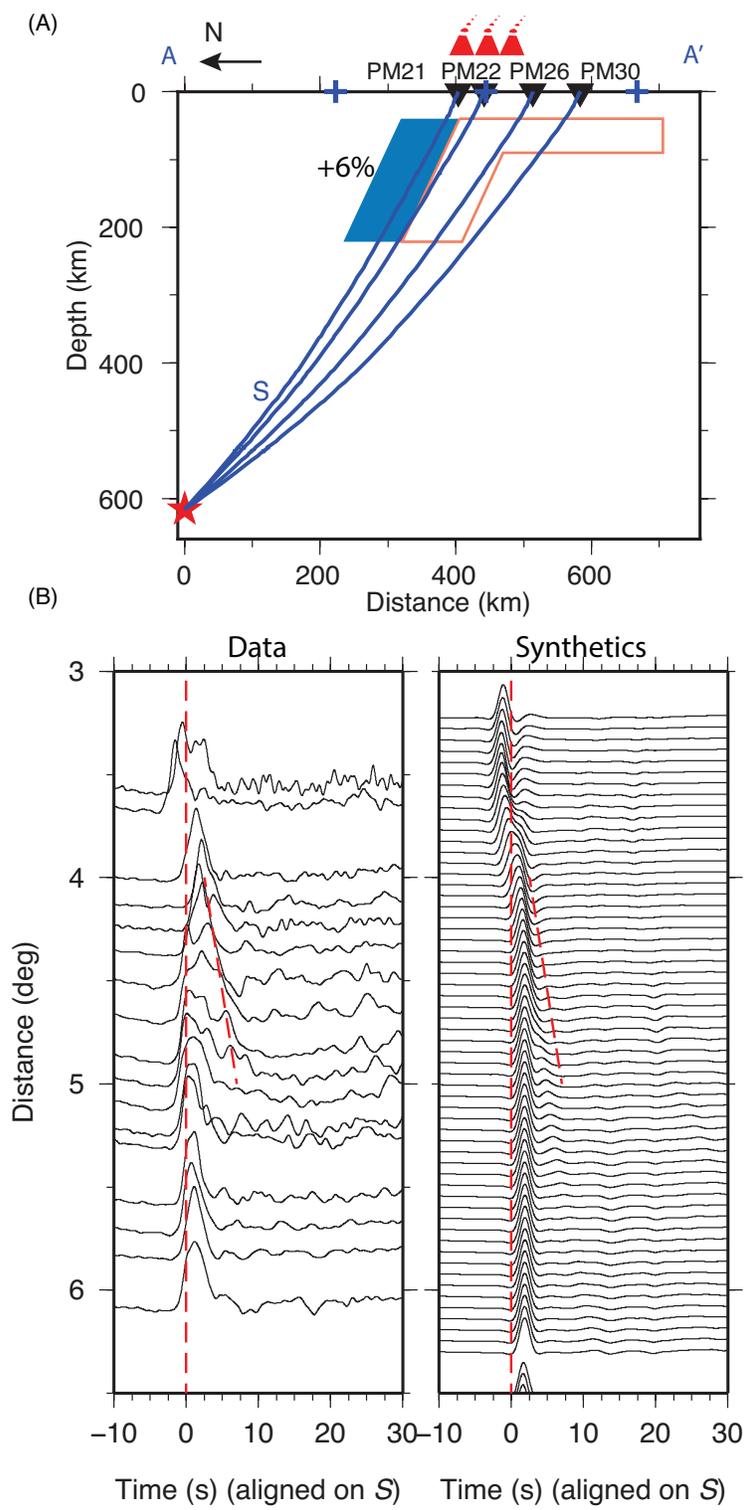


Figure S6

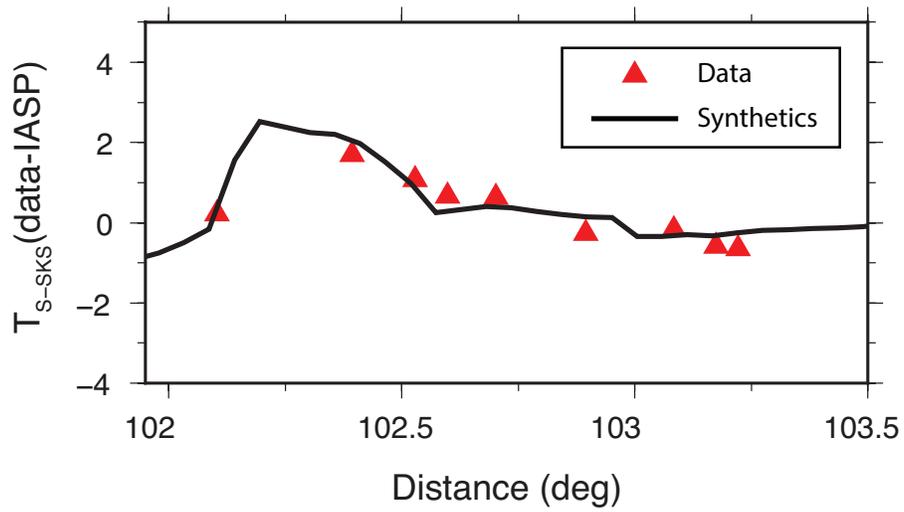


Figure S7

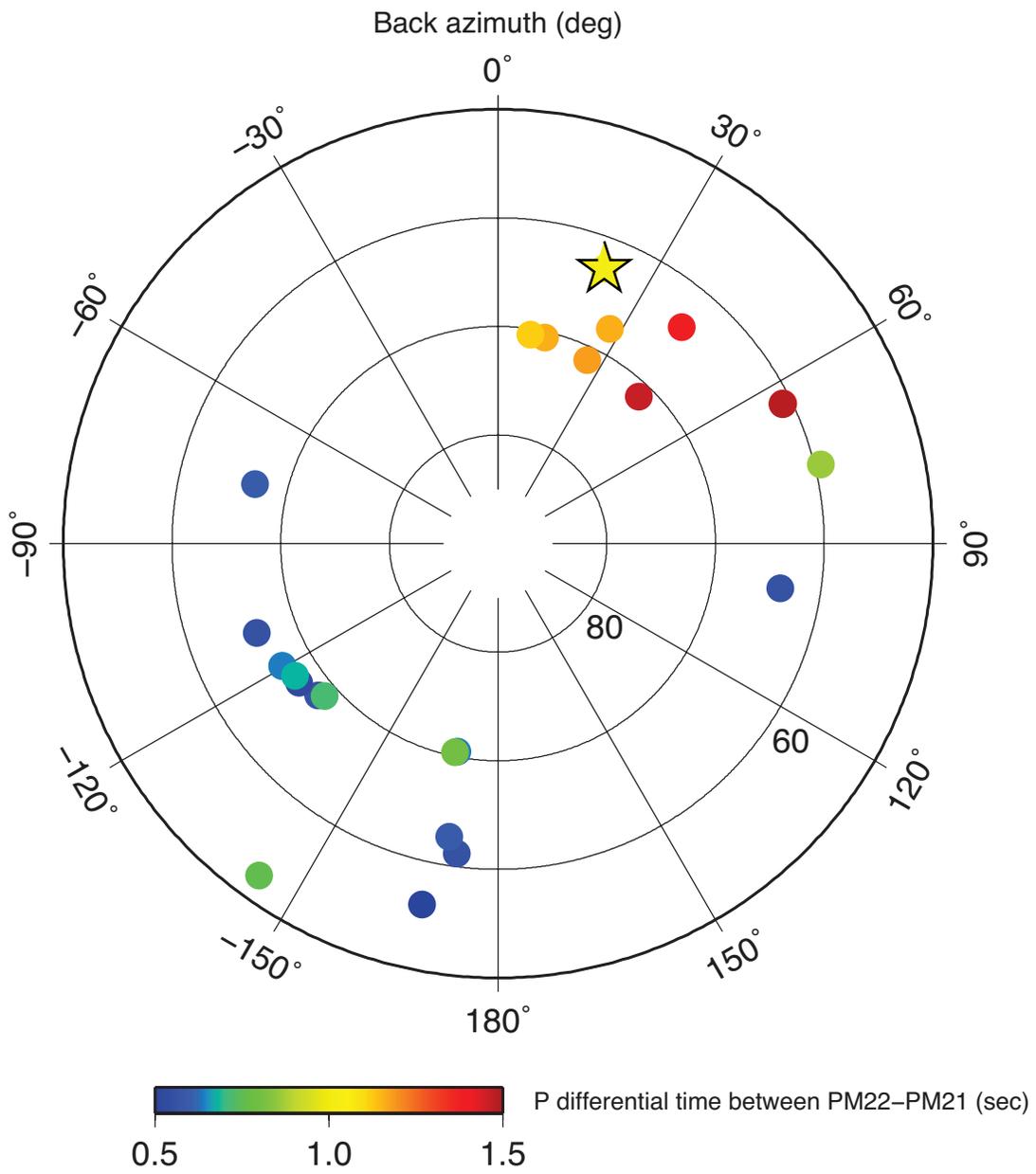


Figure S8

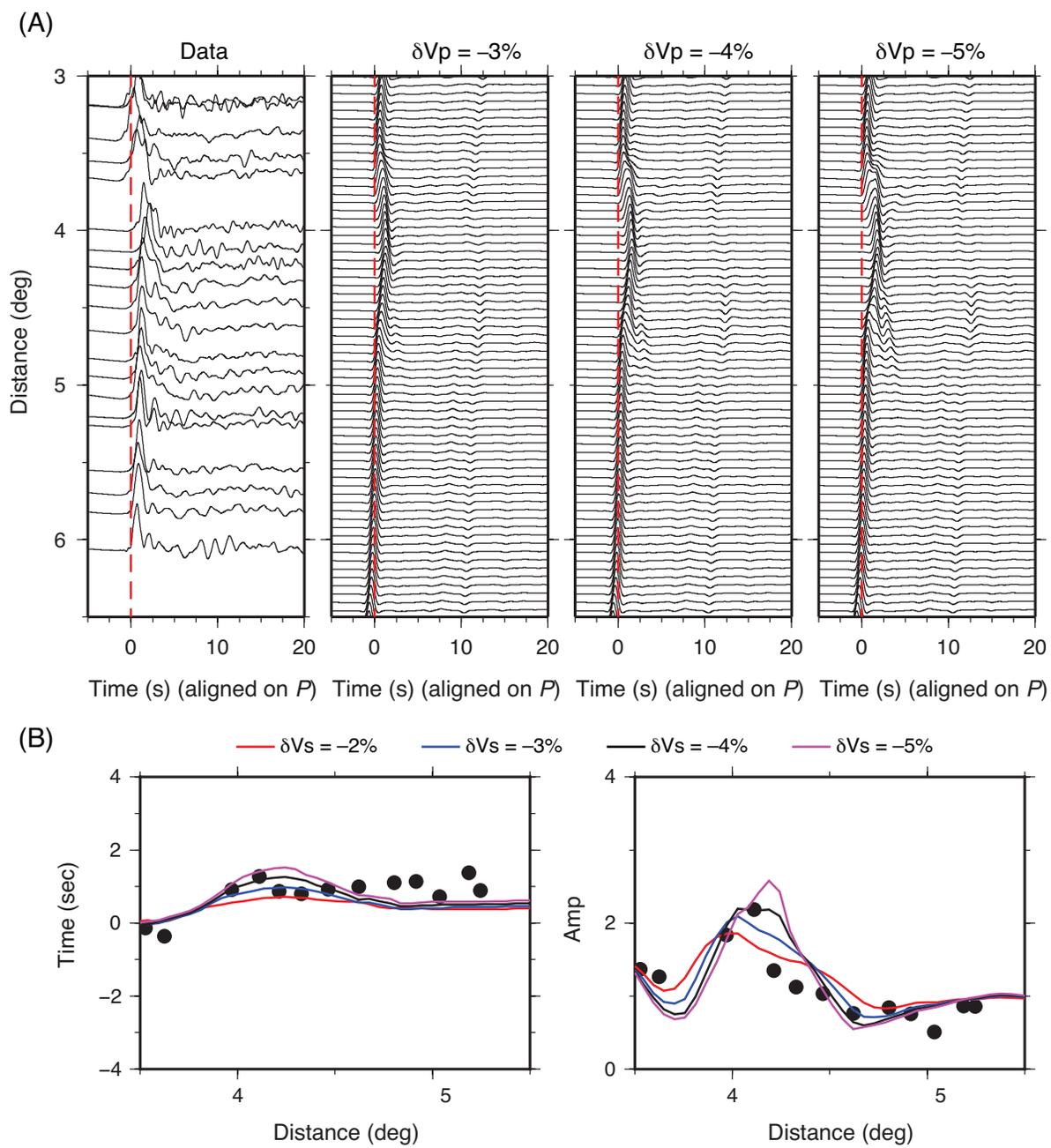


Figure S9

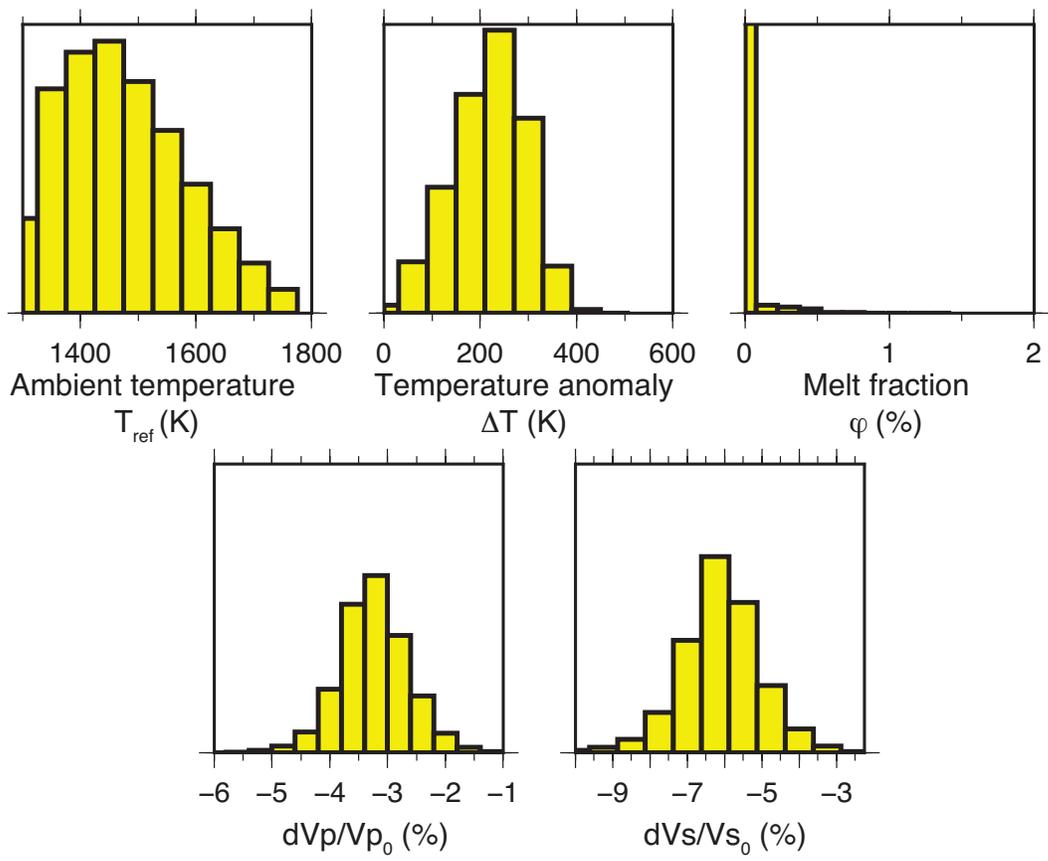


Figure S10