Global slab deformation and centroid moment tensor constraints on viscosity

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We analyze moment tensor solutions from deep subduction zone earthquakes to determine global slab deformation patterns. Inferred strain rates are compared to predicted deformation patterns from fluid models to help constrain the first-order radial and lateral viscosity structure of the Earth. While all slabs that reach the lower mantle are compressed at their tip, intermediate depth patterns are more complex. We compute 3-D spherical flow with various slab rheologies and compare the angular misfit between the compressive eigenvectors of the resultant stress field and global centroid moment tensor (gCMT) solutions. We find that upper mantle slab viscosities of \( \sim 10^{-100} \) and lower mantle viscosities of \( \sim 30^{–100} \) times the upper mantle produce the best match to gCMTs. A 0.1 viscosity reduction in the asthenosphere seems preferred. Slab geometry and lower mantle viscosity exert significant control on deformation. Inclusion of the phase changes at 410 km and 660 km increases extensional deformation at intermediate depth and compressional deformation at the lower mantle, improving the match to gCMTs for strong slabs. Our conclusions are fairly insensitive to surface boundary conditions. However, models which include net rotations of the surface with respect to the lower mantle produce compression at intermediate depths for west directed slabs and extension for east directed slabs. Without allowing for regional variations, these models yield the best match to gCMTs. While significant deviations between model and seismicity remain, our results show that seismicity provides an under-utilized constraint for slab dynamics.

1. Introduction

Subducting lithospheric slabs coupled to surface plates exert a pull force that contributes to plate motions [e.g., Forsyth and Uyeda, 1975]. In the upper mantle, flow induced by slab density exerts tractions on the base of plates that further contribute to surface motions. Together, these forces act to make subducting slabs the main driving force for plate tectonics [e.g., Lithgow-Bertelloni and Richards, 1998; Becker and O’Connell, 2001; Conrad and Lithgow-Bertelloni, 2002]. Geodynamic models with only radial viscosity variations and dense slabs are sufficient for reproducing relative plate motions [e.g., Ricard and Vigny, 1989]. However, lateral viscosity variations (LVVs) are required to speed up oceanic compared to continental plates [Conrad and Lithgow-Bertelloni, 2002; Becker, 2006] and
to induce the net rotation of the lithosphere as observed in plate motion models [e.g., Ricard et al., 1991]. The strength of slabs relative to the mantle, which may constitute an important component of LVVs, is, however, still a matter of debate (see reviews by Billen [2008] and Becker and Faccenna [2009], for example), as are variations in viscosity with depth in the mantle [e.g., Mitrovica and Forte, 2004].

[3] Geodynamic models often invoke a layer of reduced viscosity beneath the lithosphere as well as an increase in viscosity in the lower mantle. The existence of a low-viscosity layer between the lithosphere and the transition zone at ~300 km is supported by rock rheology [e.g., Hirth and Kohlstedt, 2004] and geoid studies [e.g., Hager and Clayton, 1989; Mitrovica and Forte, 2004]. Geoid studies also support an increase in viscosity with depth to support subducting slabs [Hager, 1984]. Models with a ~50× increase in viscosity applied to the lower mantle improve the match to the observed geoid [e.g., Hager and Clayton, 1989; Forte, 2007].

[4] Buoyancy forces from density anomalies in the upper mantle produce deformation within subducting slabs accommodated by seismic strain release [e.g., Vassiliou and Hager, 1988]. Strain rates calculated from seismic moments [e.g., Bevis, 1988] suggest that plates deform anelastically at intermediate depths. Slab geometry from seismicity and seismic tomography [e.g., Hager and O’Connell, 1979; van der Hilst et al., 1997] images highly contorted slab shapes inconsistent with rigid, elastic plates. Such studies justify a fluid treatment of subducting slabs in numerical models. Moreover, slabs are imaged by tomography as deflected by or penetrating through the 660 km discontinuity [e.g., van der Hilst et al., 1997; Fukao et al., 2001]. Lower mantle slabs may also contribute to plate driving forces, though the mode of force transmission is debated [e.g., Deparis et al., 1995; Lithgow-Bertelloni and Richards, 1995; Becker and O’Connell, 2001; Stadler et al., 2010].

[5] Given the uncertainties about slab strength and force transmission, additional constraints from seismicity are helpful. Coseismic strain release from deep earthquakes can be imaged by focal mechanism and moment tensor solutions, from P wave first motion studies [e.g., Isacks and Molnar, 1969] and waveform modeling [e.g., Dziewoński et al., 1981], respectively. The principal axes orientations of these solutions can be interpreted as reflecting the slab deformation state [e.g., Isacks and Molnar, 1969; Giardini and Woodhouse, 1984; Chen et al., 2004]. In the classic analysis of 204 focal mechanisms by Isacks and Molnar [1971], stress orientations suggest slabs deform by extension as they enter the low-strength upper mantle. As the slabs become supported by the higher-strength lower mantle, they become compressed. This transition is reflected in gaps in seismicity. If seismicity is continuous, compression was observed at all depths, with the inference that the slab is fully supported by the lower mantle [Isacks and Molnar, 1971]. Knowledge of the slab deformation state can guide numerical modeling in an effort to constrain the rheology of the slab and mantle that produces the best match to observations. Vassiliou and Hager [1988] compared stress orientations from 2-D fluid models to focal mechanisms, moment tensor inversions, and summed seismicity with depth. Their results suggest that slab and lower mantle viscosities are comparable and are approximately 1 order of magnitude more viscous than the upper mantle. Similar slab viscosities were inferred by Billen and Gurnis [2003] and Billen et al. [2003] when comparing stress orientations, the geoid, strain rate, and dynamic topography to three-dimensional (3-D) regional models. Two-dimensional work by Carminati and Petricca [2010] explored stress orientations resulting from viscoelastic models to investigate the role of large-scale plate motions on intermediate depth deformation.

[6] In this study, we use the gCMT catalog (Global CMT Project, 2006, available at http://www.globalcmt.org/, accessed May 2008) (formerly Harvard CMT) to build a global snapshot of the deformation state of subducting slabs. Building on the work by Vassiliou and Hager [1988], we investigate the rheology of the slab, asthenosphere, and lower mantle in a 3-D spherical mantle flow model, considering the effects of phase transitions, lower mantle density, boundary conditions, and net rotation of the lithosphere with respect to the lower mantle. We compare stresses produced by the flow model to the gCMTs by calculating the 3-D angular misfit between compressional (P) axes orientations in regions of deeply extending seismicity. Our goal is to constrain the first-order radial and lateral viscosity structure of the Earth that is complementary to the existing results from geoid, dynamic topography and plate motion studies.

[7] We find that upper mantle slab viscosities of ~10–100, lower mantle viscosities of ~30–100, and asthenosphere viscosities of ~0.1 times the upper mantle produce the best match to gCMTs.
Plate kinematic forces and the net rotation generate second-order variability in intermediate depth deformation.

[8] We first present our resurvey of gCMT solutions in slabs and follow with results from numerical modeling.

2. Slab Deformation

2.1. Methods

[9] To obtain an overview of slab deformation, we generate a new global compilation of moment tensor strain orientations as a function of depth. We use the gCMT catalog with all events up to December 2008. Earthquake distribution with depth, and the depth to the transition from predominant extension to compression in subducting slabs, suggest defining ~100–350 km as intermediate and ~350–700 km as deep [e.g., Vassiliou and Hager, 1988] (see also Appendix A). To determine the coseismic strain orientation in subducting slabs at different depths we analyze profiles from 24 geographic regions (Figure 1). Profile parameters (Appendix B) were chosen to most closely replicate the cross sections of Isacks and Molnar [1971], however, some regions were divided to investigate differences along strike. We orient profiles roughly perpendicular to the strike of seismicity contours as defined by Gudmundsson and Sambridge [1998], then choose profile lengths and widths so that the gCMTs provide the most complete slab profile (Appendix B). For each profile, we divide the gCMT solutions into 50 km depth bins between 100 and 700 km and sum the normalized moment tensors for each bin to obtain the average coseismic strain orientation. Following Isacks and Molnar [1971], we exclude solutions shallower than 100 km to avoid earthquakes related to convergence.

[10] We sum normalized moment tensors [e.g., Fischer and Jordan, 1991; Bailey et al., 2009] rather than moment tensors [Kostrov, 1974] due to the sensitivity of the latter to the largest earthquake considered. To obtain the slab dip corresponding to each of the depth bins we fit a polynomial to the mean hypocenter positions within the bins based on the Engdahl earthquake catalog from 1960 to 2007 [Engdahl et al., 1998; E. R. Engdahl, personal communication, May 2008]. The order of the polynomial in each case is given by \( \min(4, N_{\text{bins}} - 2) \), where \( N_{\text{bins}} \) is the number of bins containing hypocenter data, and the fit is performed using mean depth within bins as the dependent variable and mean horizontal position within bins as the independent variable.

[11] We compute the slab dip at the mean depths within bins of the gCMTs by taking the derivative of the polynomial at that depth. We use this dip to rotate the summed tensor into a slab coordinate system, and divide the down-dip component by the
As previously noted [e.g., Isacks and Carminati and Molnar, 1971], the global slab deformation pattern can be broadly separated into predominantly intermediate extension for east directed subduction zones and predominantly intermediate compression for west directed subduction zones (Figure 1). Previous workers have recognized intermediate depth complexities in slab geometry, or interaction of the slab with mantle heterogeneity.

2.2. Results

[12] As previously noted [e.g., Carminati and Petrica, 2010], the global slab deformation pattern can be broadly separated into predominantly intermediate extension for east directed subduction zones and complexities in slab geometry, or interaction of the slab with mantle heterogeneity.
3. Quantitative Modeling With Flow Models

3.1. Methods

3.1.1. Seismicity Analysis

Our goal is to compare stress from fluid modeling to the strain orientations of gCMT solutions to produce a model with slab and mantle rheology that results in the best match to observations. We compare stress to gCMT solutions along 30 subduction zone segments from 13 of the 24 regions shown in Figure 1. We compare only regions with deeply extending seismicity and divide most regions into adjacent transects to investigate differences along strike. Although gCMTs image strain, we assume that earthquakes at a given location are an unbiased sample of the long-term stress orientation. Comparing fluid stress to seismically recorded strain release presumes no specific mechanism for deep earthquakes, but assumes that the stresses must be oriented appropriately for an earthquake to produce the resultant moment tensor. To avoid complexities potentially related to the mantle wedge region, we compare model results to gCMT solutions from 200 km to 700 km.

While Figure 2 shows normalized moment tensor summation results, stress results from our flow models are extrapolated and compared to the nearest individual gCMT solution. The model fit to a gCMT solution is quantified by the angular difference between the gCMT P axis and the most compressive eigenvector from the fluid stresses, which we refer to as the misfit. To quantify the misfit for a single subduction zone (regional misfit) we compute the average of average misfits for 100 km depth bins. The global misfit is given by the average of all of the segments. This averaging at multiple scales allows us to avoid biases due to strong clustering of gCMT locations at certain depths. Although this measure neglects the orientation of intermediate and extensive axes, we find a strong correlation between eigenvector misfits and a measure based on the tensor inner product (Appendix C), so only the more intuitive eigenvector misfits are reported.

3.1.2. Fluid Model

To model mantle circulation we use CitcomS [Zhong et al., 2000], a spherical finite element code based on Citcom [Moresi and Gurnis, 1996], from the Computational Infrastructure for Geodynamics (geodynamics.org). CitcomS solves the equations governing mantle circulation with a variable viscosity structure. Conservation of mass (1) and momentum (2) are solved for instantaneous Stokes flow in the Boussinesq approximation:

$$\nabla \cdot \mathbf{u} = 0, \quad (1)$$

$$-\nabla p + \nabla \cdot \left[ \eta (\nabla \mathbf{u} + \nabla^T \mathbf{u}) \right] - (Ra_T T - Ra_{pc} \Gamma) \mathbf{e}_r = 0, \quad (2)$$

where $p$ is the dynamic pressure, $\eta$ is the viscosity, $\mathbf{u}$ is the velocity, $T$ is the temperature, $Ra_T$ is the thermal Rayleigh number scaled to the radius of the Earth, $Ra_{pc}$ is the phase change Rayleigh number, $\Gamma$ is the phase change function, and $\mathbf{e}_r$ is the unit vector in the radial direction. The strength of the thermal buoyancy is controlled by $Ra_T$ defined as:

$$Ra_T = \frac{\rho_0 g \alpha \Delta T R^3}{\kappa \eta_0},$$

where $\rho_0$ is the reference density, $g$ is gravitational acceleration, $\Delta T$ is the temperature difference between the surface and the asthenosphere, $\alpha$ is the coefficient of thermal expansion, $R$ is the radius of the Earth, $\kappa$ is the thermal diffusivity, and $\eta_0$ is the reference viscosity.

All parameters are listed in Table 1. Phase change values were chosen as end members to emphasize their effect [Vidale and Benz, 1982; Bina...
The temperature at which 
\( g = 1 \) is given by:

\[
Ra = \frac{g}{\frac{\Delta \rho \cdot g \cdot R^2}{\kappa \eta_0}}.
\]

where \( \Delta \rho \) is the density contrast and \( g \) is the gravitational acceleration.

The deflection of the phase change boundary due to temperature, \( \Gamma \), is given by:

\[
\Delta \rho \cdot g \cdot R^3
\]

The strength of the phase change is controlled by:

\[
Ra = \frac{g}{\frac{\Delta \rho \cdot g \cdot R^2}{\kappa \eta_0}}
\]

where \( \Delta \rho \) is the density difference of the phase change. The deflection of the phase change boundary due to temperature, \( \Gamma \), is given by:

\[
\Gamma = \frac{1}{2} \left( 1 + \tanh \left( \frac{p_r}{\rho g W_{pc}} \right) \right)
\]

where \( p_r \) is the reduced pressure, \( r \) is the radial direction, \( \gamma \) is the Clapeyron slope of the phase change, \( T \) is the temperature, and \( d_{pc} \) is the depth, \( W_{pc} \) is the width, and \( T_{pc} \) is the temperature at which the phase change occurs. The phase change temperature, depth, and Clapeyron slopes are non-dimensionalized as:

\[
T' = \frac{T}{\Delta T},
\]

\[
d'_{pc} = \frac{d_{pc}}{R},
\]

and

\[
\gamma' = \gamma \left( \frac{\Delta T}{\rho_0 g R} \right).
\]

Horizontal numerical resolution was typically ~25 km with a mesh of 129 elements in the vertical, with ~17 km spacing from the surface to 660 km, ~18 km spacing from 660 to 1200 km, ~20 km spacing from 1200 to 1800 km, and ~36 km spacing from 1800 km to the core mantle boundary (CMB). Resolution tests (Appendix D) show that our computations converge under successive refinement and that global misfit differences of 0.2° are within the error of the upper mantle element spacing for the moderate resolution computations employed for convenience. The regions which show the greatest relative variation, though minor, in absolute misfit values are those with sparse gCMT coverage within a depth bin. Higher resolution would lead to insignificant increases in accuracy with a significant increase in computational resources.

To incorporate density anomalies in a mantle circulation model, we convert the regionalized upper mantle (RUM) model of Gudmundsson and Sambridge [1998] into temperature. RUM slabs are contours of seismicity based on the Engdahl et al. [1998] earthquake catalog. RUM contours are converted to 100 km wide polygons and interpolated at 50 km depth intervals. Our goal is to quantify slab viscosity to first order, so a 3.9% density anomaly, corresponding to a 1300 K temperature anomaly, is prescribed for most models and remains constant throughout the slab. For some models, we prescribed a 1.8% density anomaly, corresponding to a 600 K temperature anomaly, to explore end-member estimations of the expected thermal anomaly for subducting slabs. The lithosphere extends from the surface to 100 km for most models and was modeled as a fluid with viscosity \( \eta_0 \), 10\( \eta_0 \), 100\( \eta_0 \), and 1000\( \eta_0 \), where \( \eta_0 \) is the upper mantle reference. For selected models, we reduced the lithosphere and slab thickness to 50 km, maintaining the same 3.9% density anomaly, further exploring younger slabs. Lithospheric age will control the thermal structure of slabs, but we vary the buoyancy and thickness separately because effective thickness might depend on rheological factors other than temperature. Below the lithosphere, we control the slab viscosity (\( \eta_{slab} \)) with a temperature-dependent rheology \( \eta = \eta_0 \exp(E(\Delta T)) \). Here, \( E \) determines the strength of the temperature dependence and \( \Delta T \) is the temperature difference between...

### Table 1. Model Parameters

<table>
<thead>
<tr>
<th>Variable Name</th>
<th>Symbol</th>
<th>Value</th>
</tr>
</thead>
<tbody>
<tr>
<td>Reference viscosity</td>
<td>( \eta_0 )</td>
<td>10^{21} Pa s</td>
</tr>
<tr>
<td>Radius of the Earth</td>
<td>R</td>
<td>6371 km</td>
</tr>
<tr>
<td>Thermal diffusivity</td>
<td>( \kappa )</td>
<td>10^{-8} m^2 s^{-1}</td>
</tr>
<tr>
<td>Thermal Rayleigh number</td>
<td>( Ra )</td>
<td>3.4 × 10^8</td>
</tr>
<tr>
<td>Temperature change between surface and asthenosphere</td>
<td>( \Delta T )</td>
<td>1300 K</td>
</tr>
<tr>
<td>Phase change Ra #(410 km)</td>
<td>( Ra_{410} )</td>
<td>4.07 × 10^8</td>
</tr>
<tr>
<td>Phase change Ra #(660 km)</td>
<td>( Ra_{660} )</td>
<td>7.24 × 10^8</td>
</tr>
<tr>
<td>Gravitational acceleration g</td>
<td>( g )</td>
<td>10 m s^{-2}</td>
</tr>
<tr>
<td>Coefficient of thermal expansion</td>
<td>( \alpha )</td>
<td>3 × 10^{-3} K^{-1}</td>
</tr>
<tr>
<td>Reference density</td>
<td>( \rho_0 )</td>
<td>3500 kg m^{-3}</td>
</tr>
<tr>
<td>Slab/upper mantle density contrast</td>
<td>( \Delta \rho )</td>
<td>3.9%</td>
</tr>
<tr>
<td>Clapeyron slope, 410 km</td>
<td>( \gamma_{410} )</td>
<td>3.0 MPa K^{-1}</td>
</tr>
<tr>
<td>Clapeyron slope, 660 km</td>
<td>( \gamma_{660} )</td>
<td>−3.5 MPa K^{-1}</td>
</tr>
<tr>
<td>Nondimensional ( \gamma ), 410 km</td>
<td>( \gamma'_{410} )</td>
<td>0.0175</td>
</tr>
<tr>
<td>Nondimensional ( \gamma ), 660 km</td>
<td>( \gamma'_{660} )</td>
<td>−0.0204</td>
</tr>
<tr>
<td>Density change at 410 km</td>
<td>( \Delta \rho_{410} )</td>
<td>4.5%</td>
</tr>
<tr>
<td>Density change at 660 km</td>
<td>( \Delta \rho_{660} )</td>
<td>8.0%</td>
</tr>
<tr>
<td>Width of the phase transitions</td>
<td>( W_{pc} )</td>
<td>20 km</td>
</tr>
<tr>
<td>Nondimensional ( T' ) of PC at 410 km</td>
<td>( T_{410} )</td>
<td>0.9965</td>
</tr>
<tr>
<td>Nondimensional ( T' ) of PC at 660 km</td>
<td>( T_{660} )</td>
<td>0.99879</td>
</tr>
</tbody>
</table>

*The Rayleigh number is defined using the Earth’s radius as in the work by Zhong et al. [2000].
the slab and the ambient mantle. Since our nondimensional slab temperature is zero and the upper mantle temperature is unity, the maximum slab viscosity is $\eta_0 \exp(E)$. Values for $E$ were chosen to most closely match the lithosphere viscosity which is controlled by a preexponent factor ($1-1000$) applied to the lithospheric layer.

[21] The asthenosphere is applied globally as part of the radial viscosity structure, extending from the base of the lithosphere to 300 km. The asthenosphere was modeled with viscosities ($\eta_{\text{asth}}$) of $0.1\eta_0$, 0.01$\eta_0$, and $0.001\eta_0$. In models with a reduced viscosity asthenosphere, the temperature dependence of the RUM slabs was increased in the asthenospheric layer inversely proportional to the viscosity reduction in the surrounding asthenosphere to maintain the same slab viscosity throughout the upper mantle. We also explored limiting the asthenosphere to beneath oceanic regions only, by excluding continental areas defined by 3SMAC [Nataf and Ricard, 1996] when applying the reduction in viscosity to the asthenospheric layer. Results were similar to models with a globally applied asthenosphere and we have excluded these from this discussion.

[22] The lower mantle extends from 660 km to 2891 km and was modeled with viscosities ($\eta_{\text{LM}}$) of $0.1\eta_0$, $10\eta_0$, $30\eta_0$, and $100\eta_0$. Models with no viscosity increase in the lower mantle produced consistently poor matches to gCMTs and we have excluded these results from the discussion.

[23] Surface boundary conditions are either free slip (FS models) or prescribed NUVEL-1A [DeMets et al., 1990] plate motions, either in the no net rotation reference frame of DeMets et al. [1990] (NNR models), or the HS3 absolute plate motion model of Gripp and Gordon [2002] (HS3 models). For NNR and FS models, we use a free slip boundary condition at the core mantle boundary (CMB). For HS3 models, we use a no-slip CMB boundary condition to induce net rotation related shearing [e.g., Conrad and Behn, 2010]. Fixing the CMB and prescribing absolute plate motion models with net rotation, channels the net rotation flow into the upper mantle, roughly consistent with the dynamically induced net rotations from geodynamic models [Zhong, 2001; Becker, 2006]. Our stiff slab models also induce net rotations self-consistently, as discussed below.

[24] Phase transitions in the upper mantle enhance or reduce the effect of buoyancy forces within cold slabs [e.g., Turcotte and Schubert, 2002]. The olivine-spinel phase change at 410 km aids subduction due to the upward deflection of the phase change within the cold slab, increasing negative buoyancy and promoting increased extension at intermediate depths. The postspinel to perovskite + magnesiowüstite phase change at 660 km produces a downward deflection of the phase boundary in the slab, decreasing negative buoyancy. Penetration of the slab into the lower mantle can be hindered by the magnitude of the negative Clapeyron slope of the 660 km phase change [e.g., Christensen and Yuen, 1984; Christensen, 1996; Kirby et al., 1996] or promote ponding at the lower mantle [Steinbach et al., 1993; Christensen, 1996]. We explore the role of phase changes by testing several NNR and HS3 models with and without phase changes, using end-member parameters given in Table 1.

[25] While upper mantle slabs are defined by seismicity and tomography as narrow features descending from oceanic lithosphere in active subduction zones, tomographic images of slabs in the lower mantle are more diffuse and of ambiguous extent. Thus, the influence of lower mantle slabs on flow and deformation is difficult to determine. Lower mantle density anomalies are here inferred from seismic tomography [e.g., Hager and Clayton, 1989]. For selected models, we combine upper mantle RUM slabs with lower mantle density anomalies inferred from the mean S wave model (SMEAN) of Becker and Boschi [2002]. In these models, the RUM slabs are used for buoyancy and rheology in the first 700 km and the SMEAN tomography is used for buoyancy only in the lower mantle. Motivated by previous, similar studies (e.g., review by Forte [2007]), we scale velocity anomalies to density as $\frac{d\ln V_s}{d\ln \rho} = 0.25$.

### 3.2. Results

[26] We now provide an overview of all models discussed here, explore similarities between the best performing models in terms of global and regional misfit variations, show how they are affected by our global parameter choices, and discuss what this implies for the radial and lateral viscosity structure of the Earth. We then discuss the effect of phase changes, lower mantle density contributions, boundary conditions, net rotation, and regional variations.

#### 3.2.1. Global Model Performance

[27] We evaluated 517 flow models in total and present in Table 2 the global angular eigenvector misfit averages for 216 selected models. The global misfits presented in Table 2 represent the average of the misfits from the 30 subduction zone segments.
evaluated. In Figure 3 we show the regional misfit for the globally best performing HS3 model for slabs with a 3.9% density anomaly. Models consistently perform near random in areas of extreme arc curvature, where extension persists to ~500 km. These regions are best matched with strong slabs, but the difficulty of generating significant compression at the lower mantle to match gCMTs results in high misfit.

[28] Several general trends are reflected in Table 2. Best performing models, when evaluated based on global misfit, are dominated by those with upper mantle slab viscosities of 10–100\(\eta_0\) and at least a 1 order of magnitude reduced viscosity asthenosphere. Deviations from solely buoyancy-driven deformation occur from bending and unbending of the slab, and the compressional effect of the lower mantle viscosity. As slab strength increases, bending deformation dominates and slab geometry controls deformation. While slab geometry remains constant between models, lower mantle viscosity exerts a secondary control on deformation. The propagation of compression upward within the slab is controlled by the strength of the lower mantle and of the slab. Weak slabs (1–10\(\eta_0\)) are more easily deformed by the lower mantle viscosity increase and perform slightly better with a 30\(\eta_0\) lower mantle viscosity. As slab strength increases, comparable, and often stronger, lower mantle viscosities are preferred. Phase transitions act to enhance the same deformation patterns obtained from models without phase transitions.

Figure 3. Regional misfit for a single, globally best performing HS3 model with a 3.9% density anomaly and 100 km thick slab and lithosphere. Slab viscosity is 10\(\eta_0\), with no viscosity reduction in the asthenosphere, and a 100\(\eta_0\) lower mantle viscosity. The average misfit for the model is 36.0° with a minimum of 14.2° and a maximum of 57.6°.
misfit, but improve misfits for regions with dominant compression.

[29] Surface free slip models perform similar to prescribed velocity models, though our results highlight the importance of net rotations for global flow models. HS3 models, with a no-slip CMB, perform better than NNR and surface free slip models as a whole (Table 2, HS3 models), substantiating the suggestion by Carminati and Petricca [2010] that the net rotation of the lithosphere can induce compression in west directed slabs and extension in east directed slabs.

[30] We follow with an evaluation of the global results in terms of each of the modeling parameters, and then discuss regional variations, describing regions that illustrate our interpretations. The latter are, however, based on numerical evaluation of all subduction zones analyzed here.

3.2.2. Slab Strength

[31] Global model performance as a function of slab viscosity is presented in Figure 4 for each boundary condition. Thin, cool slabs (50 km thick and average ΔT of 1300°) and thick, warm slabs (100 km thick, average ΔT of 600°), perform similar to thick, cool slabs (average ΔT of 1300°) in terms of global misfit. However, well-performing thin (TH models of Table 2) and warm (RA models of Table 2) slab models have slab strengths up to 100η0 whereas thick slabs often show increased misfit values at these strengths (Table 2). From our flow models, we observe that in the absence of kinematic forces, slabs are predominantly extensive at intermediate depths and compressive at the lower mantle. The stronger the slab, the deeper the extensive signal and the stronger the lower mantle must be to generate compression. Best performing surface free slip models have slab viscosities less than 100η0. For kinematic models, additional force imposed by the surface velocity propagates into the slab (representing incompletely modeled global plate driving forces). These models perform best for slab viscosities of 100η0 if lower mantle viscosities are comparable.

[32] Regional variations (Figure 3) partially reflect the variable slab geometries seen in Figure 2. As expected from theory and modeling such as Houseman and Gubbins [1997], we find that compression dominates in areas of unbending and extension in areas of bending. With increasing viscosity, the compressive unbending signal strengthens.
and, in some cases, propagates throughout the slab (Kurile slab (Figure 5b)). Extension dominates for strongly concave slabs like the Marianas, but with increasing viscosity compression propagates from unbending regions of the slab toward bending regions, resulting in regions of compression on the less concave surface, and extension beneath (Marianas (Figure 5a)).

3.2.3. Asthenospheric Strength

Global misfit values improve for most models with the addition of an asthenospheric layer of reduced viscosity below the lithosphere (Table 2). Therefore, we explored parameter ranges in more detail for these models. Strong slabs are more sensitive to the asthenosphere viscosity and show improvement with increasing asthenosphere viscosity reduction (Figure 6). With increasing lower mantle viscosity, the differences are minimized and the models become less sensitive to asthenospheric viscosity, except for HS3 models, where the match degrades with the addition of an asthenosphere for lower mantle viscosities greater than $10\eta_0$ (Figure 6). This results from channeling the net rotation component into the asthenosphere, discussed below.

Regionally, a low-viscosity asthenosphere reduces intermediate depth extension and allows
compression from the lower mantle to propagate upward through the slab. For regions with intermediate depth compression, the match to gCMTs is improved. The effect is significant for 1000 $\eta_0$ slab viscosities (Table 2) where extension is otherwise dominant.

### 3.2.4. Lower Mantle Strength

[36] Global model performance as a function of lower mantle viscosity is presented in Figure 7. Results show that lower mantle viscosities of $\approx 30$–100 $\eta_0$ produce the best match to gCMTs. Increasing the lower mantle beyond 30 $\eta_0$ has a negligible effect for most models, except for very strong slabs. Testing lower mantle viscosities of 200 $\eta_0$ produced the same global misfit as models with a 100 $\eta_0$, with insignificant regional differences. This is in agreement with Vassiliou and Hager [1988] who found that increasing the lower mantle viscosity beyond 50 $\eta_0$ did not significantly change resulting stress magnitudes. As seen in Table 2, the best performing models in terms of global misfit are those in which the slab and lower mantle viscosity contrasts are less than 1 order of magnitude.

[37] Regionally, for all slabs in our fluid models, an increase in viscosity in the lower mantle will promote compression at the slab tip that propagates upward through the slab with increasing lower mantle viscosity, improving the match to gCMTs for regions with dominant compression. Our resurvey of gCMTs indicates that all deeply extending slabs undergo predominantly compressional deformation upon interaction with the lower mantle (Figure 2). Weak slabs and thin slabs are more easily compressed and are therefore less sensitive to increasing lower mantle viscosity (Figure 7). Strong slabs (100–1000 $\eta_0$) require a greater increase in lower mantle viscosity to become compressed in the absence of unbending (Figure 8).

### 3.2.5. Phase Transitions

[38] Global misfits for models with phase transitions included (PC models of Table 2) are improved with the addition of a phase change only for slab viscosities of 1000 $\eta_0$. The addition of the 410 km and 660 km phase changes generate intermediate extension and deep compression, which are patterns readily attained in flow models without phase transitions. Strong slabs require high lower mantle viscosities to become compressed so the compression generated by the 660 km phase change is noticeable in the global misfit results.

[39] Regionally, for weak slabs, the 410 km phase change improves intermediate depth misfit while the 660 km phase change is less significant because the lower mantle viscosity increase generates enough compression to produce a good match to gCMTs. For strong slabs whose geometry induces intermediate compression contrary to observations, the 410 km phase change improves the misfit at intermediate depths. Accordingly, regions with observed intermediate depth compression show poorer matches with the addition of the phase changes due to the extension generated by the 410 km transition.
For stronger lower mantle viscosities, the effect of the 410 km phase change is overwhelmed by the compression induced by the lower mantle viscosity.

### 3.2.6. Lower Mantle Density Anomalies

[40] Globally, including lower mantle density anomalies produced insignificant changes in the match to gCMTs for weak slabs (Table 2). For strong slabs and weak lower mantle viscosity, global model results worsened with the inclusion of lower mantle density anomalies. Regionally, in areas with lower mantle downwellings inferred from SMEAN, the pull from the lower mantle increased extension throughout the slab. For strong slabs and weak lower mantle viscosity, the increased extension could not be overcome and the slab remains in extension at depth, producing poorer matches to gCMTs. With increasing lower mantle viscosity, values are improved as compression increases at the slab tip. In regions of lower mantle upwelling inferred from SMEAN, increased compression from the lower mantle improves the match to gCMTs for dominantly compressive regions such as Tonga. This supports the work of Gurnis et al. [2000] who infer from tomography that the Tonga region over-lies the edge of the Pacific superplume and attribute the significant compressional deformation in the Tonga slab to the rising plume.

### 3.2.7. Surface Boundary Conditions and Net Rotations

[41] Global misfit values from Table 2 show that surface free slip models perform similar to NNR and HS3 for most models. For slab viscosities of $1000\eta_0$, free slip models perform the best. NNR and surface free slip models improve with the addition of a reduced viscosity asthenosphere. HS3 model results are similar to, or improve with a $0.1\eta_0$ asthenospheric viscosity reduction, but results are worse for a higher-viscosity reduction.

[42] Regionally, differences result from the ability of the model to affect intermediate depth deformation.
patterns and to better reflect the asymmetry observed in Figure 1. Surface free slip models generate intermediate extension, NNR models generate some intermediate depth compression, and HS3 models generate strong intermediate depth compression for west directed regions and extension for east directed regions. The two factors controlling this pattern are the propagation of additional force from the kinematic surface boundary, and the amount of induced net rotation of the lithosphere with respect to the lower mantle.

Surface free slip models generally produce intermediate depth extension that increases with increasing lithospheric strength, and compression at the lower mantle that propagates upward with increasing lower mantle strength. Strong, free slip slabs are difficult to compress and therefore perform poorly compared to NNR or HS3 models that generate intermediate depth compression. However, strong, free slip slabs induce a net rotation of the lithosphere with respect to the lower mantle.

Misfit values for surface free slip models with weak zones are intermediate between NNR and surface free slip models without weak zones. Weak zones were used only for comparing net rotations generated in surface free slip models and were implemented by reducing the lithospheric viscosity 2 orders of magnitude 100 km to both sides of the NUVEL-1A [DeMets et al., 1990] plate boundaries.

Figure 9. Mean surface velocity versus slab viscosity for surface free slip and surface free slip with weak zone models with a $\eta_{LM}/\eta_0 = 30$ and a $\eta_{asth}/\eta_0 = 0.1$. Right axis (dashed lines) shows net rotation in percent of HS3 computed by $\langle \omega_{HS3} \cdot \omega_{model} / |\omega_{HS3}|^2 \rangle$, where $\omega$ is the Euler pole of the net rotation (mean mantle no-net-rotation (MM-NNR) reference frame [cf. Zhong, 2001]). Slabs with unity viscosity compared to the upper mantle produce a slight net rotation due to LVVs produced by the reduced viscosity asthenosphere. Negative values for the net rotation as a percent of HS3 are produced by a net rotation opposite in sense to that of HS3.
from the surface to the base of the lithosphere. Weak zones produce more plate-like motion of the surface, but reduce the strength of the net rotations generated by LVVs (Figure 9). The amount of net rotation induced by stiff, upper mantle slabs is comparable to that induced by stiff continental keels [Becker, 2006] and significantly larger that the net rotations based on global, lower-resolution slab models [Zhong, 2001; Becker and Faccenna, 2009].

The high net rotation HS3 models with a fixed CMB generate greater amounts of compression for west directed slabs, and the induced upper mantle shear generates extension in east directed slabs (Figure 10), reducing the compressive signal generated by the imposed surface velocity. HS3 models then perform better than NNR for regions that are predominantly compressive (Figure 11), but they cannot generate enough extension in east directed slabs to perform better than surface free slip models in dominantly extensive regions. HS3 models without an asthenosphere perform as well as models with a 0.1 reduced viscosity asthenosphere. However, if the asthenosphere is weakened beyond 0.1, the net rotation is channeled into the asthenosphere instead of smoothly decreasing through the upper mantle, reducing the effect of the induced shear below 300 km. As HS3 is thought to invoke too large a net rotation component based on global anisotropy modeling [Becker, 2008; Kreemer, 2009; Conrad and Behn, 2010], we explored various net rotation magnitudes for a suite of models. While the changes are minor, matches to gCMTs scale with increasing net rotation (Figure 12). Global misfit results for these models are provided in Appendix E.

Figure 10. Comparison of stress orientations between (a) NNR and (b) HS3 models on (left) the northwest directed Kurile-3 slab and (right) the east directed Sumatra slab. Slab viscosity is $100\eta_0$, and lower mantle viscosity is $30\eta_0$. The induced net rotations in the HS3 models generate increased compression at intermediate depths for west directed slabs and increased extension for east directed slabs.
misfit substantiates the work by Carminati and Petricca [2010], although the mechanics of force transmission are different in our global models.

However, as shown next, regional variations in slab strength may reduce the gCMT misfit of both NNR and HS3 to similar levels, without having to invoke net rotations.

### 3.2.8. Regional Model Performance

Figure 13 shows the selection of global models that lead to the best regional misfit for HS3 boundary conditions, which shows the spatial variability of our best fit models. By mixing flow models this way, the misfit can be reduced significantly (from 36° to 28.8°), though the proper test would be, of course, to allow for regional variations in a single global model, which we do not attempt here.

Figure 14 shows that best fitting slab viscosities are depth dependent. Here, we compare NNR and HS3 best fitting slab strengths by region for three depth ranges. From 200 to 350 km, most regions are best matched with slab viscosities of 100–1000η₀; from 350 to 500 km, most regions are best matched by slab viscosities of 10–100η₀; and from 500 to 700 km, most regions are best matched by slab viscosities of 100η₀. Average misfit values for each boundary condition and depth range (Figure 14) indicate that HS3 models perform slightly better in terms of regional misfit than NNR models at the 350–500 km depth range. This depth range represents the transition from dominantly extension to dominantly compression for most regions. From 500 to 700 km, HS3 models dominate the best fit.
Our best performing global models have a misfit of ~36°, as expressed by the global misfit measure (Table 2). The best models always perform much better than expected from randomness (Figure 11), and this is also the case for almost all regional subduction zone profiles for slab viscosities below 1000η₀. Moreover, there are robust trends of misfit values with parameters such as lower/upper mantle viscosity contrast. This gives us confidence that our interpretation as to the dynamic role of these parameters in a fluid modeling context are robust and useful, and that the global circulation model can be used as a backdrop against which to improve a dynamic description of subduction zone stress. The viscosity increase in the lower mantle that is preferred by our models is consistent with estimates obtained from geoid modeling [e.g., Hager and Clayton, 1989; Forte, 2007], and the relatively low effective slab strength (10–100 times the upper mantle) supports “weak slab” models (see, e.g., review by Becker and Faccenda [2009]).

While it is difficult to make such statements with statistical rigor, the remaining misfit of our best models appears too large to be explained by uncertainties in the gCMT parameters or modeling inaccuracies (Figure 3). Although we are somewhat limited by the resolution of our input models and the representation of plate boundaries, based on our assessment of numerical stability (Appendix D), we consider it more likely that the physical assumptions are incomplete. There are a range of effects that were not considered, including elasto-plastic rheologies, the effect of the crust, phase transitions in minerals other than olivine, compositional heterogeneity, and the preexistence of planes of weakness in the slab. Incorporation of some of these effects (lithospheric age, crustal thickness and sediment load) are better suited to regional studies. Other effects relate to more complicated rheology (e.g., effects of elasticity and mechanical anisotropy) which could be incorporated into future models. Effects of elasticity could clearly be important, and would introduce stress memory into the system.

Besides elasticity we feel that two effects are perhaps most important when trying to refine our models further, that of regional variation and that of existing planes of weakness (e.g., faults). Figure 13 shows that regionally, misfit can be improved by using different parameters, and one may expect that locally varying effects such as lithospheric age, crustal thickness, or sediment load will play an important role, for example in affecting regional slab viscosity. While such effects could be included in our models [cf. Wu et al., 2008], we decided not to optimize our global model in such a way, so as to not complicate the model too far.

Zones of weakness may, for example, be due to volatile assisted faulting [e.g., Ranero et al., 2003; Faccenda et al., 2008] and existing structures may still play a role in controlling faulting at depth [e.g., Jiao et al., 2000; Warren et al., 2008]. If mechanical anisotropy in subducted lithosphere matters below depths of 200 km, our simplified association of
The true evaluation of the relative importance of the effects which are ignored here has to wait until we have models that are able to incorporate a more complete rheological description of slabs. However, the results here indicate that our initial mechanical treatment is suitable to explain the first-order features of subduction zone deformation consistently over a global scale.

5. Conclusions

The global deformation patterns of subducting slabs show that strain release complexity is ubiquitous at intermediate depths. Broadly, regions can be separated into predominantly intermediate extension for east directed subduction zones and intermediate compression for west directed subduction zones. Best performing models suggest that slabs may not be stronger than ~10–100 times the upper mantle viscosity to match the observed deformation pattern if the fluid model provides a valid baseline.
Regionally, one control on deformation is slab geometry, as viscous bending produces compression in areas of unbending and extension in areas of bending. Increasing slab strength increases the importance of viscous bending in controlling deformation. Most models improve with the addition of a reduced viscosity asthenosphere and our results support at least a 1 order of magnitude reduction in viscosity. Slab strength may be spatially variable as well as depth dependent.

To obtain compression at the slab tip while maintaining some intermediate depth extension, lower mantle viscosities must be ~30–100 times the upper mantle viscosity, as suggested by Vassiliou and Hager [1988]. In addition to radial and lateral viscosity variations, the inclusion of phase changes at 410 km and 660 km produce increased intermediate depth extension and deep compression. Phase changes affect the global misfit only for strong slabs and weak lower mantle models, where compression at the slab tip is otherwise negligible. Lower mantle density anomalies also increase extension in areas of downwellings, and increase compression in upwelling regions, such as underneath Tonga [cf. Gurnis et al., 2000].

Our best performing global models are those that have net rotations of the surface with respect to the lower mantle larger than those typically excited by the slabs themselves [cf. Carminati and Petricca, 2010]. These models generate compression in west directed slabs and the induced shear produces extension in east directed slabs, a symmetry breaking effect that may alternatively be explained by regional variations in no net rotation models.

Table B1. Subduction Zone Parameters Used to Construct Profiles

<table>
<thead>
<tr>
<th>Region Name</th>
<th>Longitude (deg)</th>
<th>Latitude (deg)</th>
<th>Azimuth (°CW From N.)</th>
<th>Width (km)</th>
<th>Length (km)</th>
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<tr>
<td>Marianas</td>
<td>140</td>
<td>17</td>
<td>90</td>
<td>200</td>
<td>1000</td>
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<td>−20</td>
<td>120</td>
<td>150</td>
<td>900</td>
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<td>Japan</td>
<td>129</td>
<td>42</td>
<td>100</td>
<td>450</td>
<td>1700</td>
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<tr>
<td>Kurile</td>
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<td>52</td>
<td>140</td>
<td>650</td>
<td>1200</td>
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<td>Izu-Bonin North</td>
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<td>900</td>
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<td>137</td>
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<tr>
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<tr>
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<tr>
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<td>Banda</td>
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</tr>
<tr>
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<td>400</td>
<td>800</td>
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<td>Sumatra</td>
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<td>240</td>
<td>500</td>
<td>1200</td>
</tr>
<tr>
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<td>−58</td>
<td>90</td>
<td>300</td>
<td>700</td>
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<tr>
<td>Hindu Kush</td>
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<td>34</td>
<td>330</td>
<td>200</td>
<td>800</td>
</tr>
<tr>
<td>Luzon</td>
<td>126</td>
<td>20</td>
<td>270</td>
<td>100</td>
<td>1000</td>
</tr>
</tbody>
</table>

*a* Degrees clockwise from north.
These results point to the need for further study of combined high-resolution slab and keel models including realistic plate margins and proper match to relative and absolute plate velocities. Our study shows that global circulation models explain first-order features in slab seismicity and so provide useful constraints on upper mantle dynamics.

Appendix A: Earthquake Distribution With Depth

We define intermediate depth as 100 km to 350 km and deep as 350 km to 700 km, based on the global distribution of gCMT solutions with depth. Figure A1 shows the centroid depth distribution for the gCMT catalog (Global CMT Project, 2006, available at http://www.globalcmt.org/, accessed May 2008) from 1976 to 2008. The seismicity minima at ∼350 km corresponds as well to the transition depth at which predominately extensional deformation becomes predominately compressional deformation, as noted by Isacks and Molnar [1971].

Appendix B: Subduction Zone Parameters

In Table B1 we provide the parameters used to construct our cross sections for the purpose of identifying the slab deformation state shown in Figure 2. Parameters were chosen to most closely replicate the cross sections of Isacks and Molnar [1971] but were occasionally adjusted when structural complexities required a distinct profile to investigate differences along strike.

Appendix C: Correlation Between 3-D Vector and Tensor

Figure C1 shows the correlation between the P axes, eigenvector misfits, as used to characterize model performance, and the more complete tensor inner product misfit. The eigenvector misfit is calculated as arccos(|e_{cmt} · e_{stress}|), where e_{cmt} is the compressional eigenvector of the gCMT solution and e_{stress} is the compressional eigenvector of the stress solution from the fluid model. The tensor misfit is calculated using the inner product as arccos(T_{cmt} : T_{stress}), where T_{cmt} is the gCMT and T_{stress} is the full stress tensor solution from the fluid model. Considering the correlation, we report only the more intuitive eigenvector misfits in Table 2 and the main text.

Appendix D: Resolution Tests

In an effort to evaluate the error associated with the numerical resolution, we compare various upper mantle resolution models and find that our computations converge under successive refinement. Global misfit differences of 0.2° are within the error of the upper mantle element spacing (Figure D1).
Table E1. Misfit Results in Degrees as a Function of Increasing Net Rotation Component, From NNR to HS3

<table>
<thead>
<tr>
<th>(\times \eta_L)</th>
<th>(\eta_{\text{frac}}/\eta_B = 0.1): Net Rotation Magnitude and Global Misfit (deg)</th>
</tr>
</thead>
<tbody>
<tr>
<td>NNR</td>
<td>0.25 NR</td>
</tr>
<tr>
<td>1</td>
<td>100</td>
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<tr>
<td>1</td>
<td>30</td>
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<td>1</td>
<td>100</td>
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<tr>
<td>1000</td>
<td>100</td>
</tr>
</tbody>
</table>

*Models with net rotation have a fixed CMB while NNR models have a free slip CMB. Models have a 0.1\(\eta_{\text{frac}}/\eta_B\) applied globally. We also evaluated an NNR model with a fixed CMB and found that the results matched to within the error of our resolution the results of the same model with a free slip CMB.

CitcomS computations were performed on a multi-grid solver with a minimum overall accuracy of \(10^{-3}\) for the divergence in the Uzawa iterations for incompressibility.

[62] For our study, the vertical resolution would most likely have an influence on our phase change models as the phase change transition width (20 km) is similar to our mesh resolution in the transition zone (~18 km). To address this, we performed a model with a 40 km transition width. The global misfit difference between the models (0.15°) is within our overall resolution. Regional differences were insignificant, with a maximum difference of <1.9°, minimum of zero, and average differences of 0.33°.

Appendix E: Global Model Misfit as a Function of Increasing Net Rotation Component

[63] To test the effect of net rotation magnitude on global eigenvector misfit, we tested a suite of models with a globally applied asthenosphere and prescribed surface velocities with variable amounts of net rotation. Our results indicate that the high net rotation absolute plate motions of HS3 produce the best match to gCMTs (Table E1).

Acknowledgments

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