

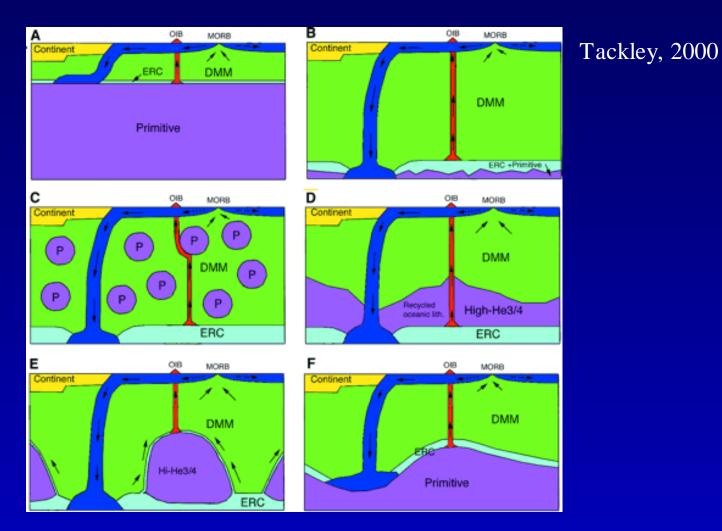




## Noble Gas Constraints on Mantle Structure and Convection

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## Which (if any) of these views of the mantle are correct?



What are the constraints on mantle reservoirs provided by the noble gases?

- what we know, what we infer, and what we do not understand

## Outline

- Helium, Neon, and Argon isotopic composition of Mid Ocean Ridge Basalts (MORBs) and Ocean Island Basalts (OIBs) – observations and constraints
- Missing Argon problem
- He Heat paradox
- Combined noble gases and lithophile tracers (Sr, Nd, Pb)
- Noble gas concentrations and elemental ratios- challenges in reconciling isotopic and elemental composition

## He isotope geochemistry

- Two isotopes of helium: <sup>3</sup>He and <sup>4</sup>He
  <sup>3</sup>He is primordial
  <sup>4</sup>He produced by radioactive decay of U and Th
- He isotopes are a measure of time-integrated (U+Th)/<sup>3</sup>He ratio:

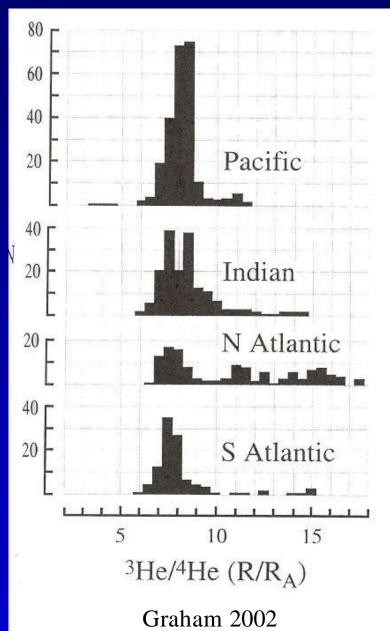
$$\frac{{}^{4}\text{He}}{{}^{3}\text{He}} = \left(\frac{{}^{4}\text{He}}{{}^{3}\text{He}}\right)_{o} + 8\frac{{}^{238}\text{U}}{{}^{3}\text{He}}\left(e^{\lambda_{238}t} - 1\right) + 7\frac{{}^{235}\text{U}}{{}^{3}\text{He}}\left(e^{\lambda_{235}t} - 1\right) + 6\frac{{}^{232}\text{Th}}{{}^{3}\text{He}}\left(e^{\lambda_{232}t} - 1\right)$$

- Helium behaves as an incompatible element during mantle melting (i.e. prefers melt over minerals)
- Helium *expected* to be more incompatible than U and Th during mantle melting
- Helium not recycled back into the mantle

If so high <sup>3</sup>He/<sup>4</sup>He ratios reflect less degassed mantle material

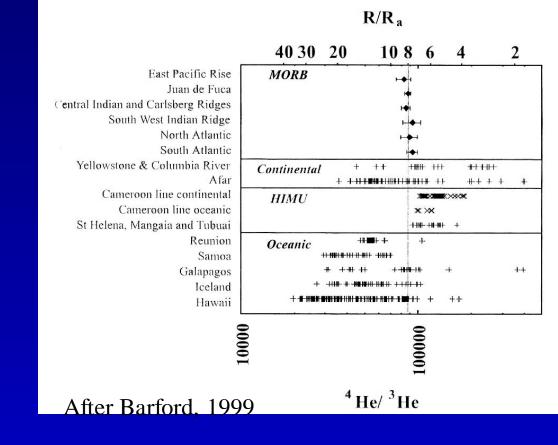
#### Histogram of He isotope ratios in mid-ocean ridge basalts (MORBs)

- <sup>3</sup>He/<sup>4</sup>He ratios reported relative to the atmospheric ratio of 1.39 x 10<sup>-6</sup>
- No relation between isotopic composition and spreading rate but the variance is inversely related to spreading rate
- Either reflects
  - efficiency of mixing in the upper mantle
  - differences in degree of magma homogenization



#### Comparison of He isotope ratios from selected MORs, OIBs, and continental hotspots

- The mean <sup>3</sup>He/<sup>4</sup>He ratio from different ridge segments is nearly identical although the variance is different
- OIBs are much more variable
- <sup>3</sup>He/<sup>4</sup>He ratios less than MORBs are frequently associated with radiogenic Pb (HIMU) and reflects recycled components in the mantle



#### He isotope ratios in ocean island basalts (OIBs)

35

30

25

20

15

Glass

Pyroxene

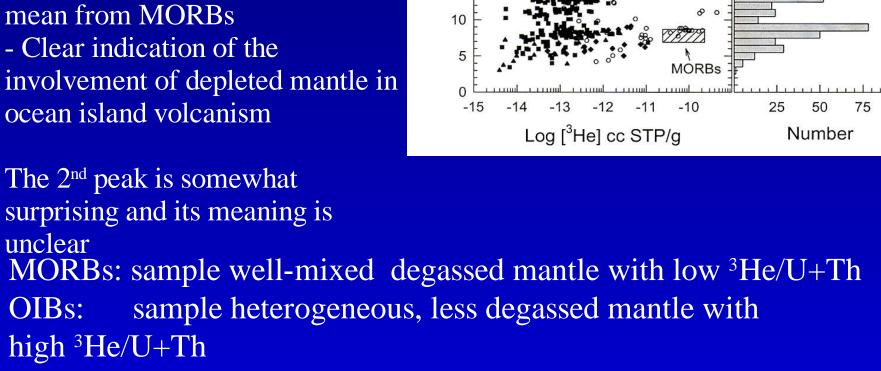
Xenolith

Farley and Neroda 1998

100

- OIBs display a very large range in  $\bullet$ He isotopic composition
- He isotopic distribution has a double-peak; maxima at 8  $R_A$  and  $13 R_{A}$
- 'He/<sup>4</sup>He (R<sub>A</sub>) The first maxima is identical to the mean from MORBs - Clear indication of the involvement of depleted mantle in ocean island volcanism

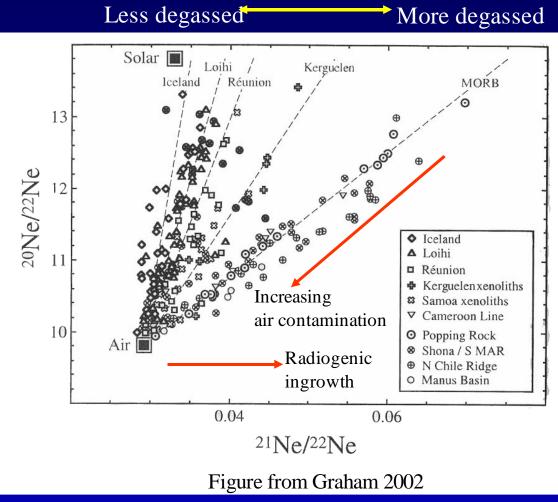
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## Geochemistry of Ne

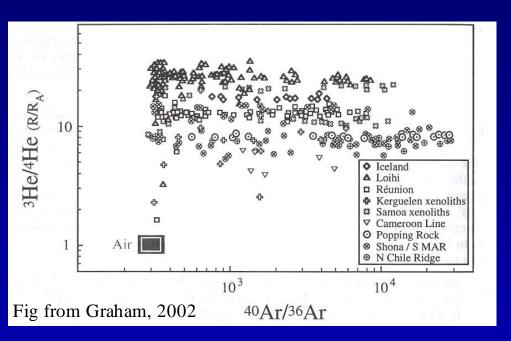
- Neon has three isotopes <sup>20</sup>Ne, <sup>21</sup>Ne, and <sup>22</sup>Ne
- <sup>20</sup>Ne is primordial
- <sup>21</sup>Ne is produced by nucleogenic reactions in the mantle:
  - ${}^{18}O(\alpha, n){}^{21}Ne$  and  ${}^{24}Mg(n, \alpha){}^{21}Ne$
  - $\alpha \text{ from U decay; neutrons from spontaneous fission;}$  $production ratio of ^21Ne/4He is ~10^{-7}$
- <sup>22</sup>Ne is primordial. There may be a small nucleogenic production of
  <sup>22</sup>Ne, [<sup>19</sup>F(α, n)<sup>22</sup>Ne] but it is likely to be negligible
- <sup>20</sup>Ne/<sup>22</sup>Ne does not vary in the mantle derived rocks; <sup>21</sup>Ne/<sup>22</sup>Ne does
- Ne is expected to be more incompatible than U and Th during mantle melting => low <sup>21</sup>Ne/<sup>22</sup>Ne ratios reflect less degassed mantle material

## Ne isotopic composition of mantle derived rocks

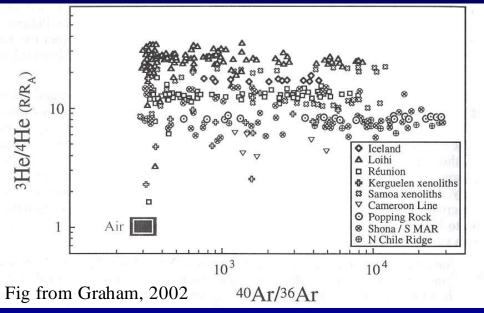


- Mantle <sup>20</sup>Ne/<sup>22</sup>Ne ratio is fixed; <sup>21</sup>Ne/<sup>22</sup>Ne varies because of radiogenic ingrowth and varying degrees of degassing
- Different ocean islands have distinct <sup>21</sup>Ne/<sup>22</sup>Ne ratios; either reflects varying amounts of MORB mantle addition to the OIB source(s) or different parts of the mantle have been degassed and processed to different degrees

- Three stable isotopes of Ar, <sup>36</sup>Ar, <sup>38</sup>Ar, <sup>40</sup>Ar
- <sup>36</sup>Ar and <sup>38</sup>Ar are primordial
- <sup>40</sup>Ar produced by radioactive decay of <sup>40</sup>K
- Ar is expected to be more incompatible than K during mantle melting
- If so high <sup>40</sup>Ar/<sup>36</sup>Ar reflects degassed mantle material

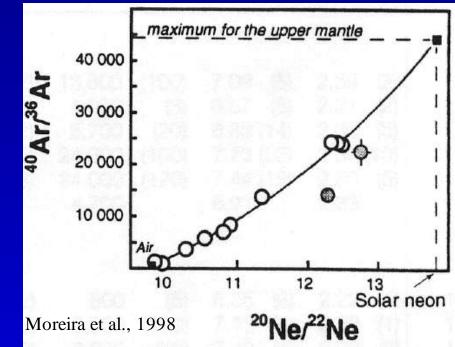


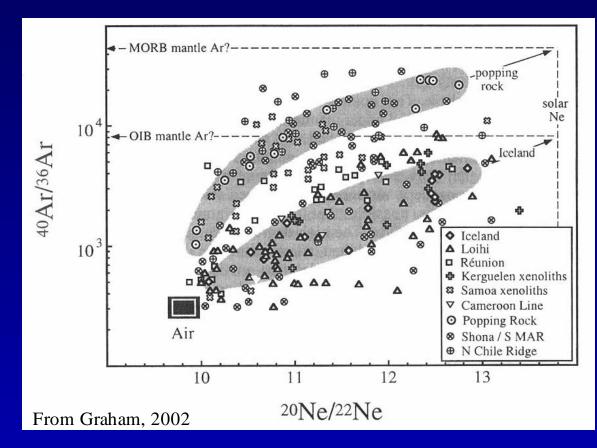
- 1% Ar in the atmosphere
- Significant air contamination for Ar
- Even when <sup>3</sup>He/<sup>4</sup>He ratios are as high as 30 R<sub>A</sub>, <sup>40</sup>Ar/<sup>36</sup>Ar ratios can be atmospheric



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- <sup>20</sup>Ne/<sup>22</sup>Ne ratio in the mantle does not vary
- Ar isotopic ratios in mantle derived rocks can be corrected for air contamination by extrapolating the <sup>40</sup>Ar/<sup>36</sup>Ar ratio to the upper mantle <sup>20</sup>Ne/<sup>22</sup>Ne value





- MORB mantle  ${}^{40}$ Ar/ ${}^{36}$ Ar values are ~ 40,000
- OIBs have lower <sup>40</sup>Ar/<sup>36</sup>Ar ratios; reasonable limit is 8000
- A value of 8000 *does not* represent pristine mantle material; must indicate some processing, although significantly less degassed than the mantle source sampled by MORBs

#### The picture that emerges so far.....

- 2. MORBs are more homogenous compared to OIBs
- 4. Many OIBs sample a mantle source that is significantly less degassed than the mantle source tapped by MORBs

# Evidence for undegassed reservoir: The missing Argon problem

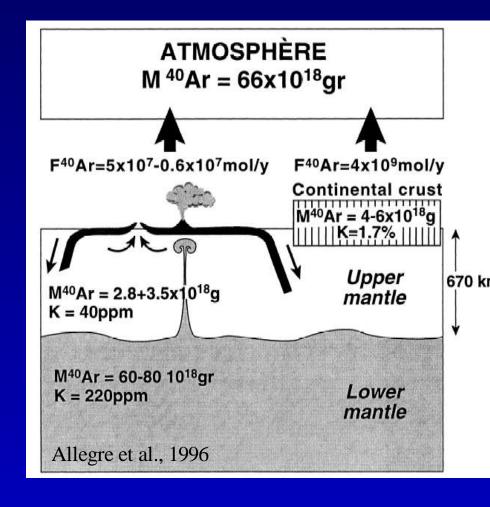
- K content of Earth derived from the K/U ratio of 12700 in MORBs and U content of 20-22.5 ppb
- Implied K content of bulk Earth is 250-285 ppm
- Total <sup>40</sup>Ar produced over Earth history =  $140-156 \times 10^{18} \text{ g}$
- ${}^{40}$ Ar in the atmosphere = 66 x 10<sup>18</sup> g (~50%)
- ${}^{40}$ Ar in the crust = 9-12 x 10<sup>18</sup> g

63-80 x 10<sup>18</sup> g of <sup>40</sup>Ar has to be in the mantle

# Evidence for undegassed reservoir: The missing Argon problem

## I) Constraints from <sup>40</sup>Ar flux

- <sup>4</sup>He flux at ridge = 9.46 x 10<sup>7</sup> moles/yr
- <sup>4</sup>He/<sup>40</sup>Ar ratio in MORBs 2-15
  => <sup>40</sup>Ar flux 0.63-5 x 10<sup>7</sup> moles/yr
- Mass of oceanic lithosphere passing through ridges = 5.76 x 10<sup>17</sup> g/yr
   If MORB mantle representative of entire mantle and if lithosphere completely degassed, <sup>40</sup>Ar content in mantle 1.4-1.8 10<sup>18</sup>g
  - Lower than the 63-81 x 10<sup>18</sup>g estimated (Allegre et al., 1996) and requires a hidden reservoir for <sup>40</sup>Ar



If MORB mantle extends to 670 km, 0.6-4.6 x 10<sup>18</sup> g of <sup>40</sup>Ar in upper mantle and 59 x 10<sup>18</sup> g of <sup>40</sup>Ar in the lower mantle, corresponding to a K concentration of about 230ppm; consistent with K content of bulk Earth

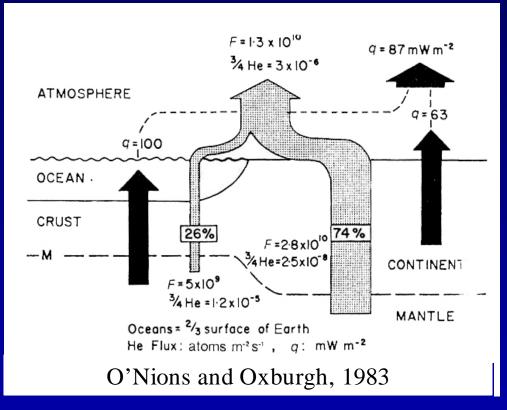
## Evidence for undegassed reservoir: The missing Argon problem

II) Constraints from Potassium content

- K content of MORB source is 40-50 ppm; if representative of entire mantle produces 22-28 x 10<sup>18</sup> g of <sup>40</sup>Ar
  -significantly less than the 63-80 x 10<sup>18</sup> g of <sup>40</sup>Ar calculated to be in the mantle
- If mantle is layered at 670 km 7.3-9. x 10<sup>18</sup> g <sup>40</sup>Ar in the upper mantle
  => 54-74 x 10<sup>18</sup> g in the lower mantle, corresponding to an K content of 230 ppm in the lower mantle

Bottom line: The constraints from <sup>40</sup>Ar require some sort of layering or a hidden reservoir in the mantle Any wiggle room? Maybe we do not know the K/U ratio of the mantle as well as we think (e.g., Albarede, 1998; Lassiter, 2002)

## The Helium Heat Paradox



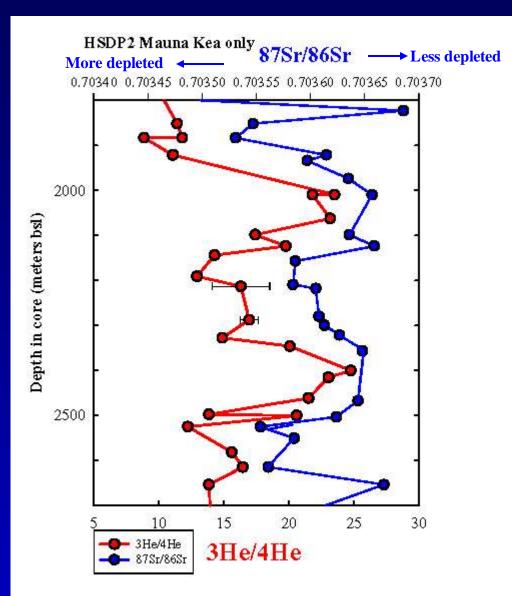
- ~75% of He entering the atmosphere is from continental crust
- ~25% from the mantle
- ~10% of the He from the mantle is primordial and the rest is radiogenic

#### The Helium Heat Paradox

- <sup>4</sup>He produced by radioactive decay of U and Th
- 10<sup>-12</sup> J of energy is liberated for each alpha decay
- The radiogenic <sup>4</sup>He flux from the mantle corresponds to 2.4 TW of heat production
- Terrestrial heat flux is 44 TW (Pollack et al., 1993) -- 5-10 TW from crust (e.g., Rudnick and Fountain, 1995), 3-7 from core (Buffett et al., 1996); and 27-36 TW from the mantle
- Of 27-36 TW from the mantle, 18–22 is secular cooling; radiogenic heat is between 9-14 TW, factor of 4-6 greater than the 2.4 TW of heat that is supported by the <sup>4</sup>He flux
- Implies a boundary layer in the mantle that passes heat but mostly retains <sup>4</sup>He (O'Nions and Oxburgh, 1983)

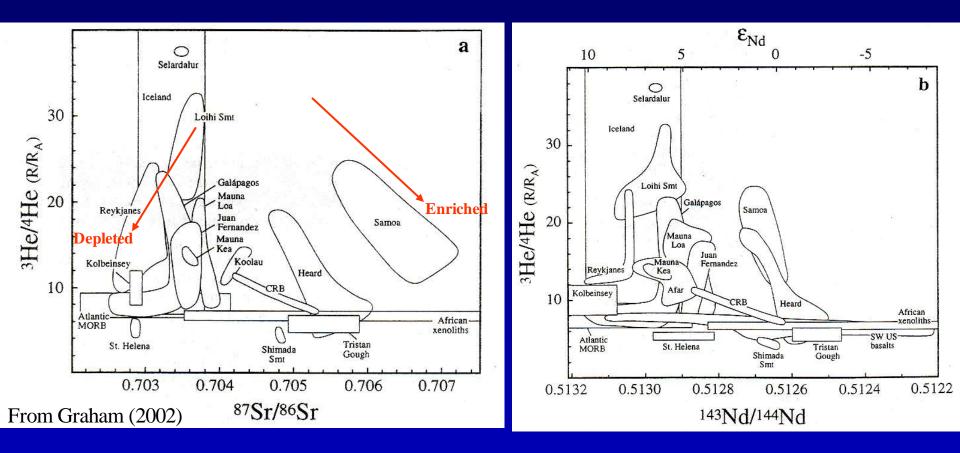
Relationship between He and other lithophile tracers

- He isotopic variations are strongly coupled to variations in other lithophile tracers (Sr, Nd, Pb)
- Higher <sup>3</sup>He/<sup>4</sup>He ratios are associated with *less* depleted <sup>87</sup>Sr/<sup>86</sup>Sr isotopic signal
- ⇒ high <sup>3</sup>He/<sup>4</sup>He ratios are indicative of less degassed mantle

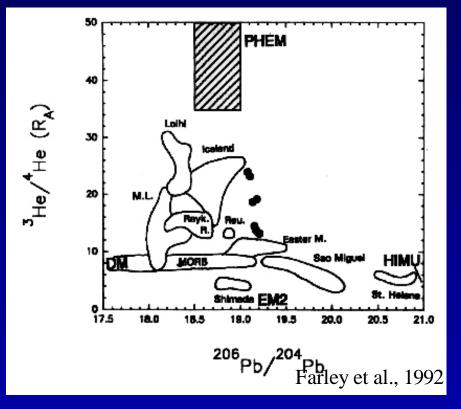


Above data is from the 3 km deep drill hole from Mauna Kea, Hawaii (Kurz et al., 2004)

## Global relationship between He and other lithophile tracers: The wormograms



## Global relationship between He and other lithophile tracers



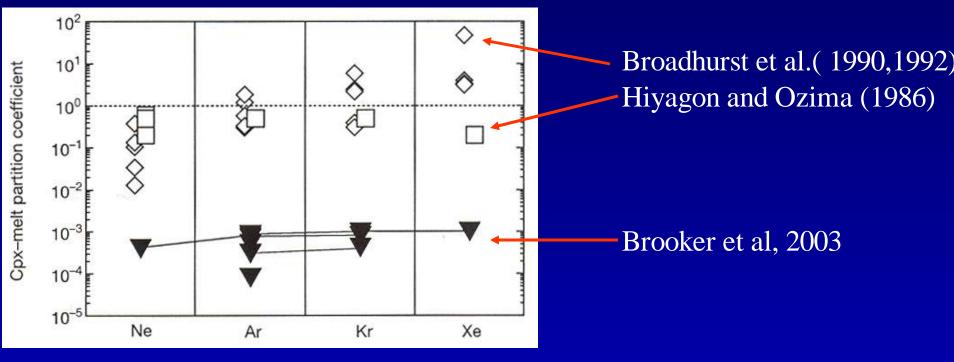
- Observations:
- Highest <sup>3</sup>He/<sup>4</sup>He ratios occur at intermediate values of Sr, Nd, and Pb and not associated with either the most depleted or most enriched mantle end-members
- Isotopic arrays from individual ocean islands convergence to a composition that is internal to the mantle end-members defined in Sr, Nd, and Pb isotopic space

Global relationship between He and other lithophile tracers

Inferences:

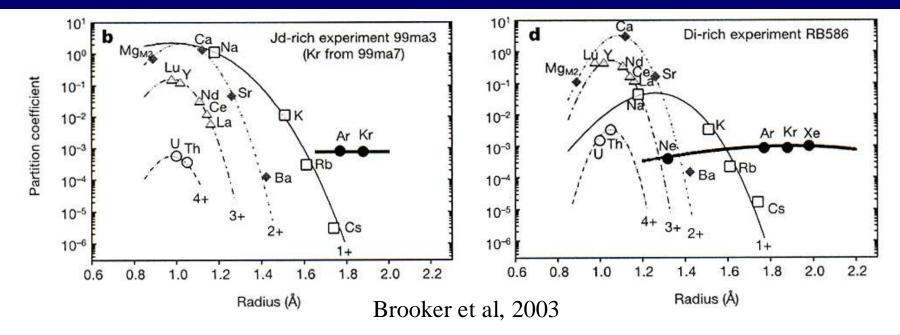
- High <sup>3</sup>He/<sup>4</sup>He ratios from a single, relatively undegassed mantle source that is characterized by well defined Sr, Nd, and Pb isotopic composition
- <sup>3</sup>He/<sup>4</sup>He is one of the reasons to come up with a component (PHEM, FOZO, C) that is internal to the other mantle endmembers in Sr, Nd, and Pb isotopic space (EM1, EM2 HIMU, DM)
- FOZO/C has Sr, Nd, Pb isotopic composition that is slightly depleted in comparison to primitive mantle; PHEM is primitive
- But there appears to be a problem: mixing hyperbolas seem to curve the wrong way... undegassed reservoir should have higher He concentration

### Partition coefficient for noble gas



Cpx-melt partition coeffcients

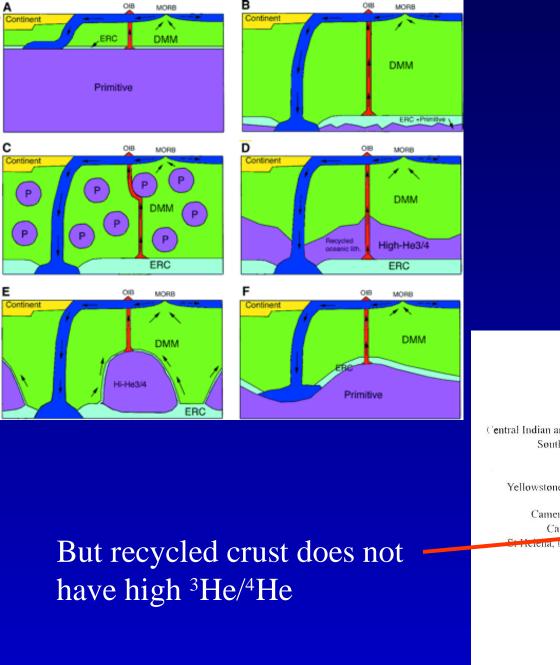
Hard to determine experimentally because of the formation of fluid inclusion; noble gases will prefer a fluid over a melt But ability to measure partition coefficient getting better **Partition coefficient of the noble gases** Are the noble gases (<sup>4</sup>He, <sup>21</sup>Ne, <sup>40</sup>Ar) really more incompatible than their radiogenic parents (e.g., U, Th, K)?

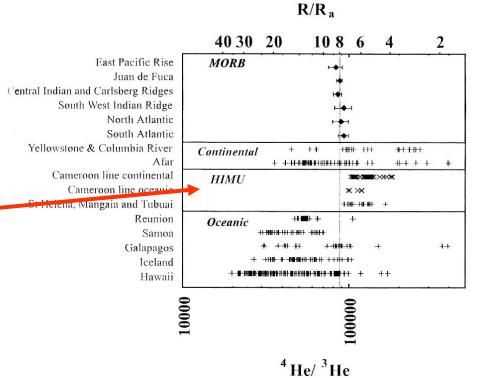


- For clinopyroxene, Ar slightly more incompatible than K
- Experimental data still not good enough to show conclusively how He behaves with
  - respect to U and Th
- Time integrated ratios however provide some insights; for example high <sup>3</sup>He/<sup>4</sup>He

ratios are never associated with the most depleted isotopic signatures of Sr and

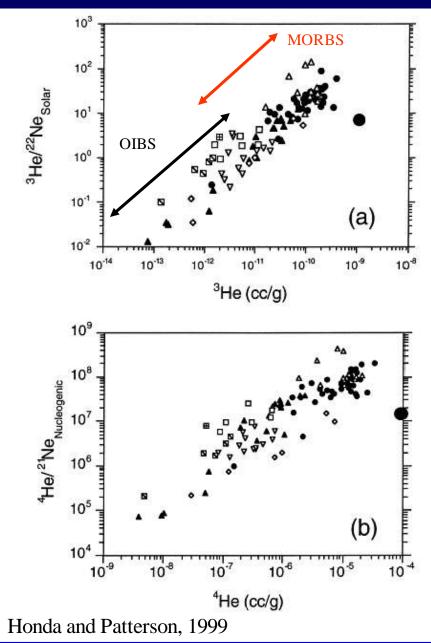
Nd,

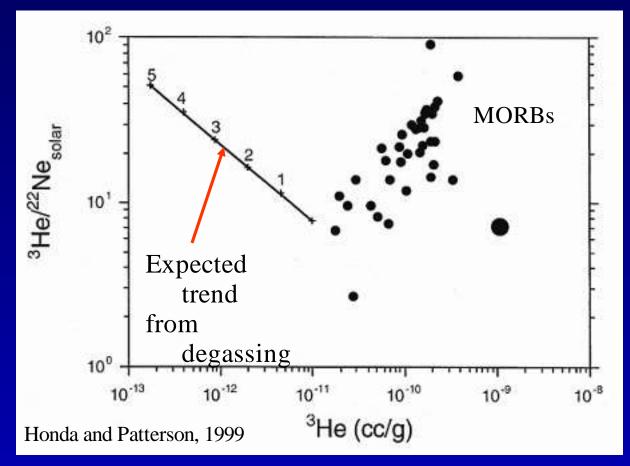




#### **Noble Gas Concentrations**

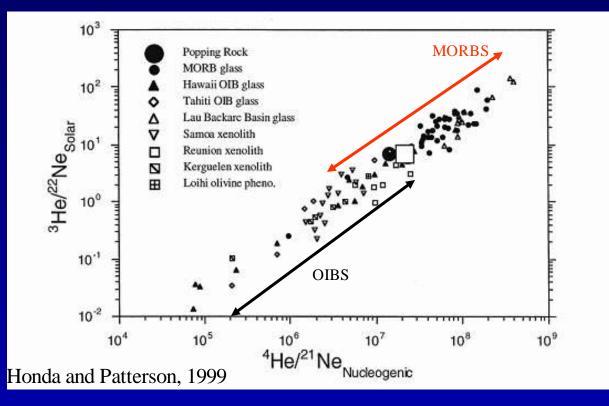
- Previously noted that based in the curvature of mixing hyperbolas He concentrations might be higher in MORBs than OIBs
- The figures show that is indeed the case
- Maybe not too surprising since most OIBs are erupted at shallower water depths than MORBs; so would be degassed more
- Turns out that such an explanation is not really tenable...



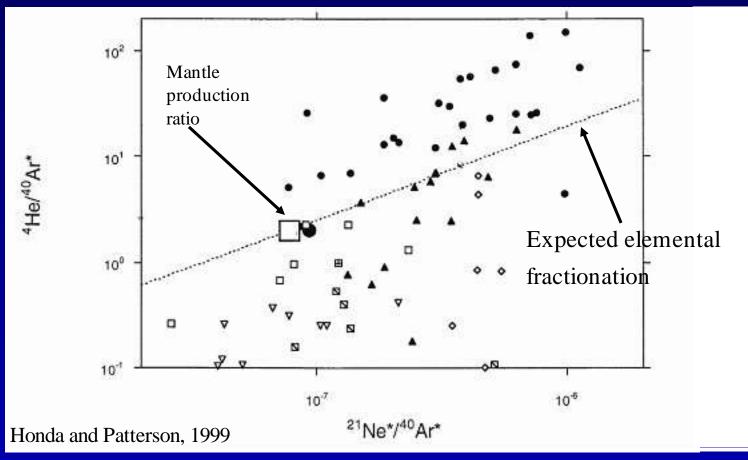


He more soluble in basaltic melt than Ne, which is more soluble than Ar => With increasing degassing He/Ne ratios increase and Ne/Ar ratios increase

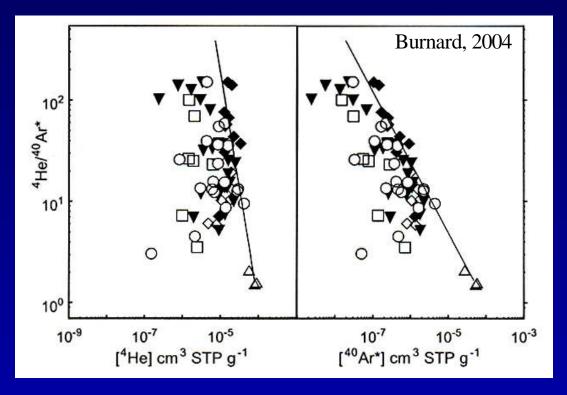
Solubility controlled degassing does not explain the differences in gas concentration between MORBs and OIBs.



 Fractionation has to be recent, otherwise the slope would not be 1 and the <sup>4</sup>He/<sup>21</sup>Ne ratio would have evolved back to the production value of ~10<sup>7</sup>



- OIBs are not depleted in He; rather MORBs are enriched in He with respect to Ne and Ar
- Enrichment of He with respect to Ar significantly greater than predicted for simple mass dependent process



- The Ar concentration decreases with degassing as expected
- For some MORBs suites, as <sup>4</sup>He/<sup>40</sup>Ar ratio increases (more degassed), He concentration increases as well!! So the problem appears to be with He

Open question: What controls the noble gas concentration and elemental ratios in oceanic basalts?

## Conclusion

- From <sup>3</sup>He/<sup>4</sup>He, <sup>21</sup>Ne/<sup>22</sup>Ne, and <sup>40</sup>Ar/<sup>36</sup>Ar ratios we know there exists a *relatively undegassed* reservoir in the Earth that is tapped at many ocean islands; MORBs sample a more degassed and processed mantle source
- An undegassed reservoir is supported by <sup>40</sup>Ar and the helium heat paradox
- Based on correlations between <sup>3</sup>He/<sup>4</sup>He ratios and other isotopic tracers (Sr, Nd, Pb), the undegassed reservoir has the composition of primitive mantle, or is slightly depleted relative to primitive mantle; the slight depletion is consistent with the inferred <sup>40</sup>Ar/<sup>36</sup>Ar ratio of ~8000 for the OIB source

## Conclusion (continued)

- Noble gas elemental ratios indicate that MORBs have higher He concentration than OIBs that cannot be explained away by simple magmatic degassing
- The higher He concentration in MORBs reflects a recent enrichment of He, and not likely to be a characteristic of the MORB source itself. The physical mechanism through which MORBs acquire a high He concentration remains unidentified

## Important questions that need to be answered

- Are the noble gases really more incompatible than their radioactive parents U, Th, and K during partial melting? Can the noble gases be partitioned into the core?
- What physical mechanism(s) control gas loss during mid ocean ridge and ocean island volcanism? Can we ' see through' such gas loss processes to infer the concentrations in the different mantle reservoirs?
- What are the characteristics of the heavy noble gases in OIBs and what role does subduction play in recycling of the heavy noble gases (Ar-Xe)?