

Subduction zone dynamics and mantle flow

Thorsten W Becker

University of Southern California

Short course at Università di Roma TRE

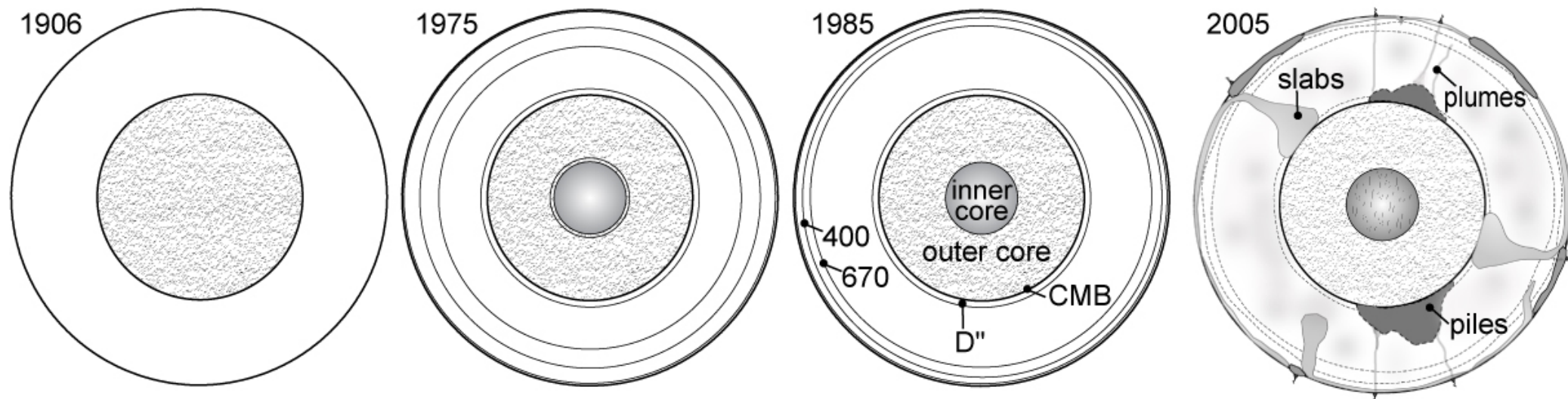
April 18 – 20, 2011

Revised syllabus

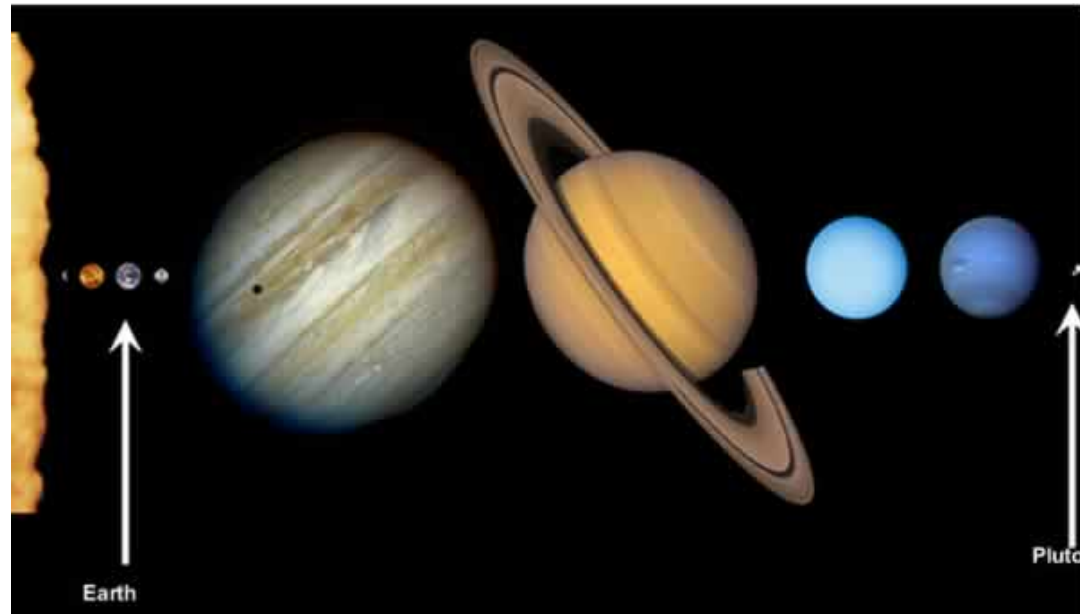
- Fundamentals of Earth Structure and Geophysical constraints for Subduction
- Fundamental of Fluid and Mantle Dynamics
- Regional Subduction Zone Modeling
 - From static to dynamic
- Global Mantle Flow Modeling
- Geological Constraints for Subduction
- Trench Migration and Upper Mantle Convection

Subduction I

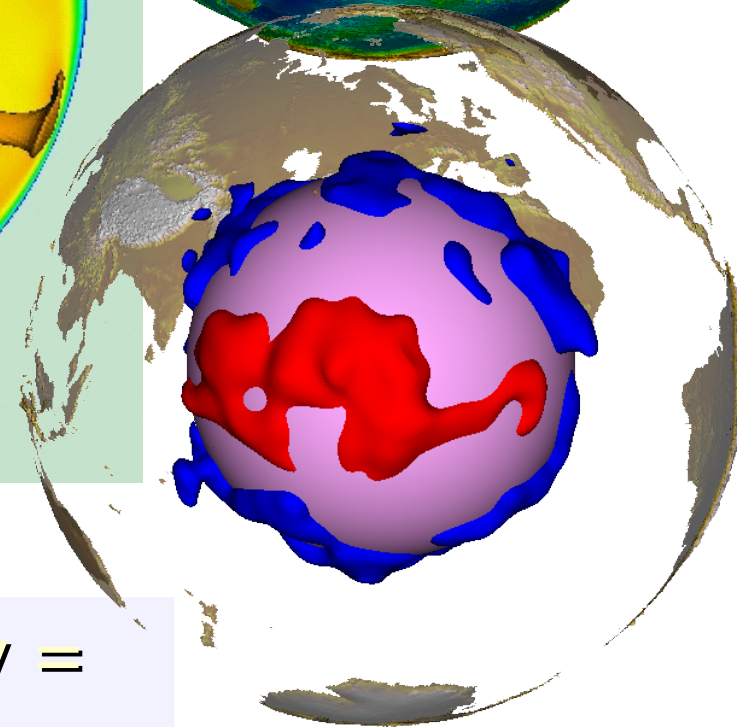
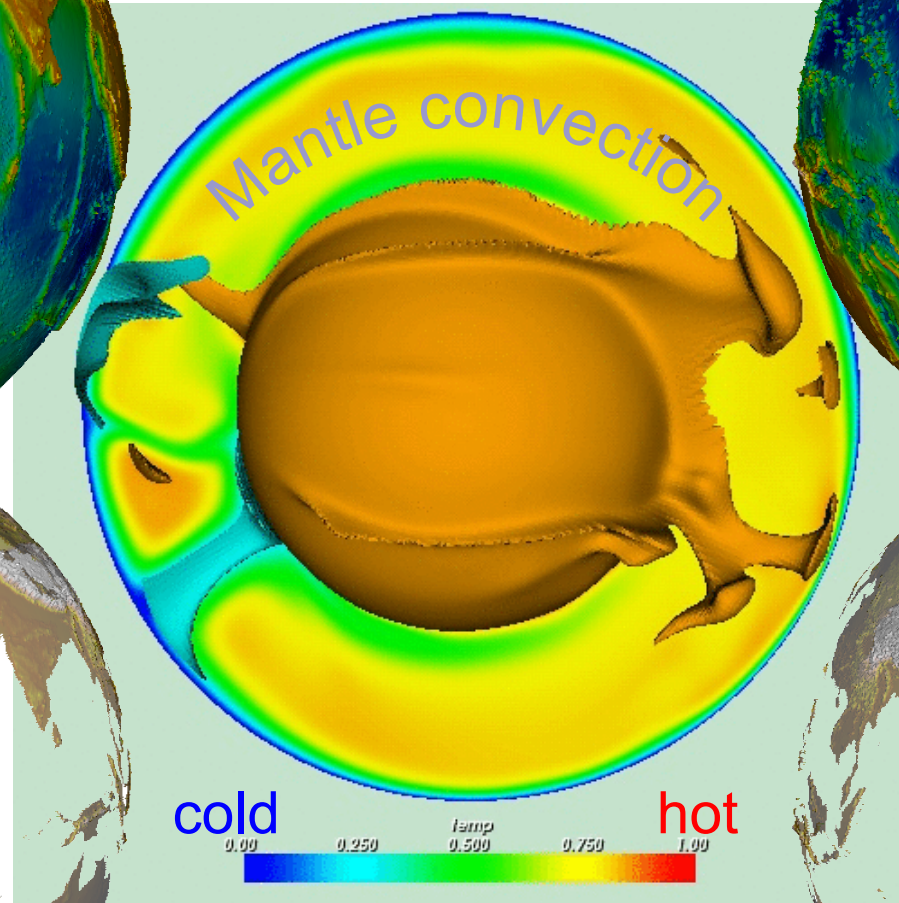
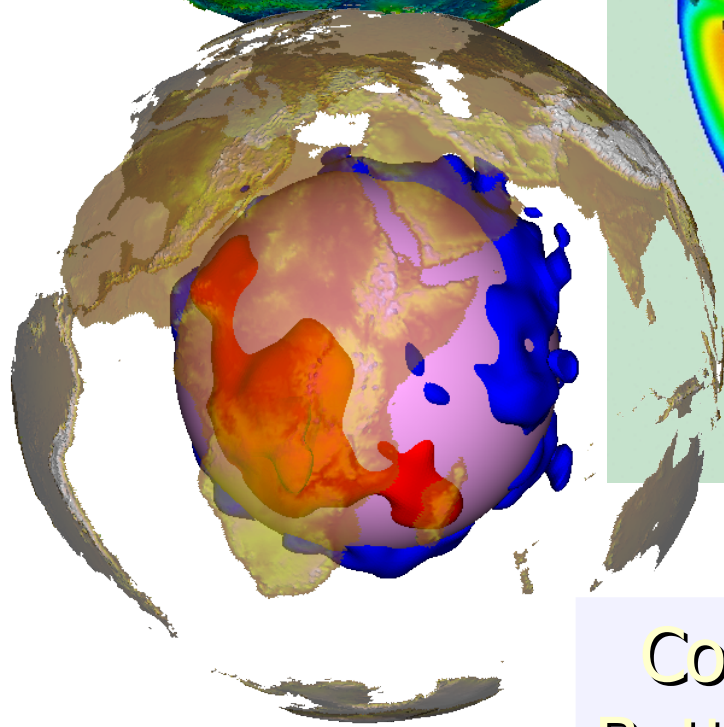
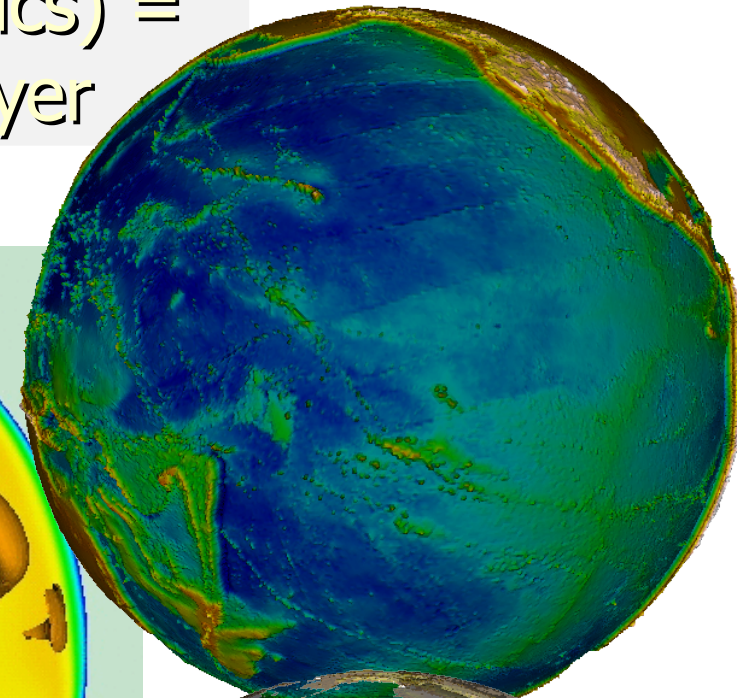
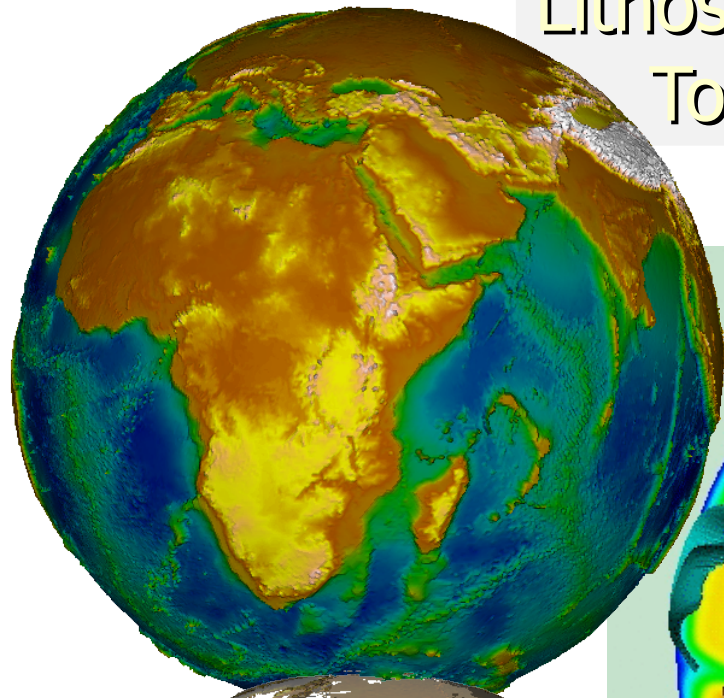
Earth Structure and geophysical subduction zone constraints



Rocky planets



Lithosphere (plate tectonics) =
Top, cold boundary layer



Core-mantle boundary =
Bottom, hot boundary layer

The mantle system

➤ **Dynamics of mantle largely thermally driven**

Lithosphere:

top thermal boundary layer of the mantle

rheology(p, T, C) →

strong, elastic with yielding (brittle/plastic)

plate tectonics

Heat sources

loss of accretional heat of mantle and core

radioactive heat production in mantle and core

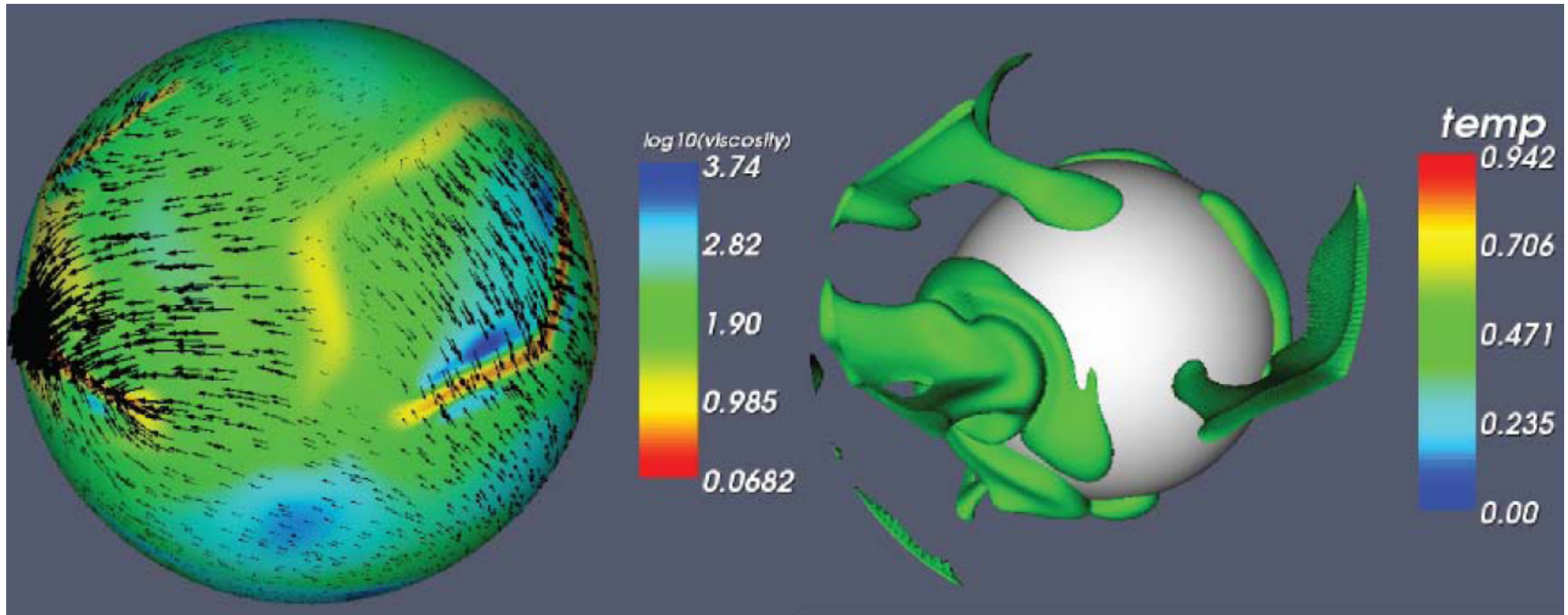
➤ **Role of composition?**

melting → compositional variations (crust)

affects geochemistry? mantle dynamics?

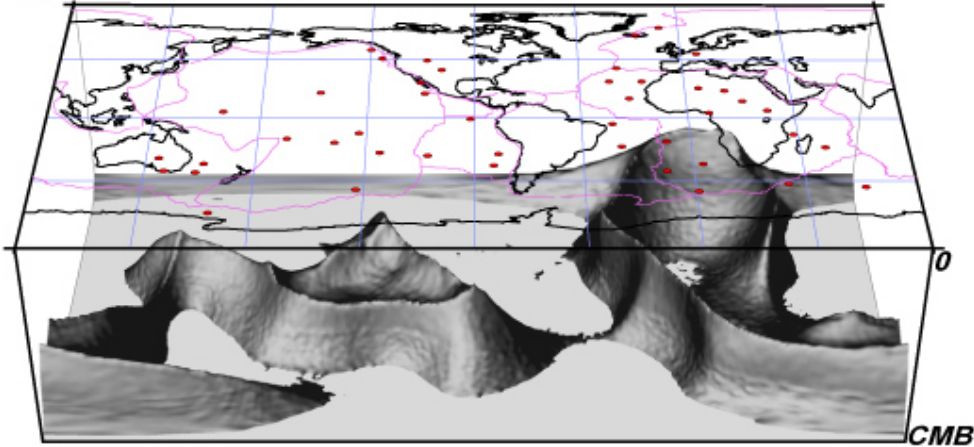
solidifying inner core → chemical source of energy

Some issues: 1) Plate generation

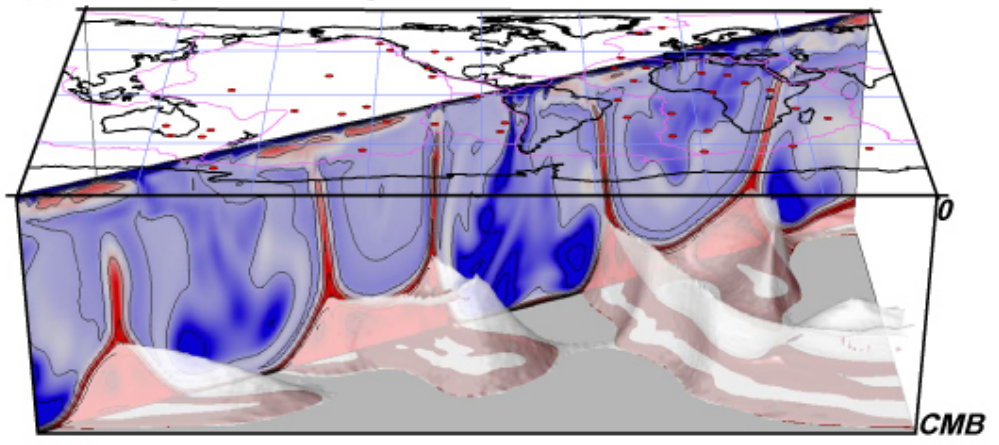


2) Thermo-chemical mantle structure

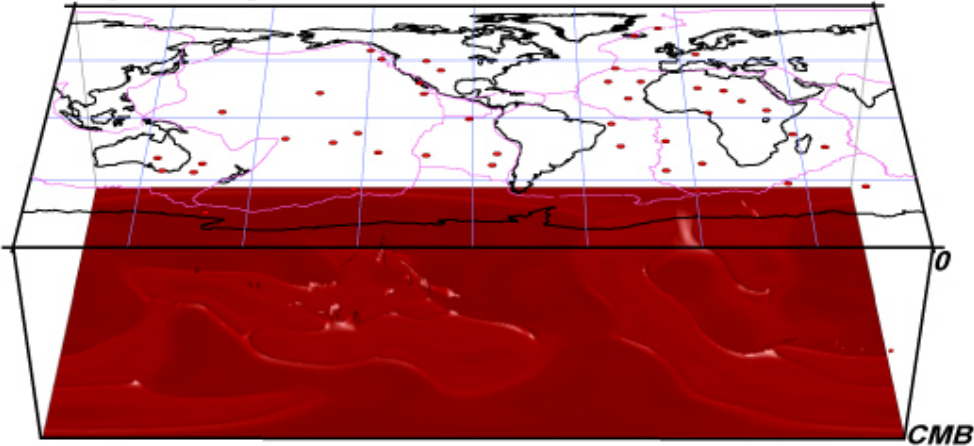
(a) Compositionally distinct, dense piles



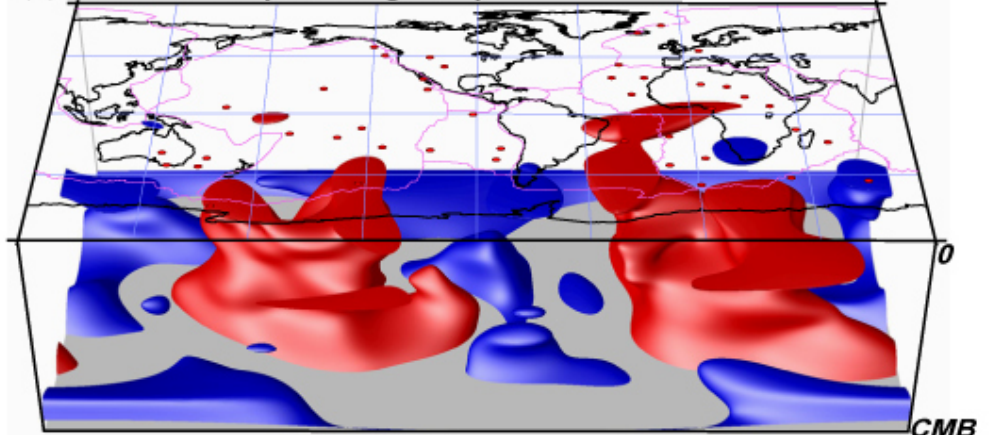
(c) Dense piles and temperatures



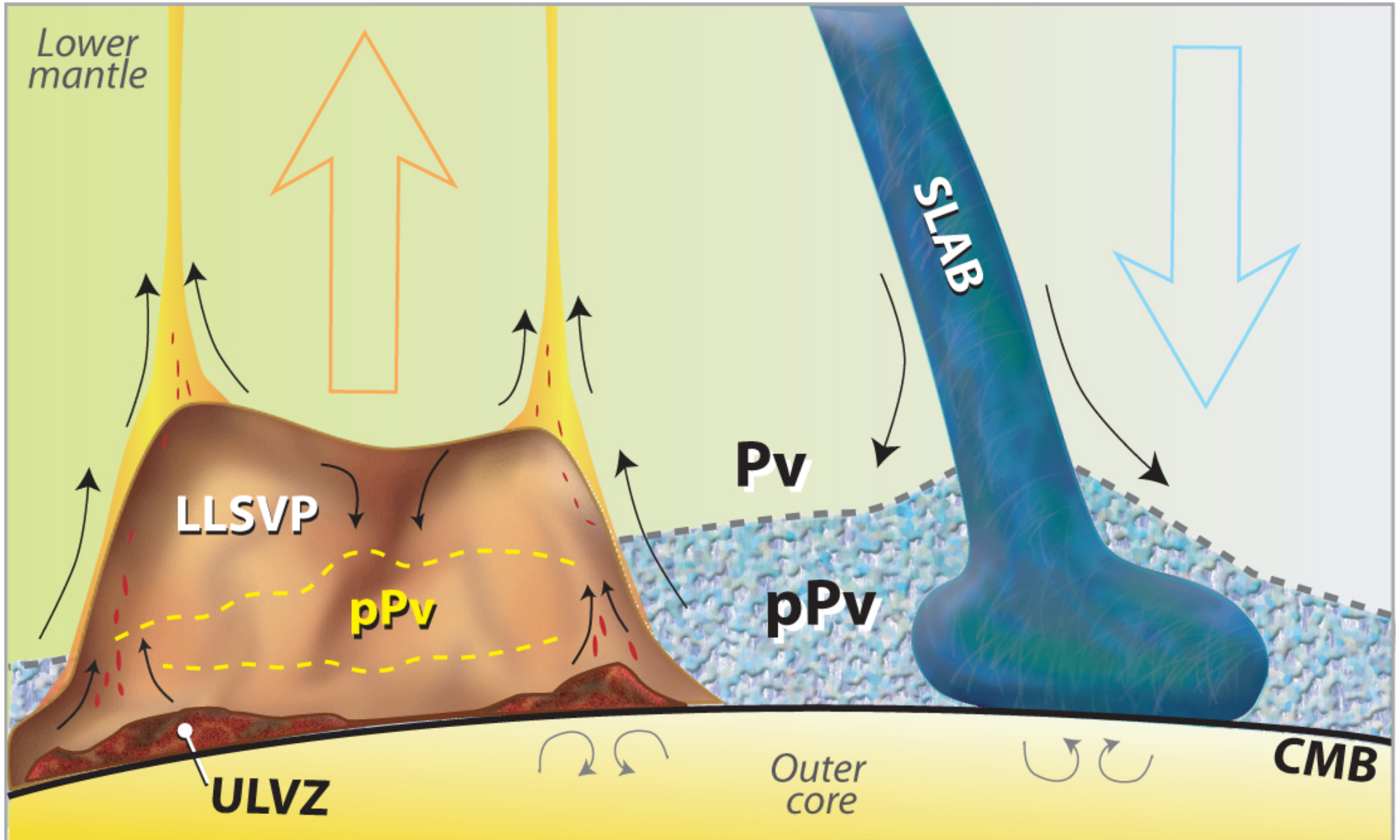
(b) Highest temperatures



(d) Shear velocity heterogeneity



3) Post-perovskite and ULVZs



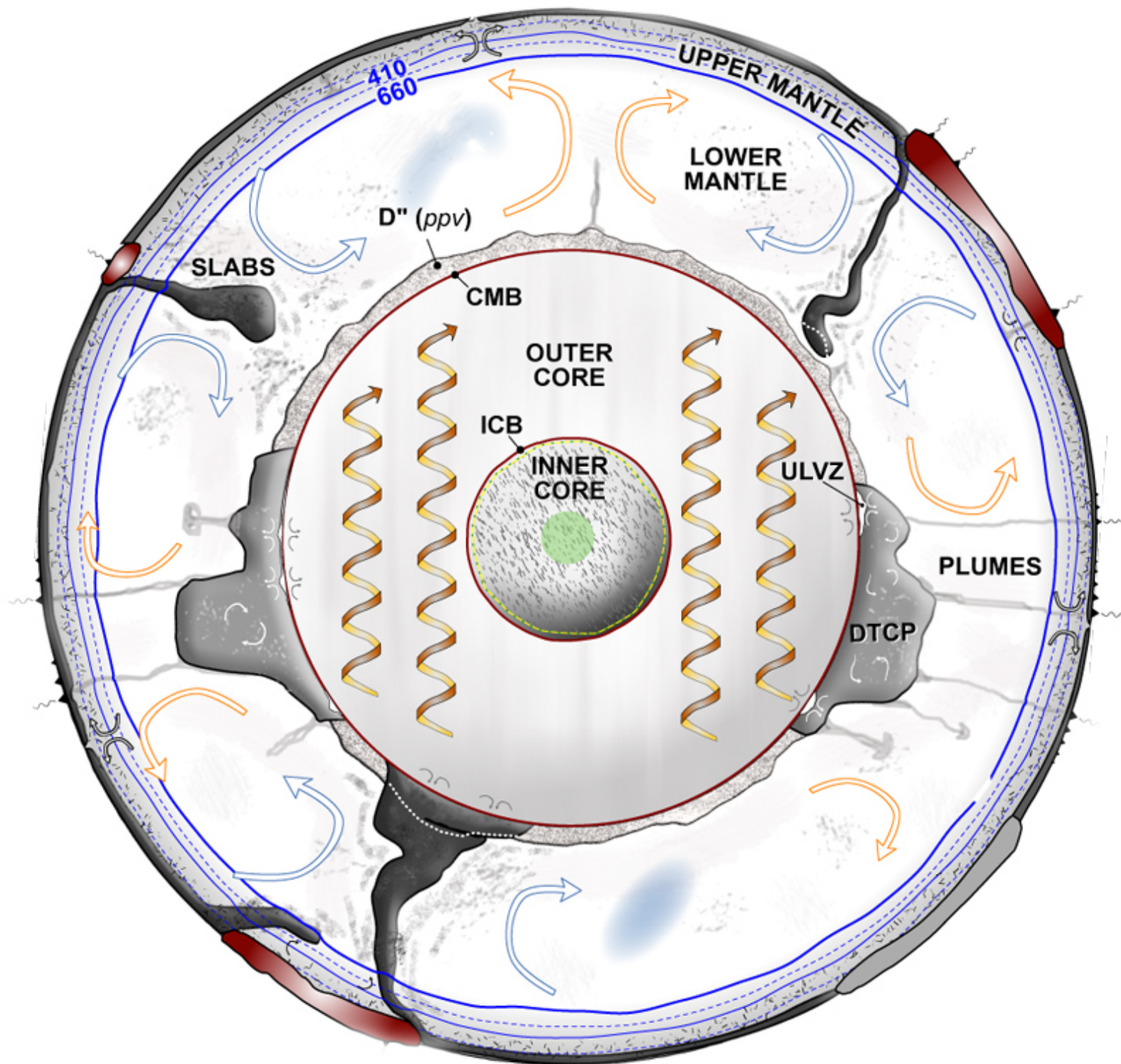
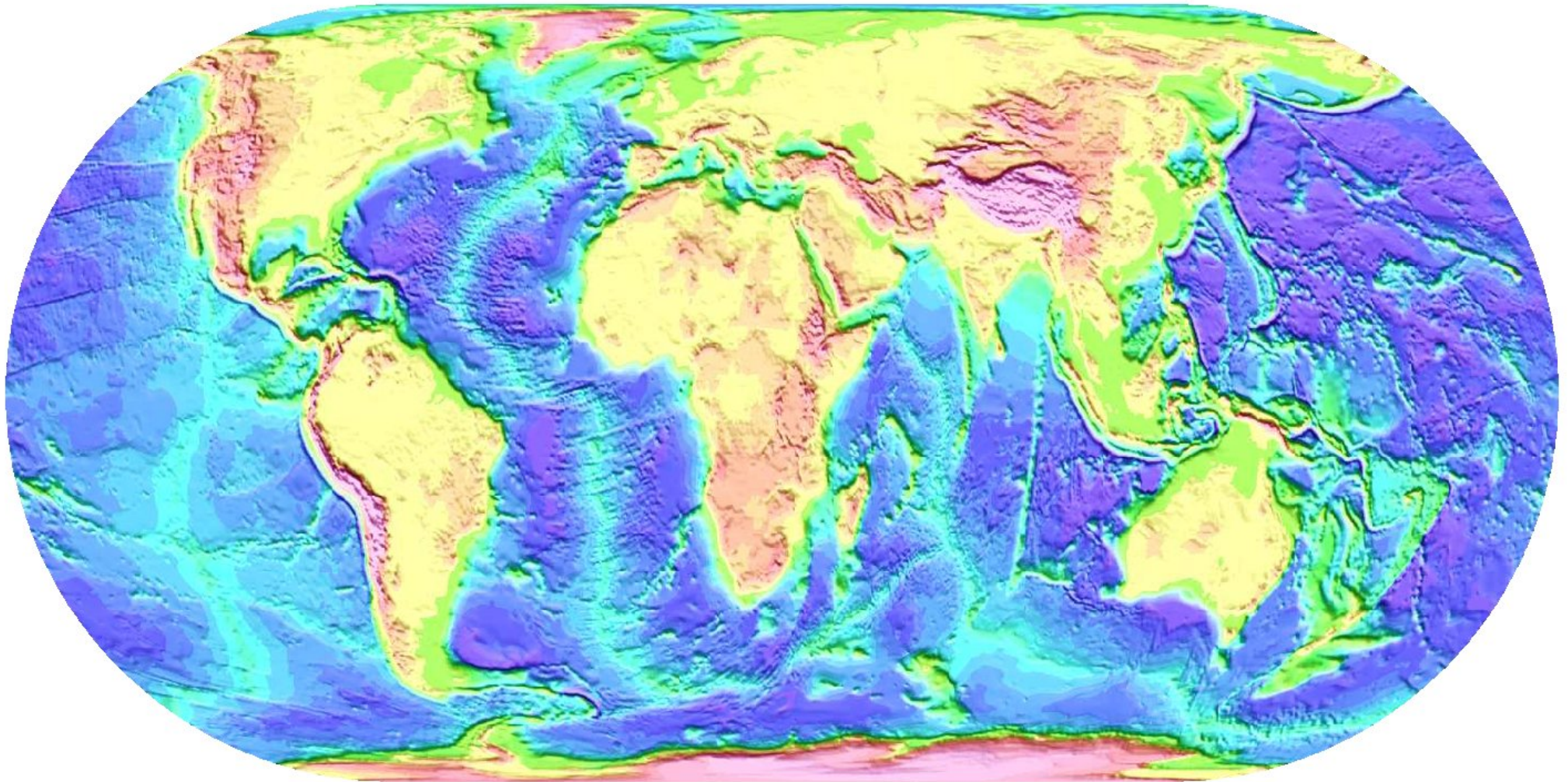
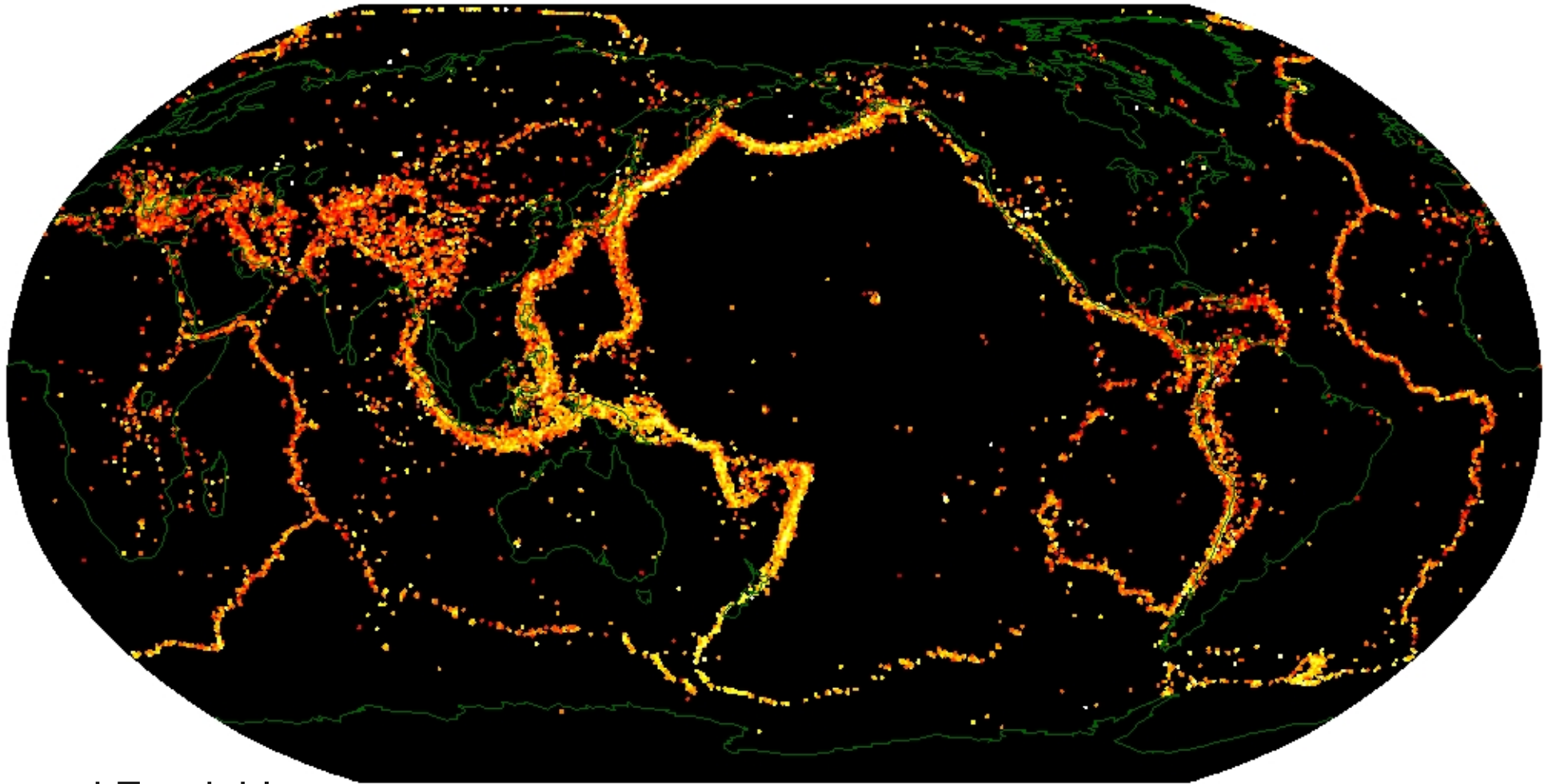


Figure courtesy Ed Garnero

Plate tectonics



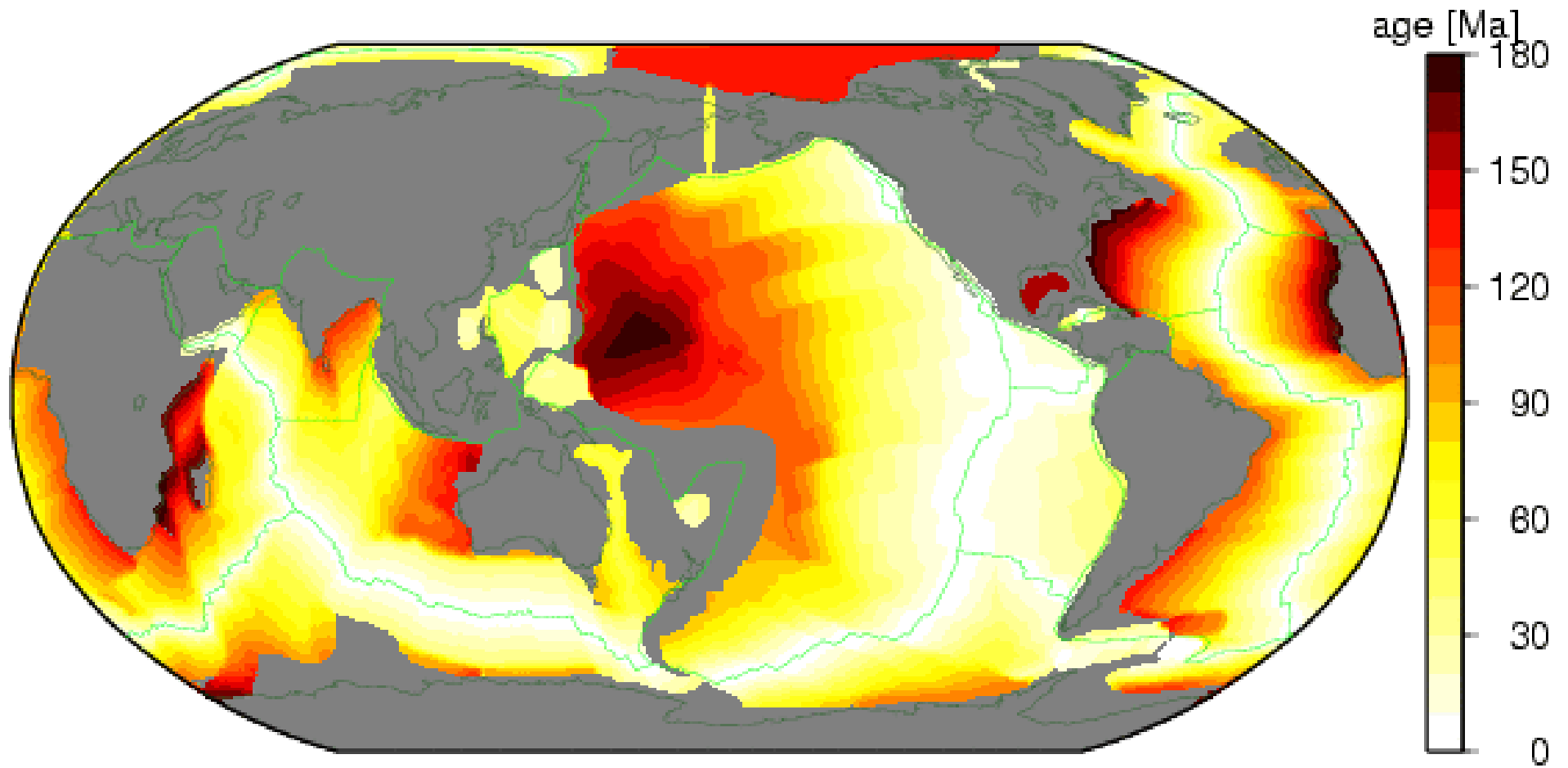
Global shallow seismic moment release



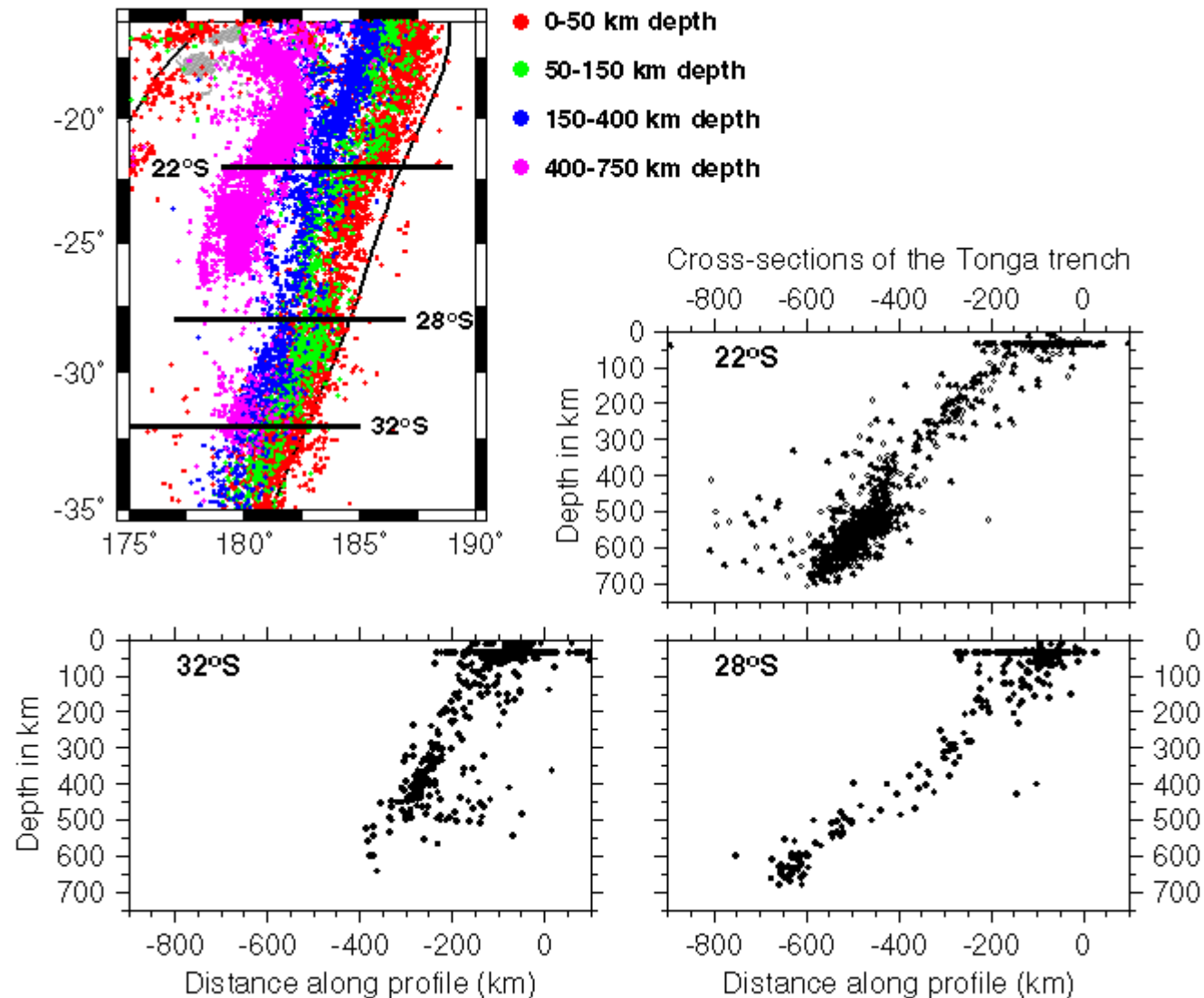
summed Engdahl
catalog < 50 km

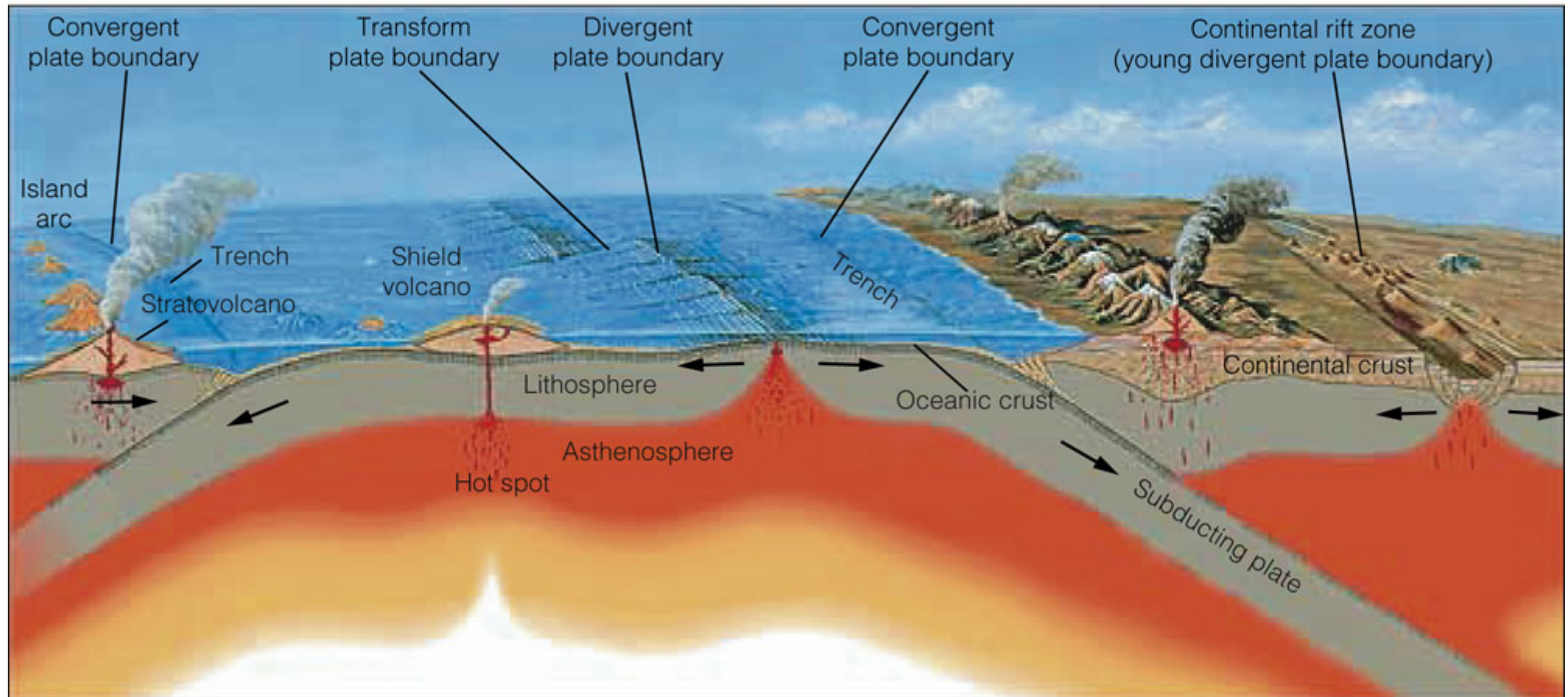
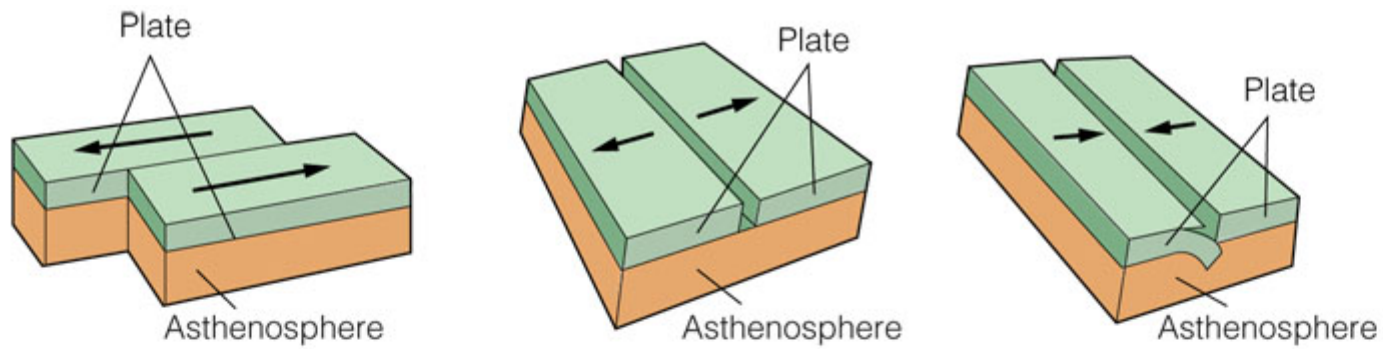


Seafloor age



Wadati-Benioff zone seismicity





The hardest part of the puzzle

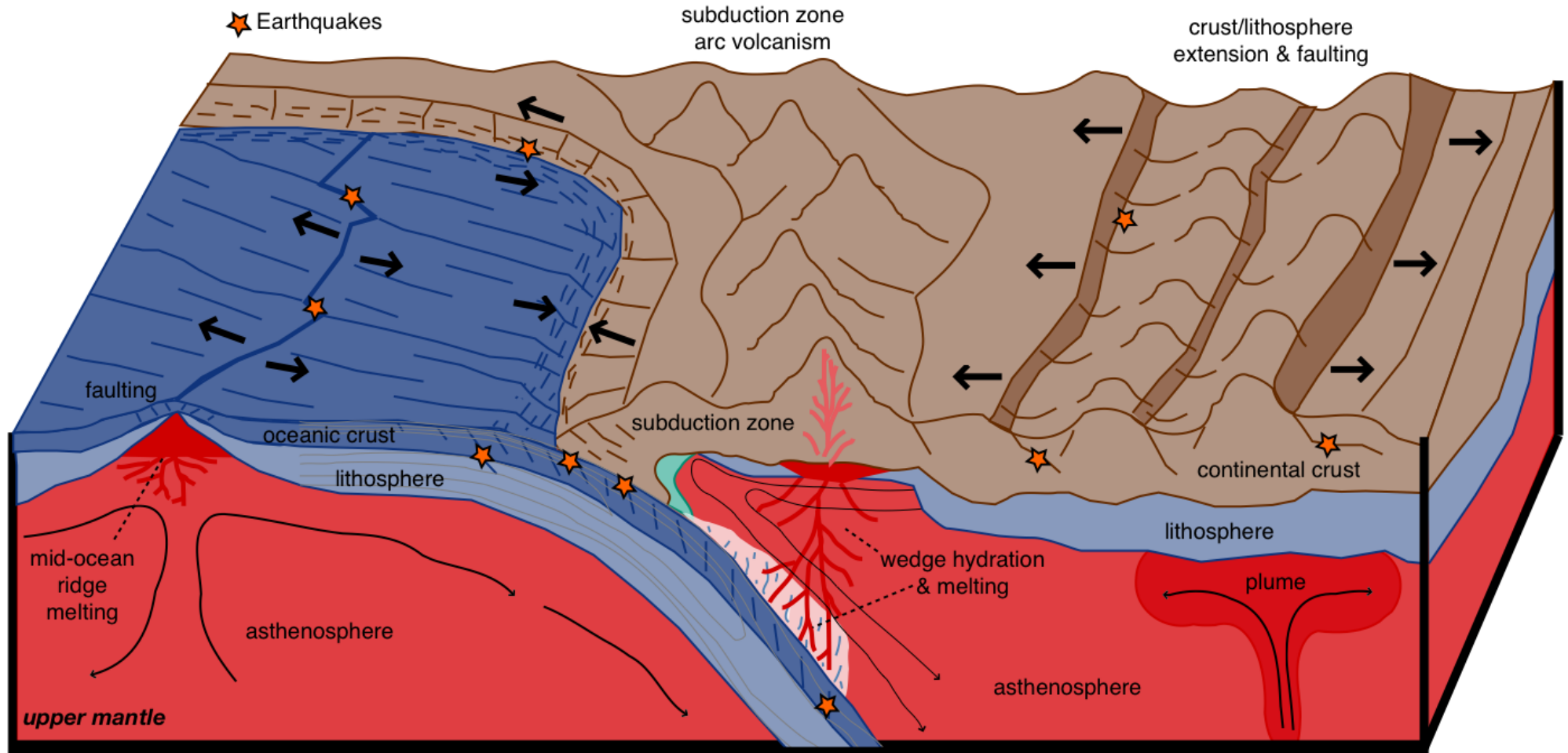
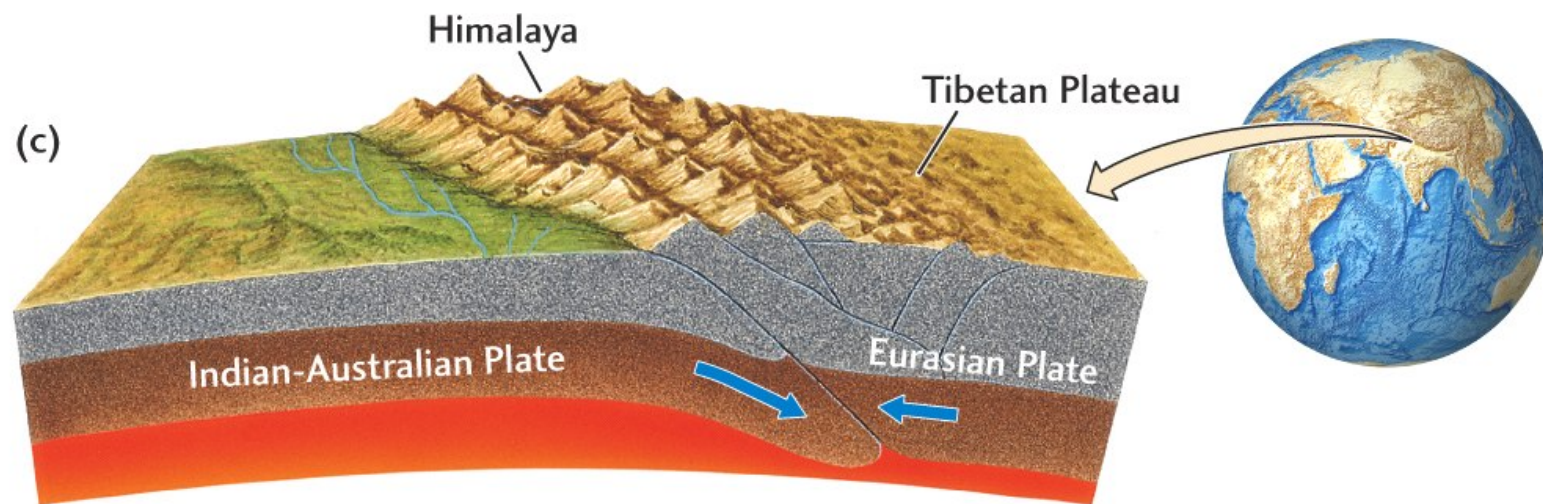
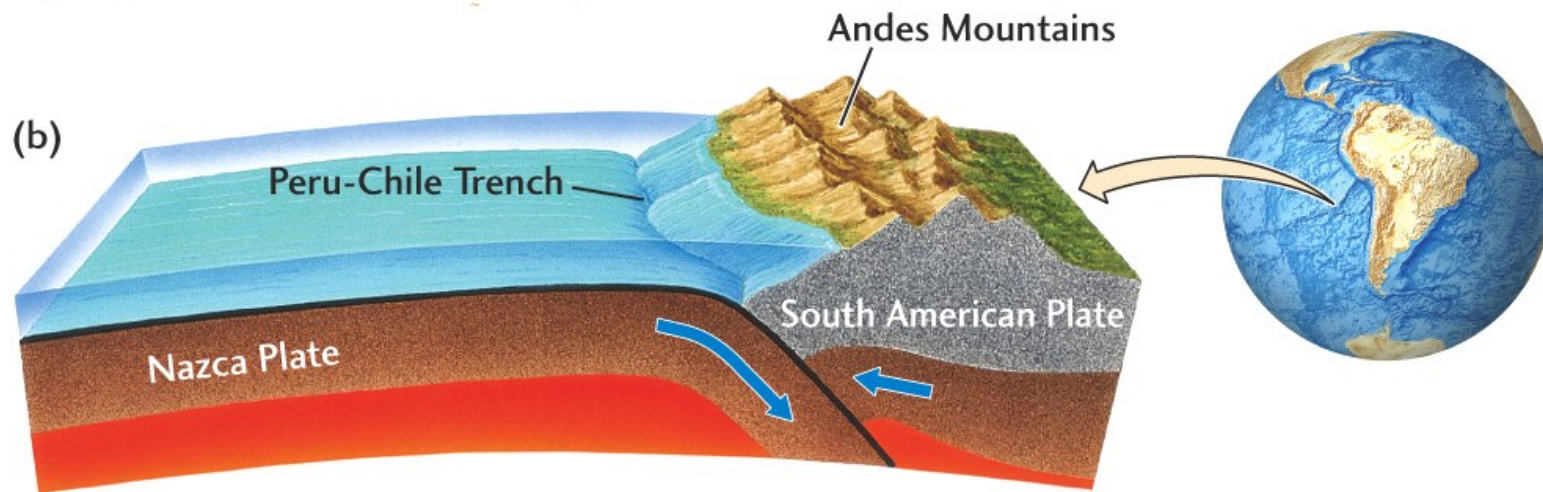
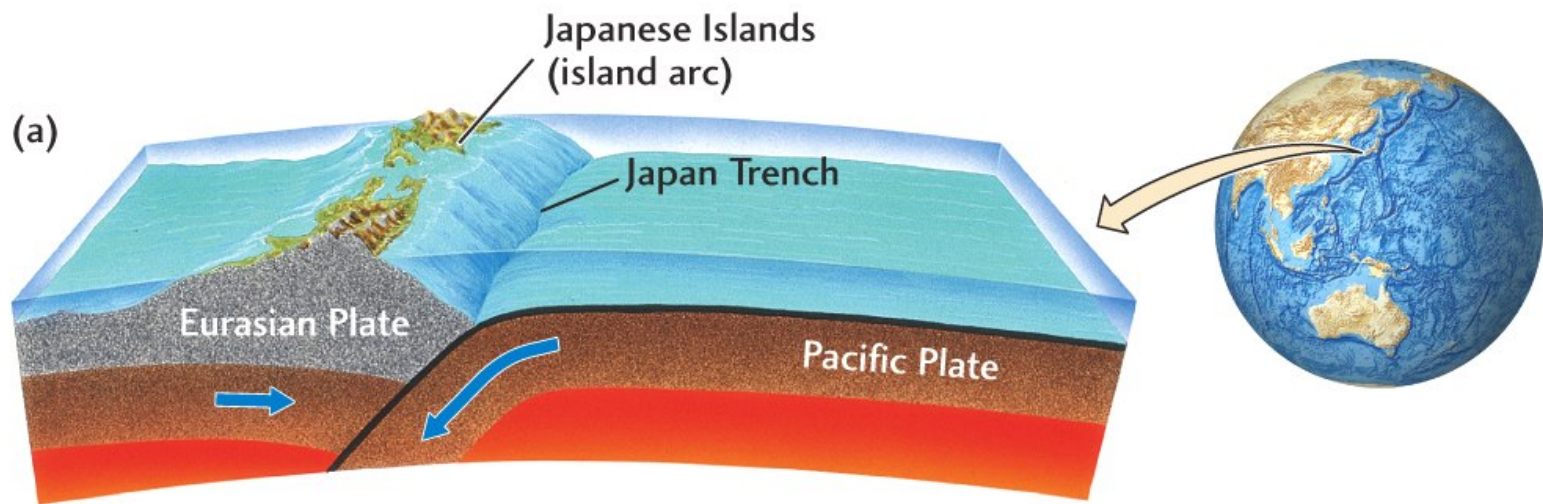
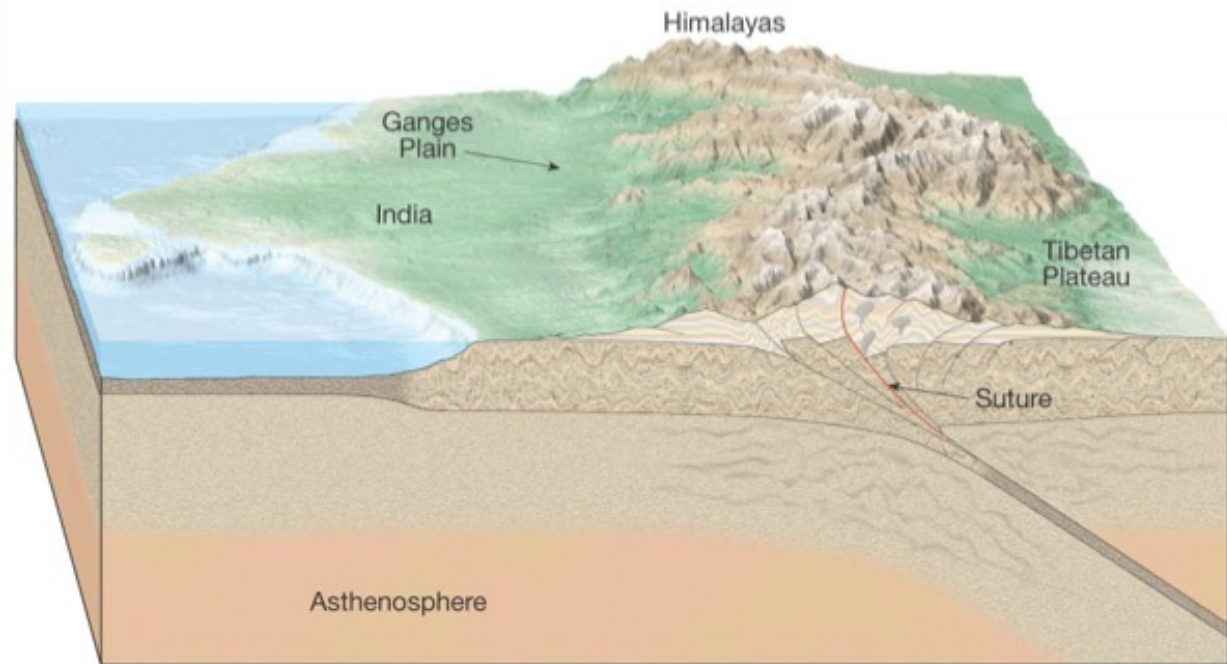
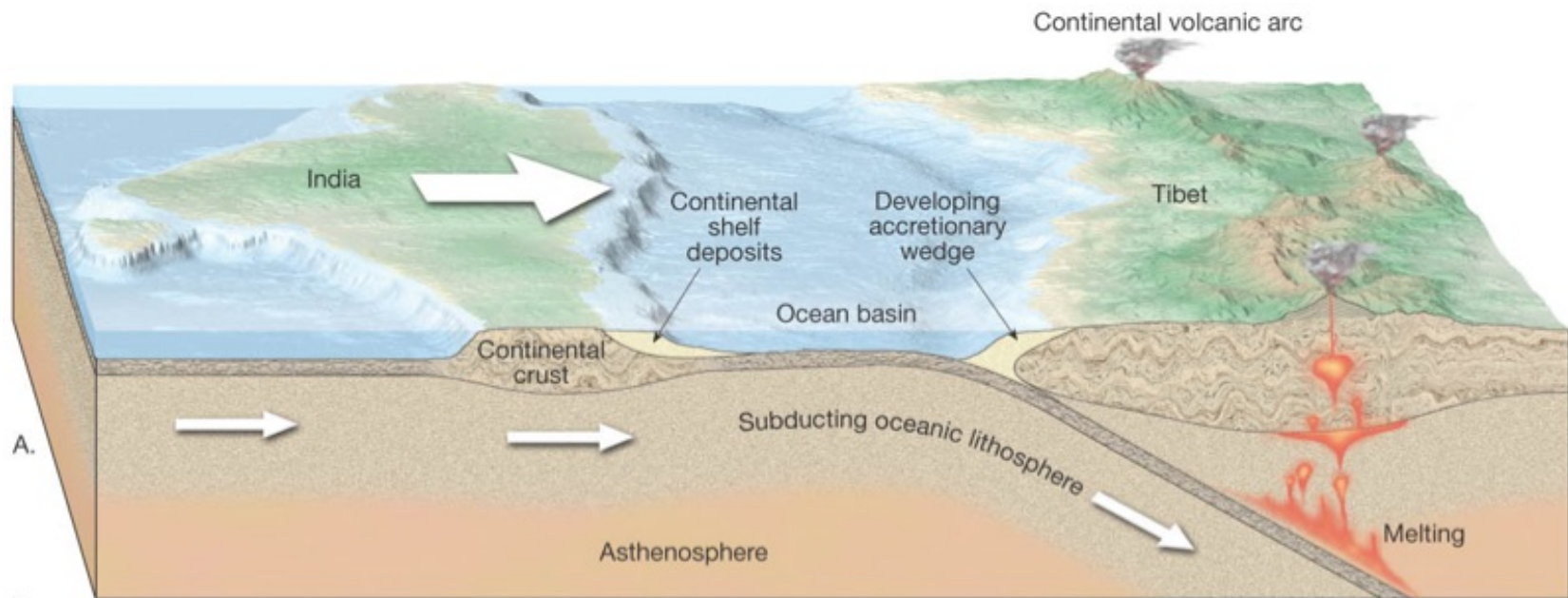
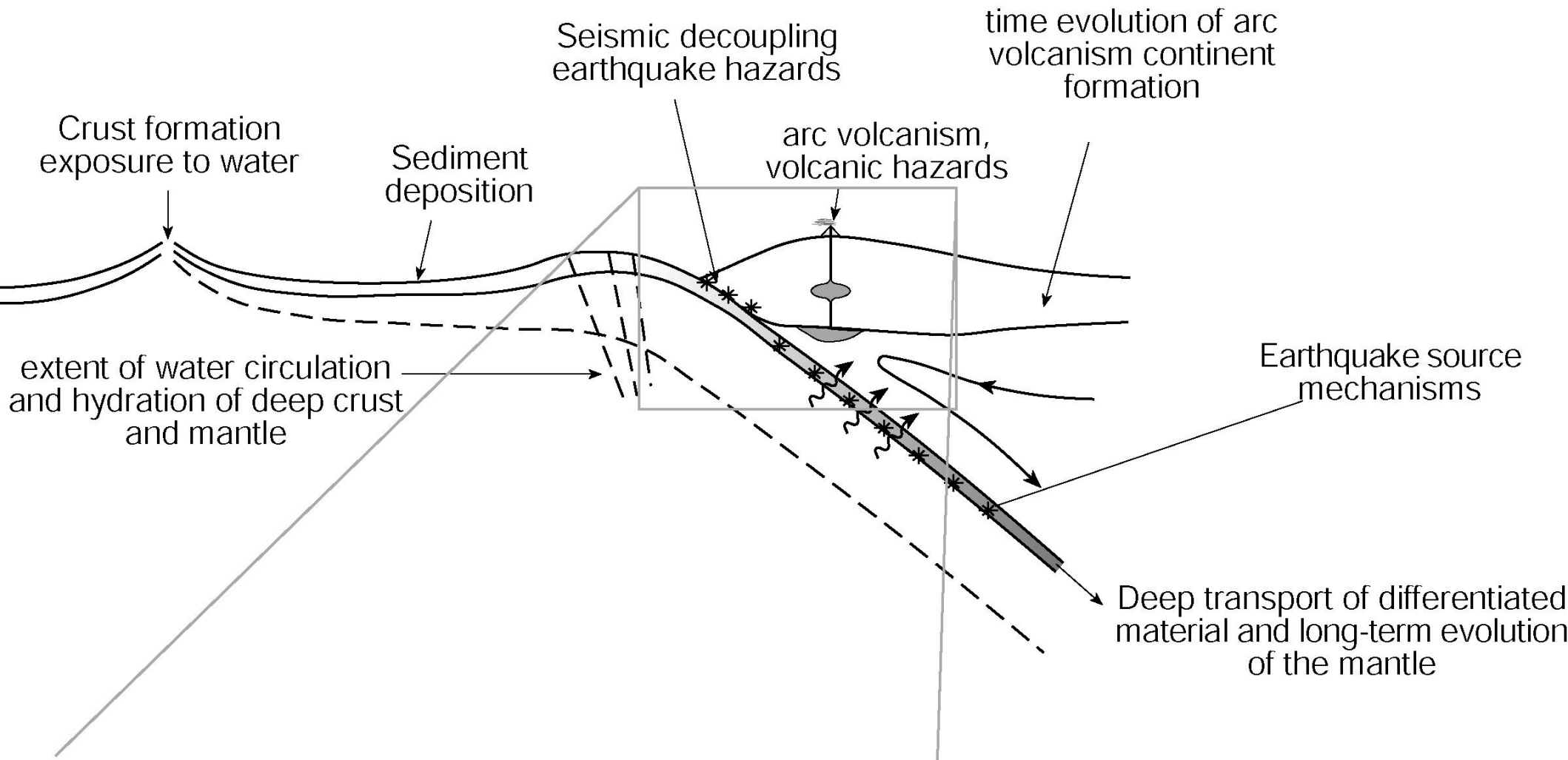


Figure courtesy of M. Billen

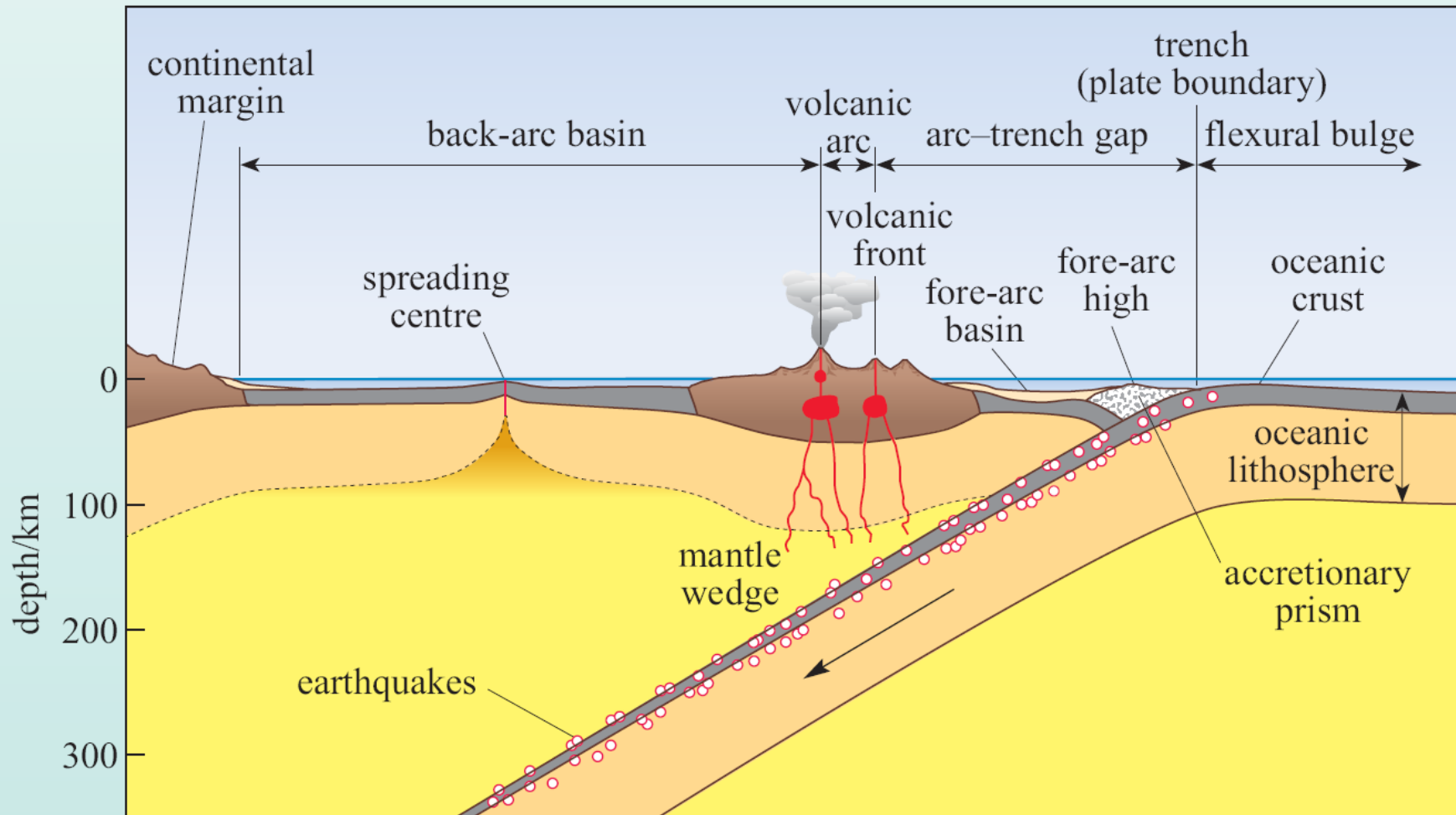


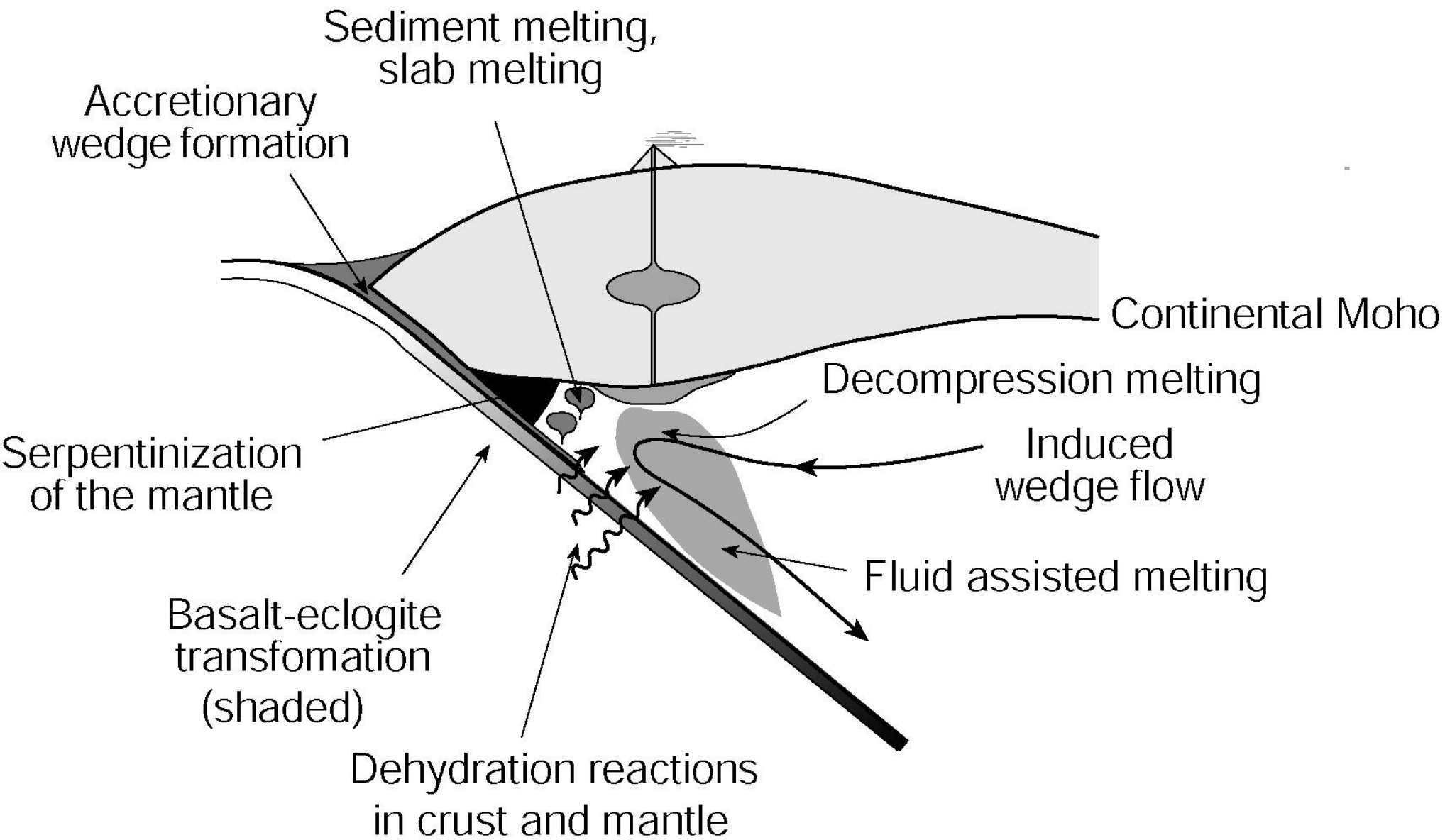


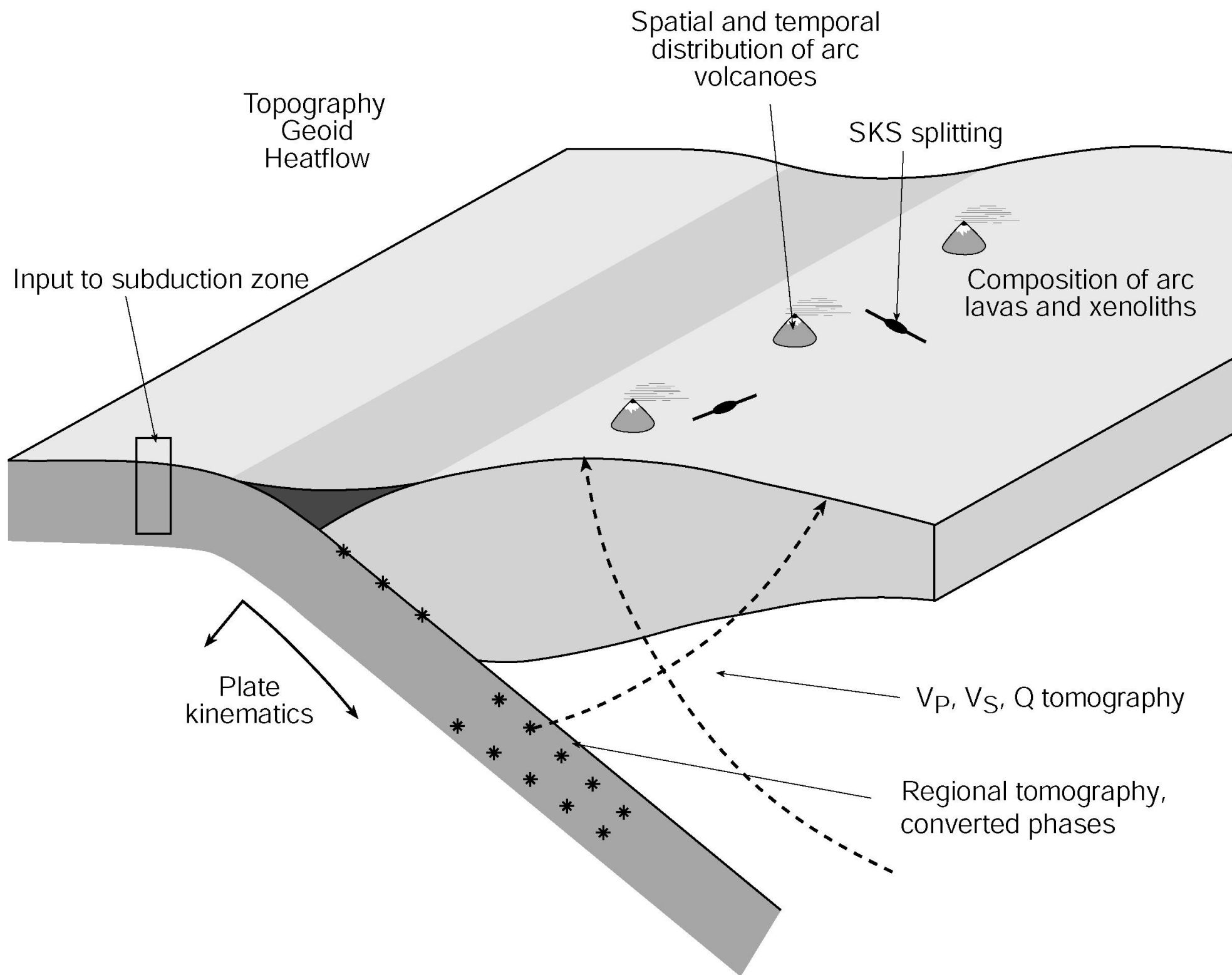
Subduction zones



Subduction zone terms







What drives the plates?

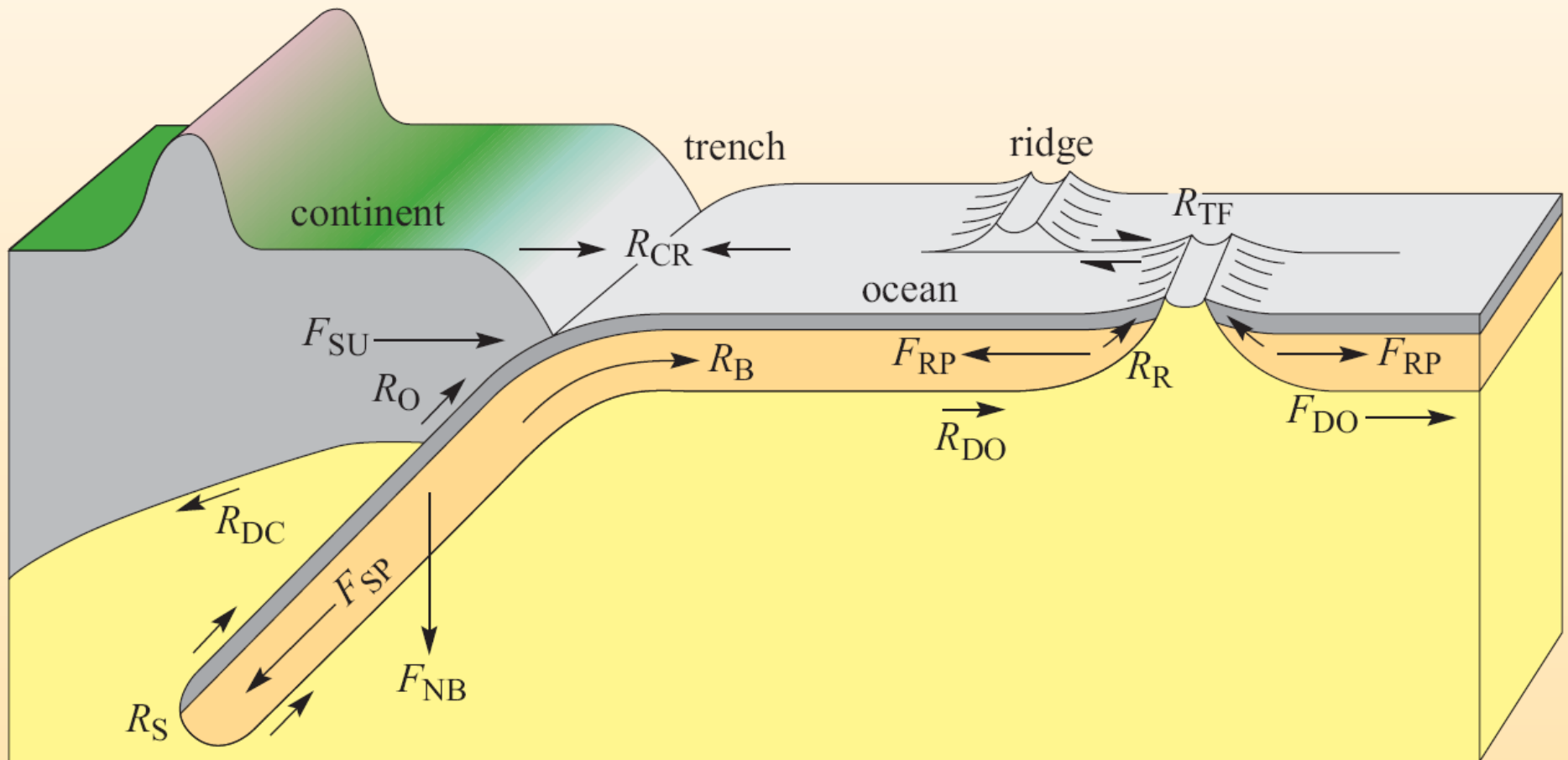
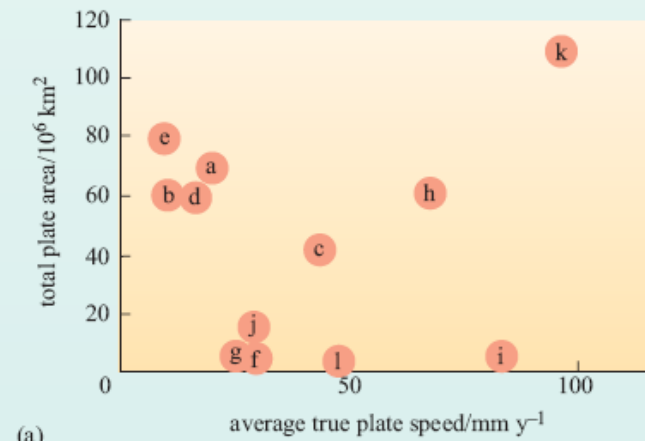
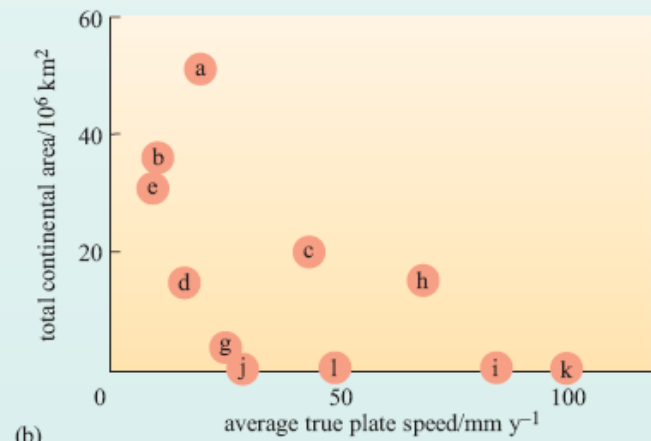


Plate driving forces

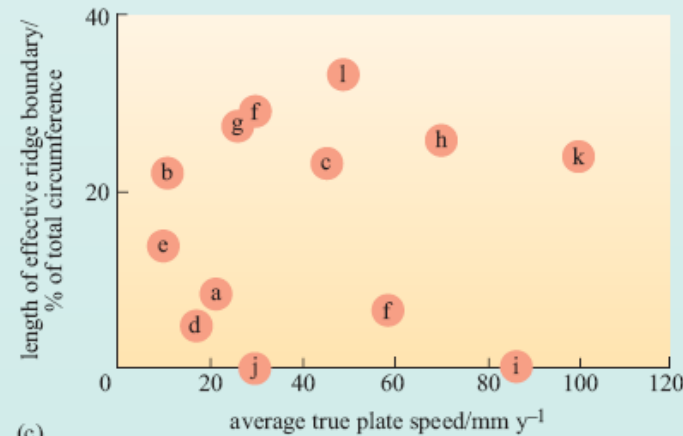
Forsyth & Uyeda (1975)



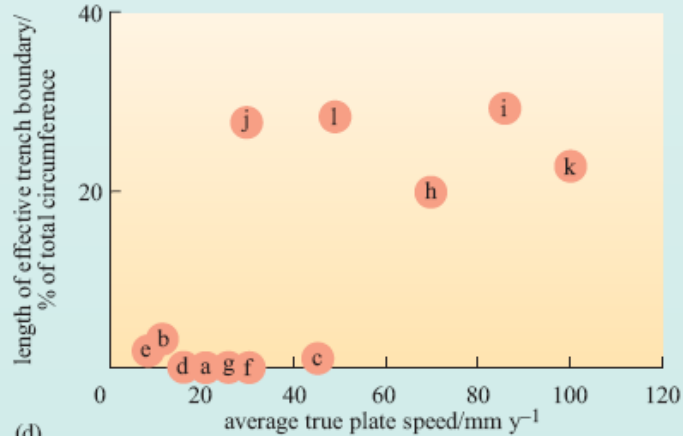
(a)



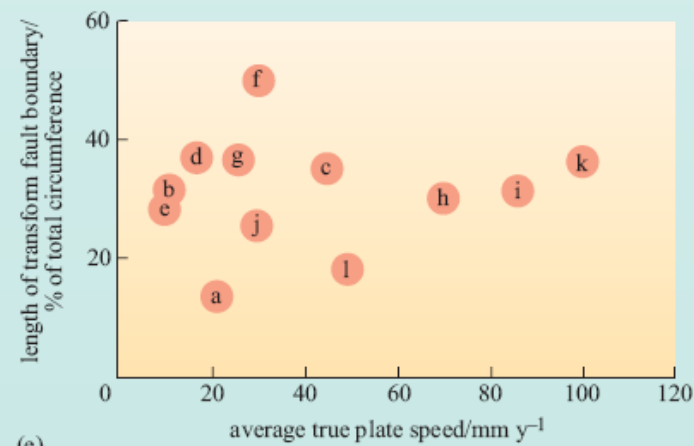
(b)



(c)



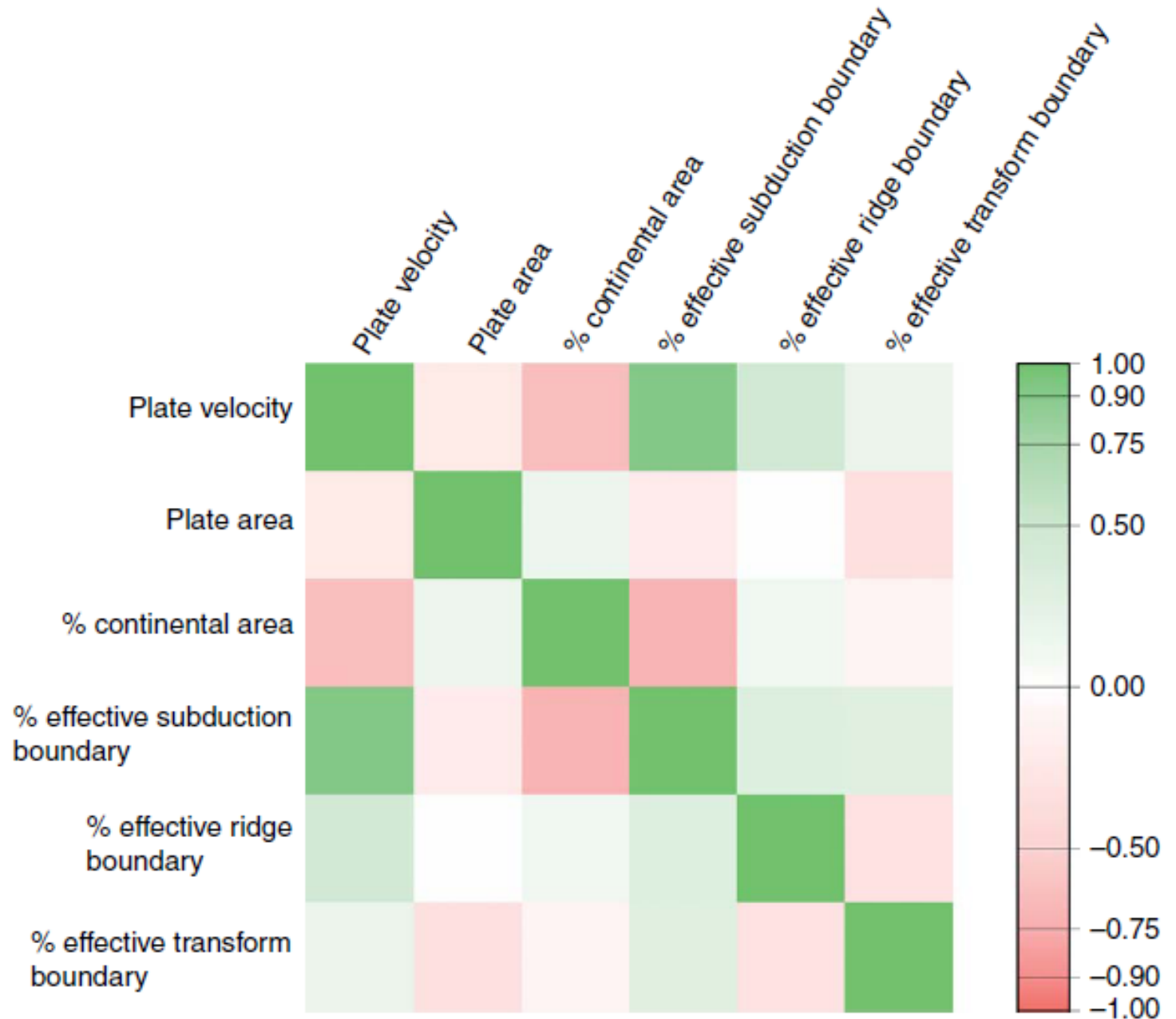
(d)



(e)

- (a) Eurasia
- (b) North America
- (c) South America
- (d) Antarctic
- (e) Africa
- (f) Caribbean
- (g) Arabian
- (h) Indo-Australia
- (i) Philippine
- (j) Nazca
- (k) Pacific
- (l) Cocos

Plate driving forces

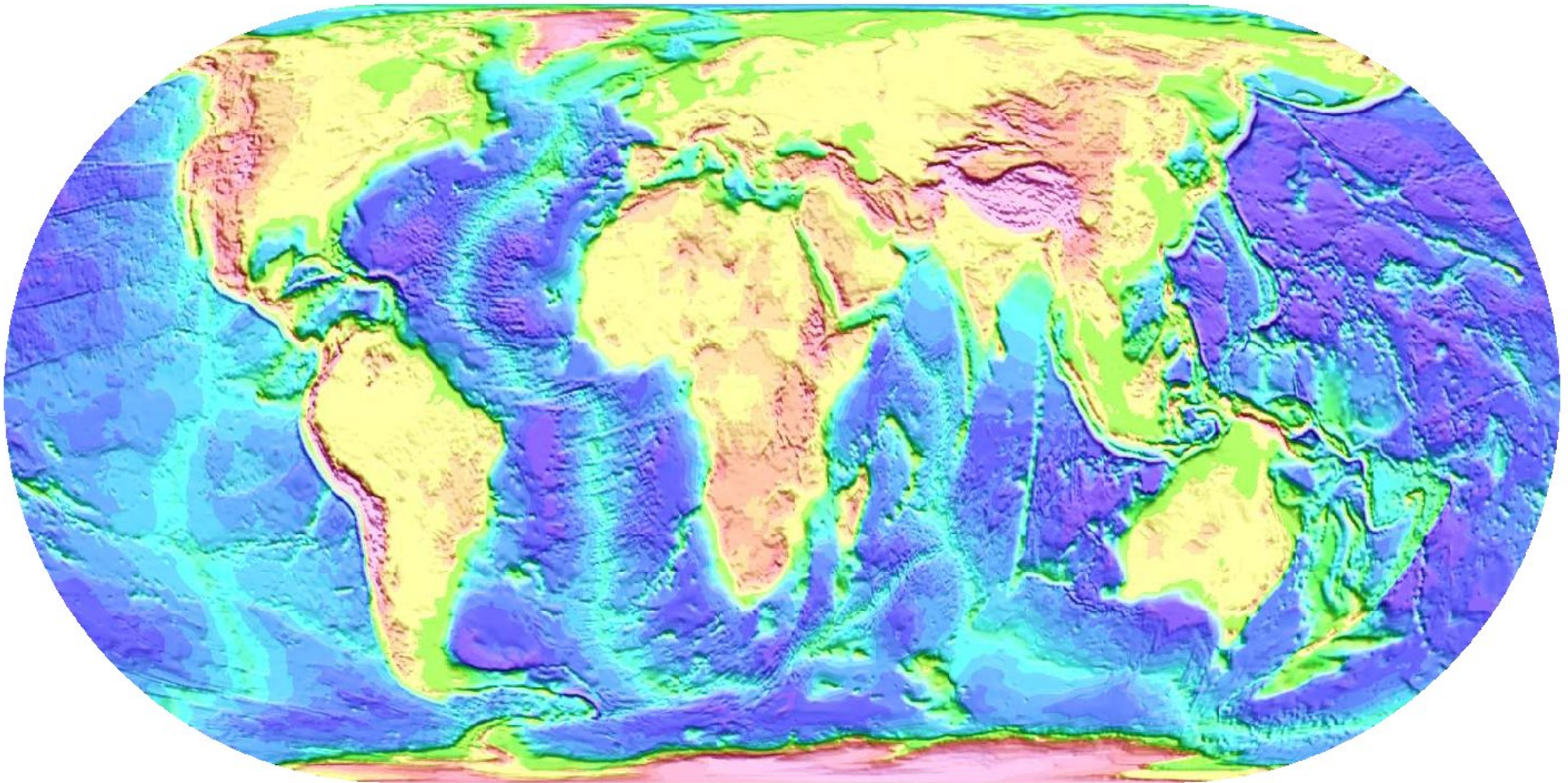


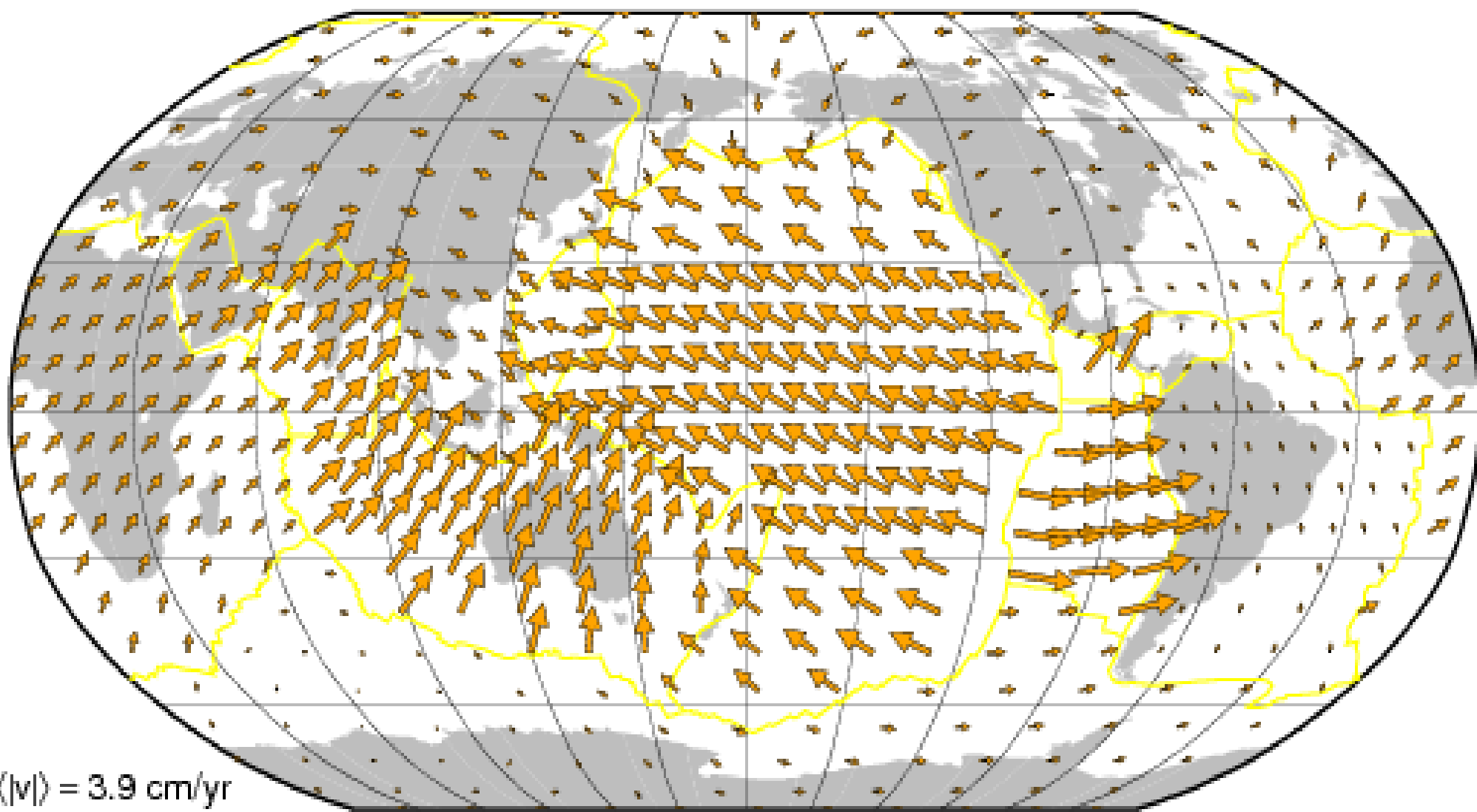
Forsyth & Uyeda (1975), plotted by King (2008)

Relevance of subduction

- Plate tectonics and Earth evolution (life) might be intrinsically linked
- Subducted slabs are responsible for ~70% of the plate driving forces
- Subduction zones host the largest earthquakes
- Subduction zone volcanism and sediment transport major players in volatile and carbon cycle

Half space cooling





Thermal structure of oceanic lithosphere

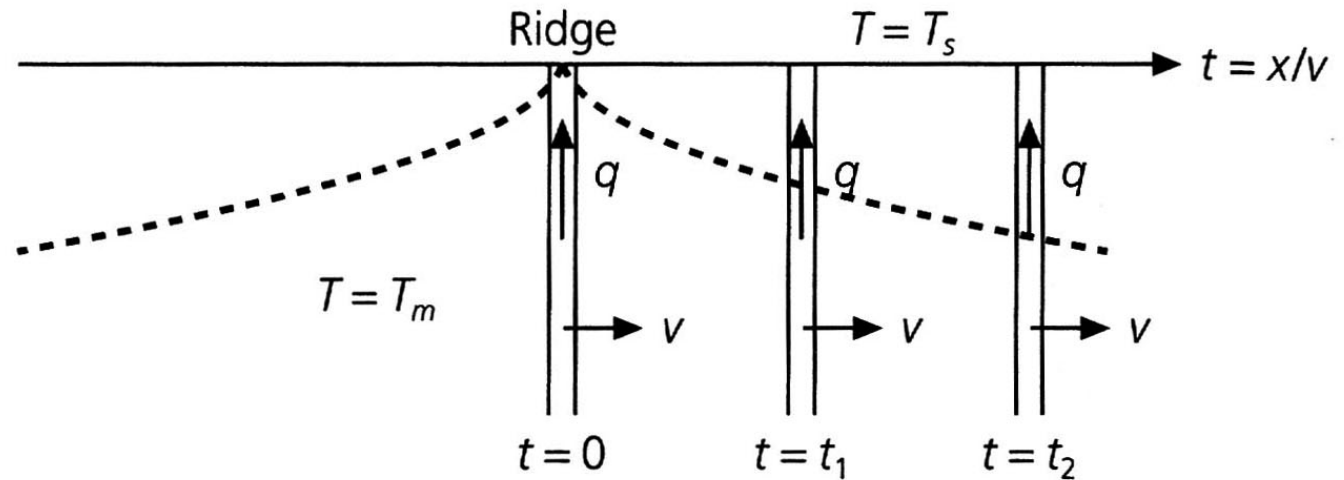
Cooling halfspace

$$\frac{\partial T}{\partial t} = \kappa \frac{\partial^2 T}{\partial z^2}$$

$$T = T_m \quad \text{at } x = 0$$

$$T = 0 \quad \text{at } z = 0$$

$$T = T_m \quad \text{at } z \rightarrow \infty$$



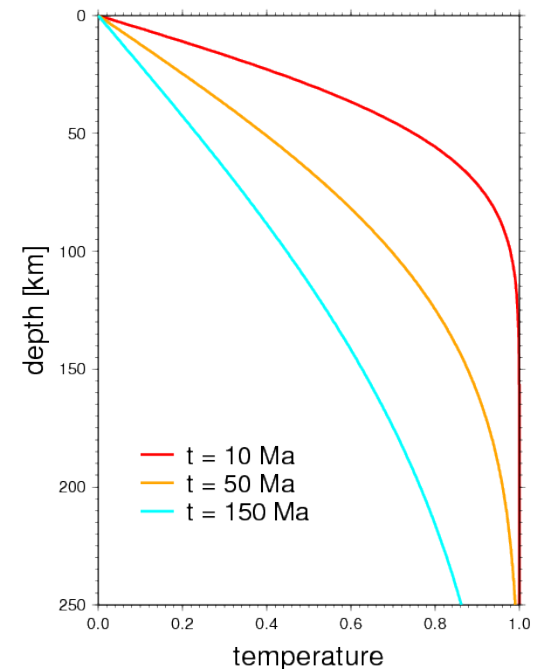
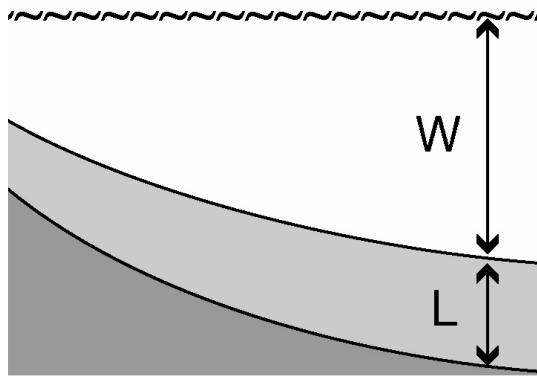
$$T(x, z) = T_m \operatorname{erf} \left(\frac{z}{2 \sqrt{\kappa t(x)}} \right) = T_m \operatorname{erf} \left(\frac{z}{2 \sqrt{\kappa x/v}} \right)$$

$$T(x, z) = T_m \operatorname{erf} \left(\frac{z}{2 \sqrt{\kappa x / v}} \right)$$

Depth of an isotherm $z_{iso} = \sqrt{\kappa x / v}$

Surface heat flow $q_s = -k \left(\frac{\partial T}{\partial z} \right)_{z=0} = \frac{k T_m}{\sqrt{\pi \kappa t}} = \frac{k T_m}{\sqrt{\pi \kappa x / v}}$

Bathymetry $w(x) = \frac{2 \rho_m \alpha_v T_m}{(\rho_m - \rho_w)} \sqrt{\frac{\kappa x}{\pi v}}$



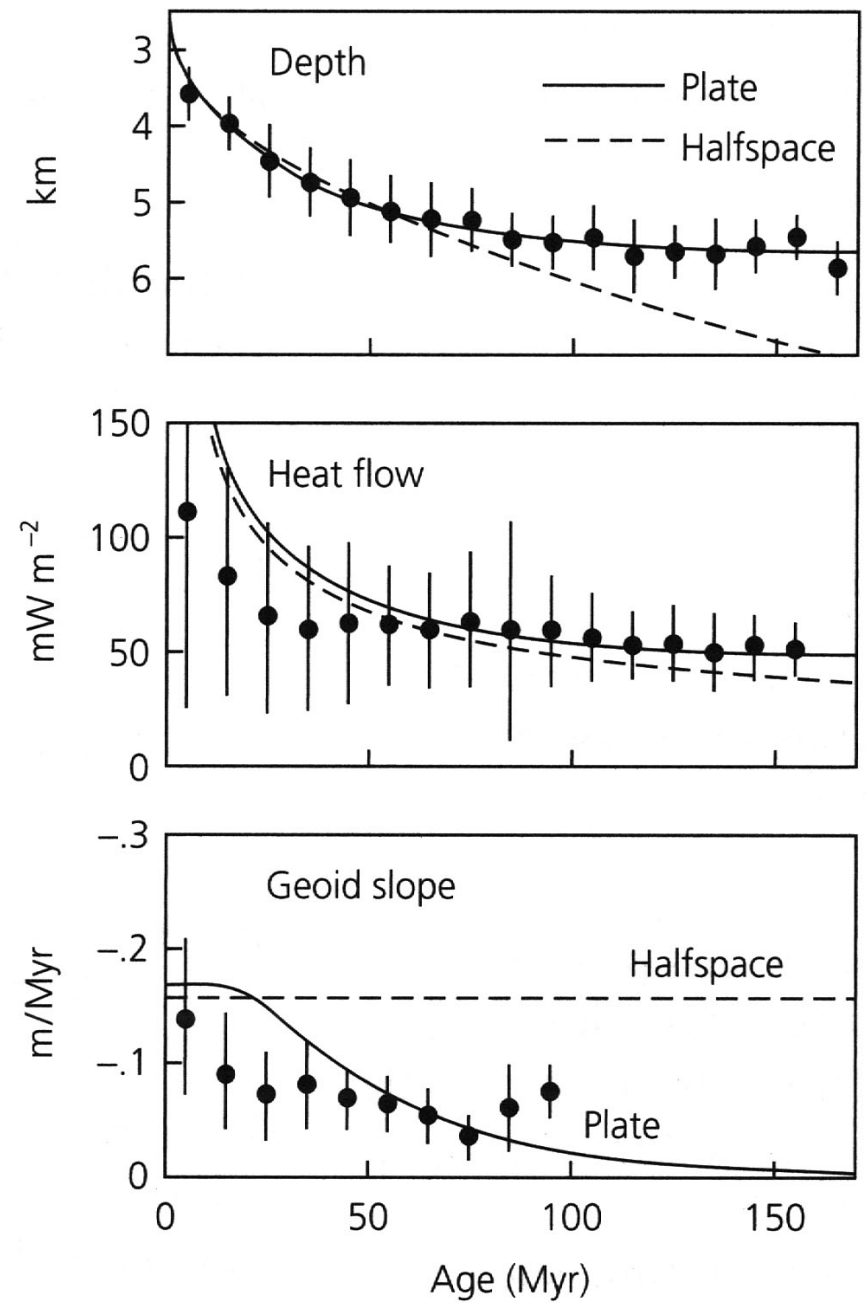
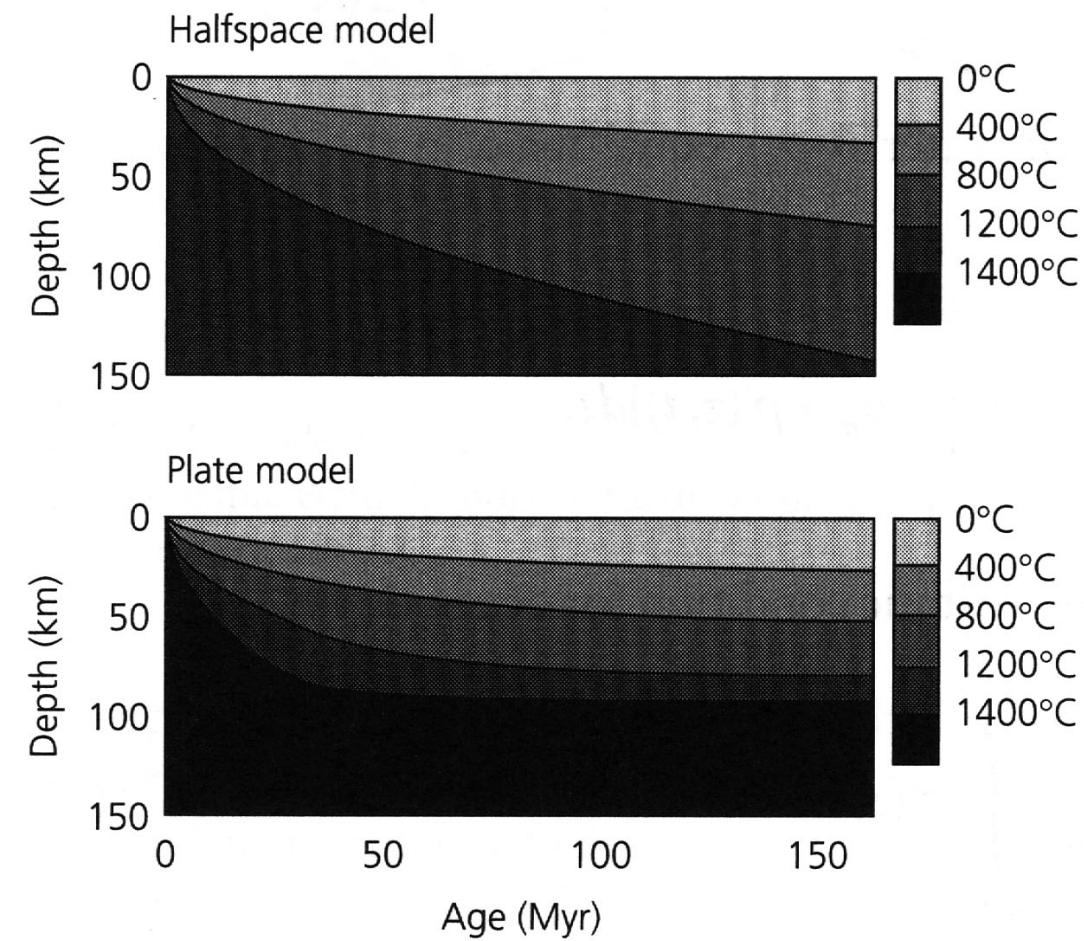
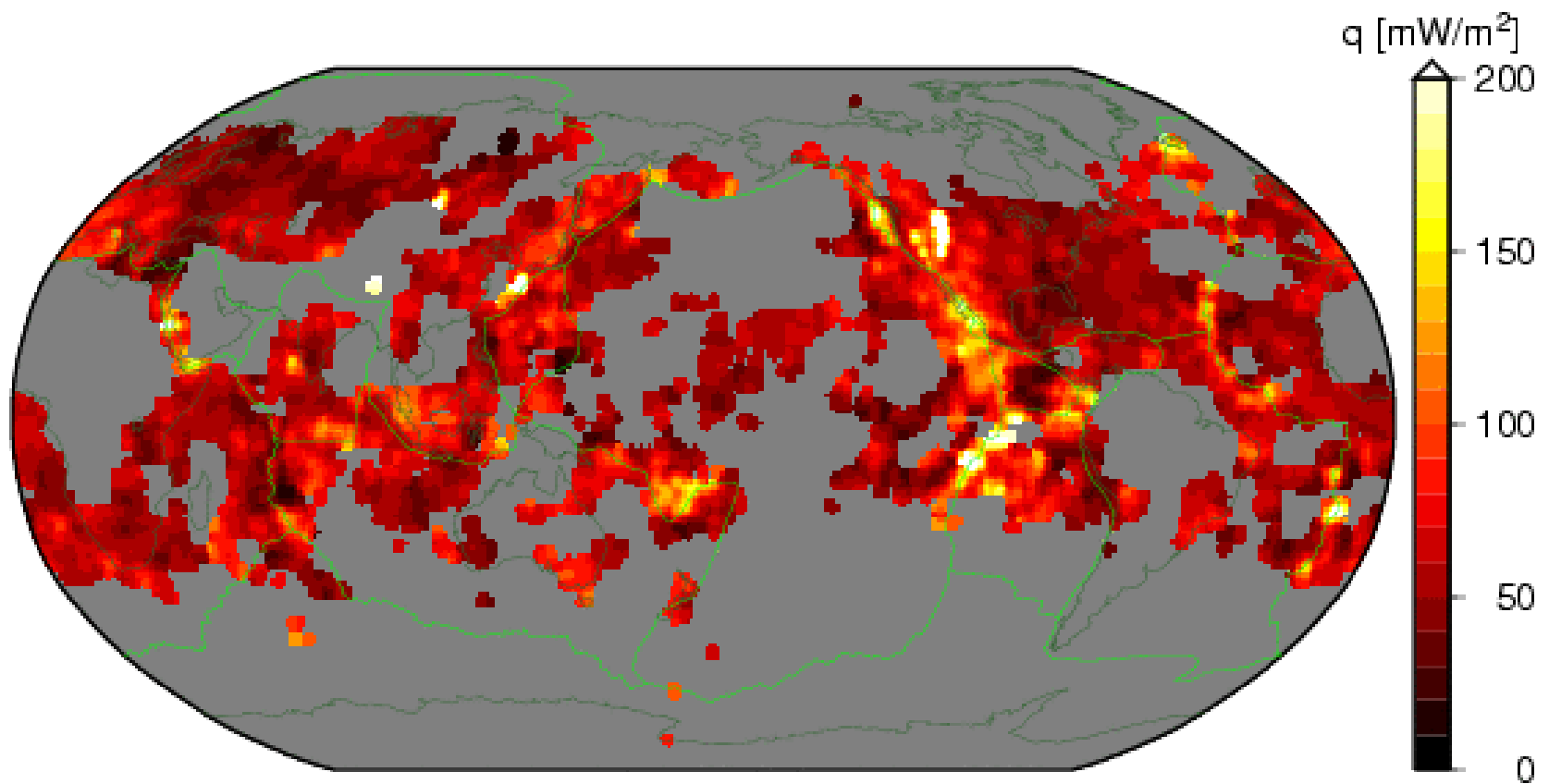


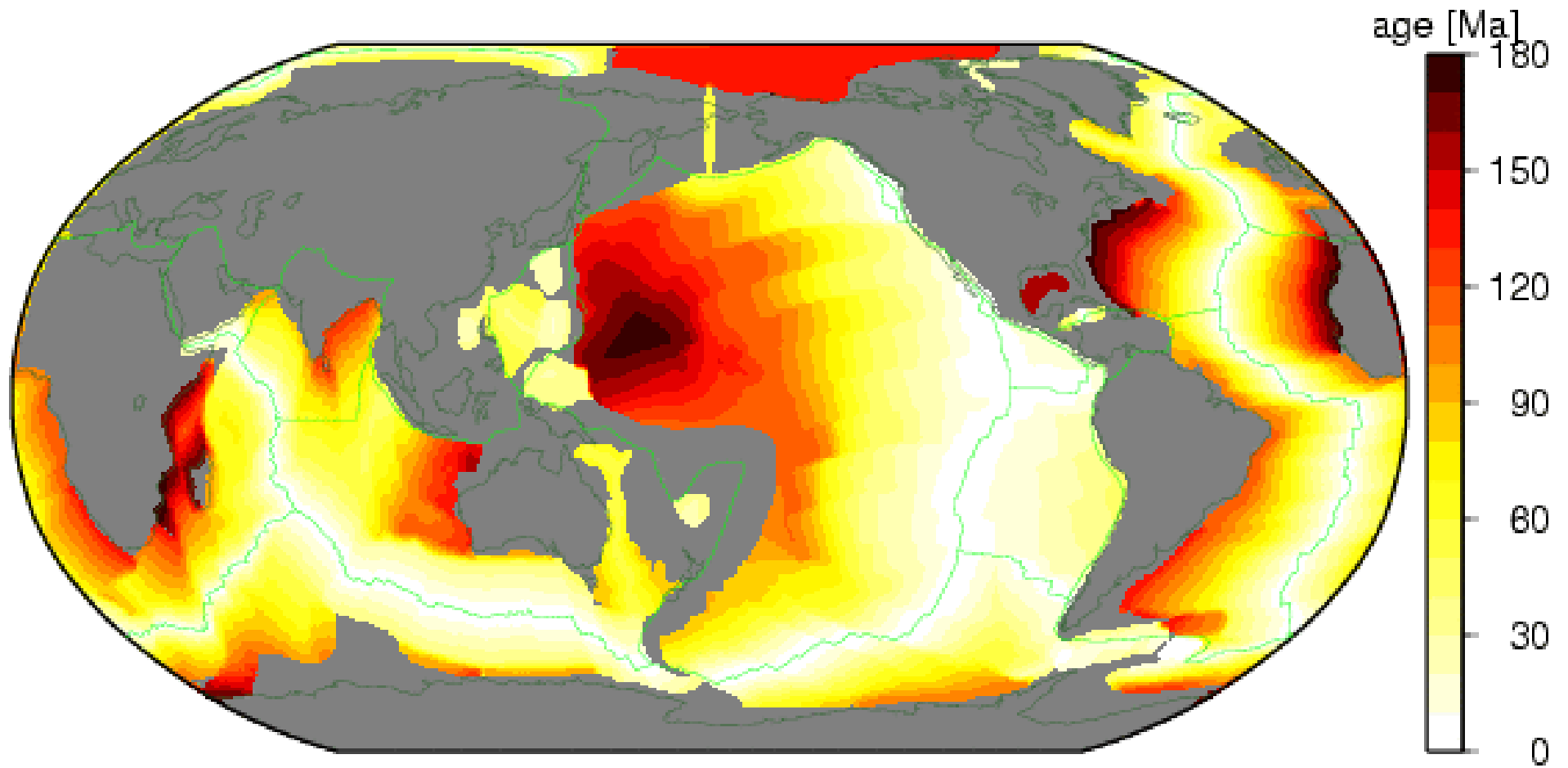
Fig. evol
Left
litho
in a
~70
a 95
plate
temp
Right
pred
litho
(but
of a
thos
(Ric
Left
Ame

Richardson et al. (1995)

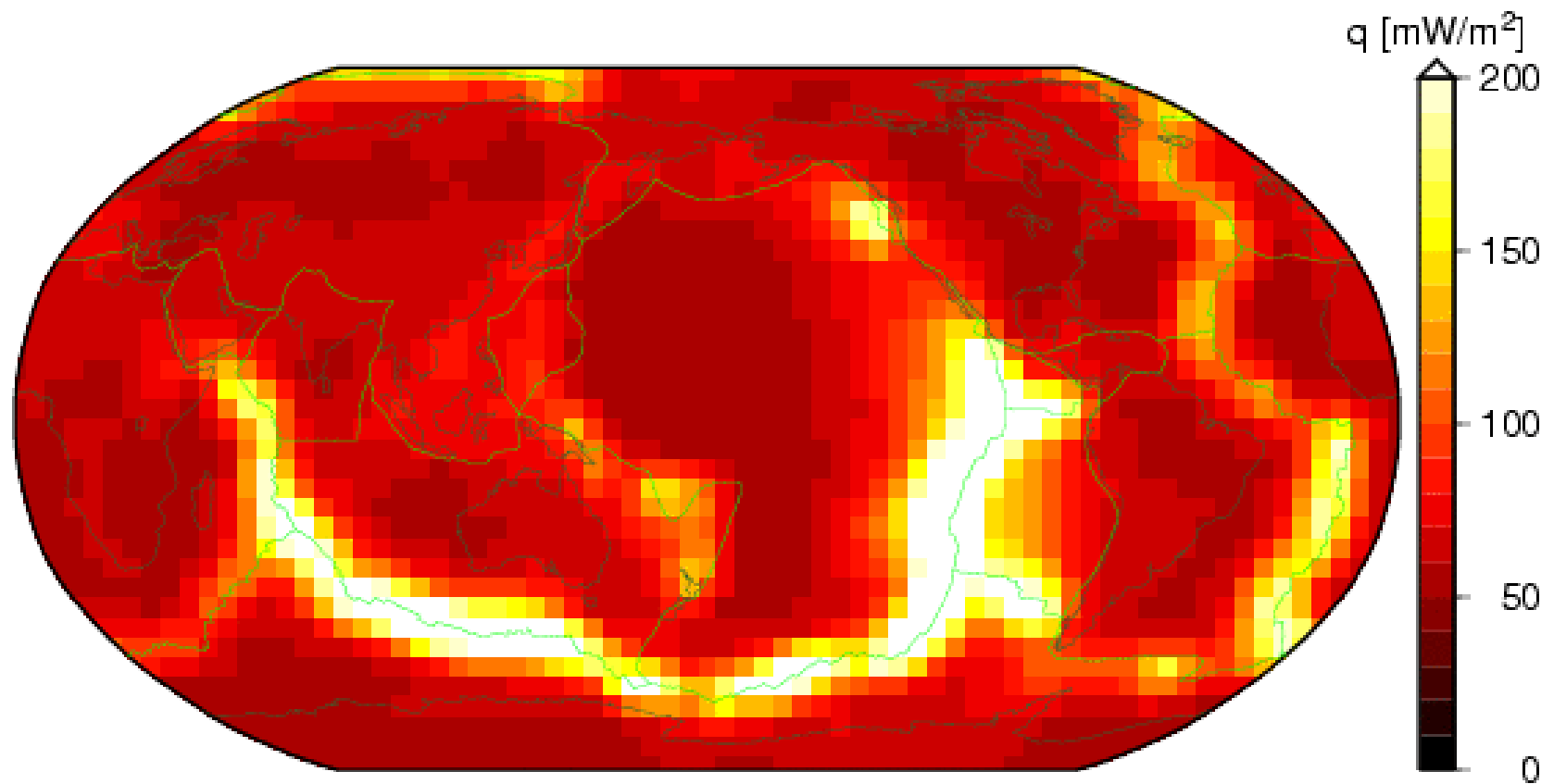
Heat flow data



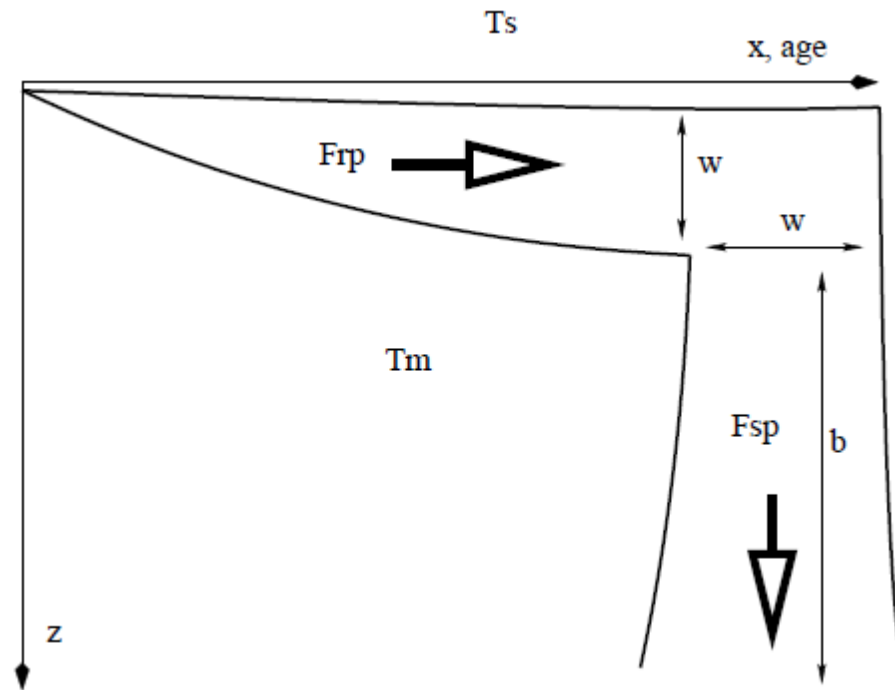
Seafloor age



Global heat flow

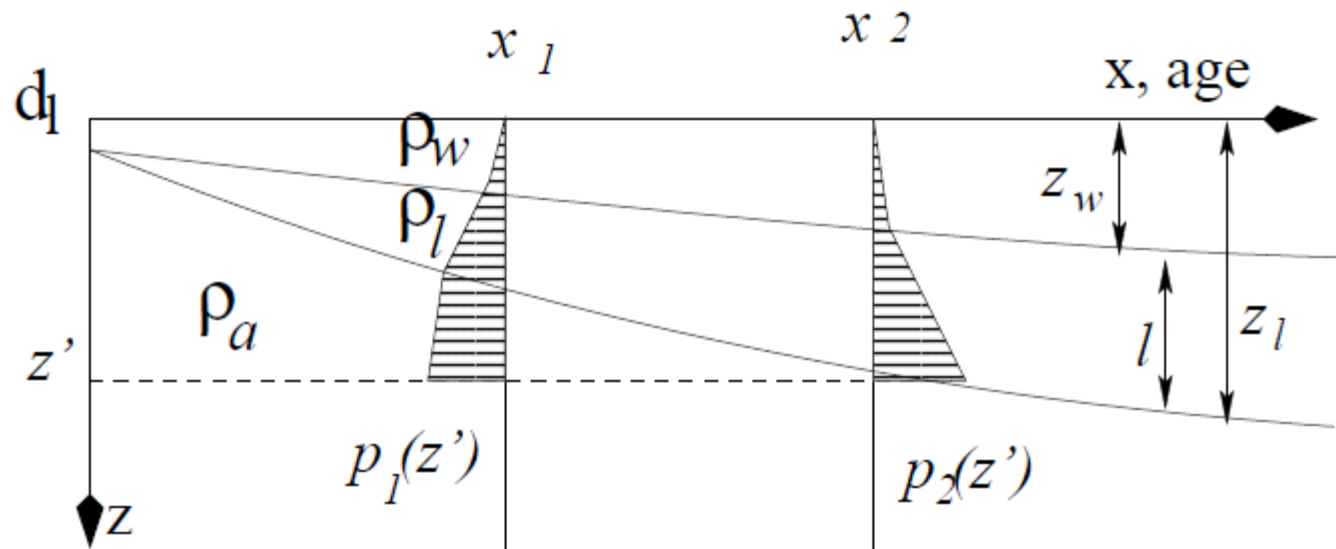


Force balance based on half-space cooling slab pull



$$F_b = -2\rho_0 g \alpha b (T_m - T_s) \sqrt{\frac{\kappa t_s}{\pi}}$$

Lithospheric thickening



$$l = a(z_w - d_l) = a\Delta z_w \quad a = \frac{\rho_w - \rho_a}{\rho_a - \rho_l} \quad z_w(x) = c\sqrt{\frac{x}{u_0}}$$

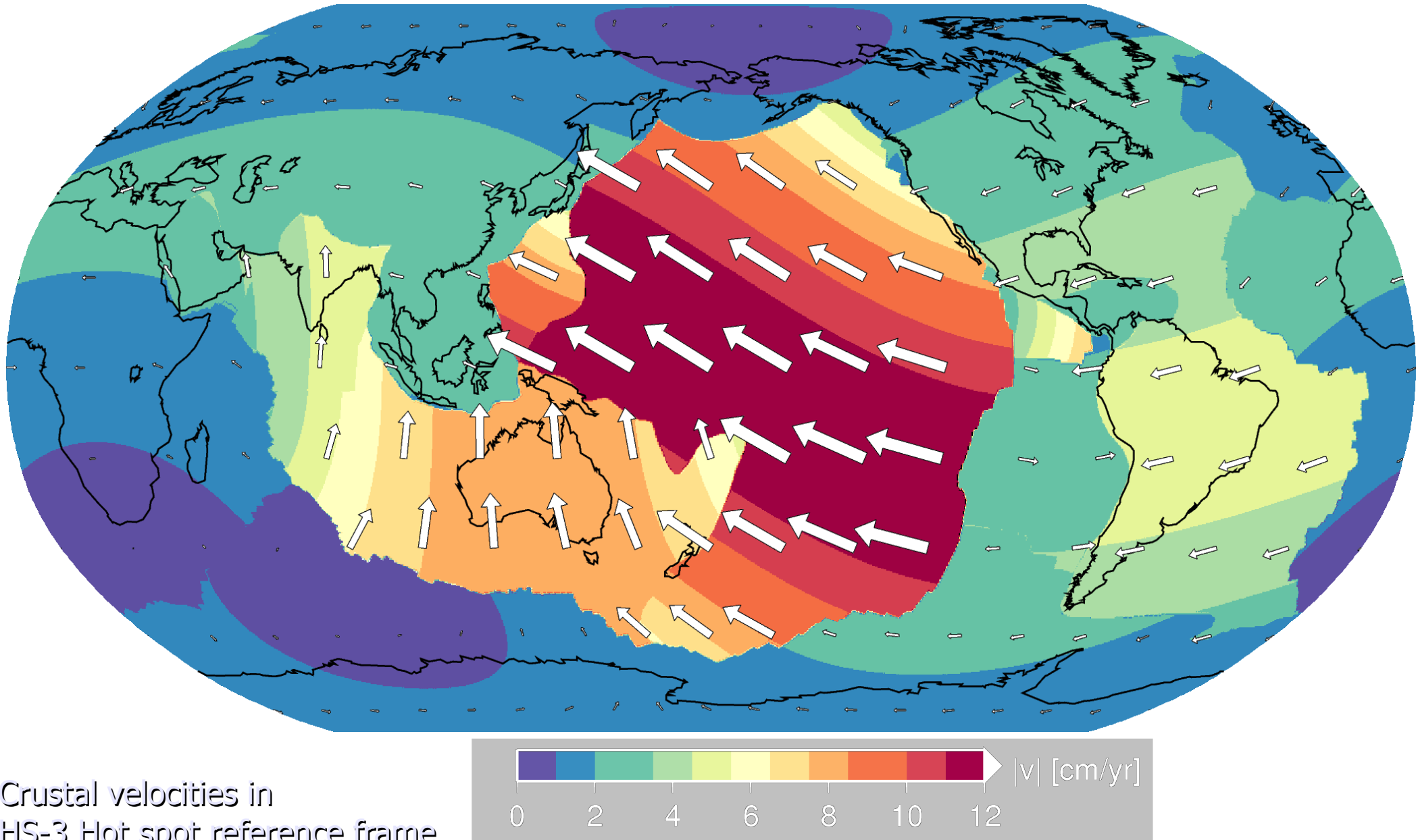
$$F = \frac{ac^2}{2u_0} g \Delta \rho \Delta x$$

Force estimates from half-space cooling

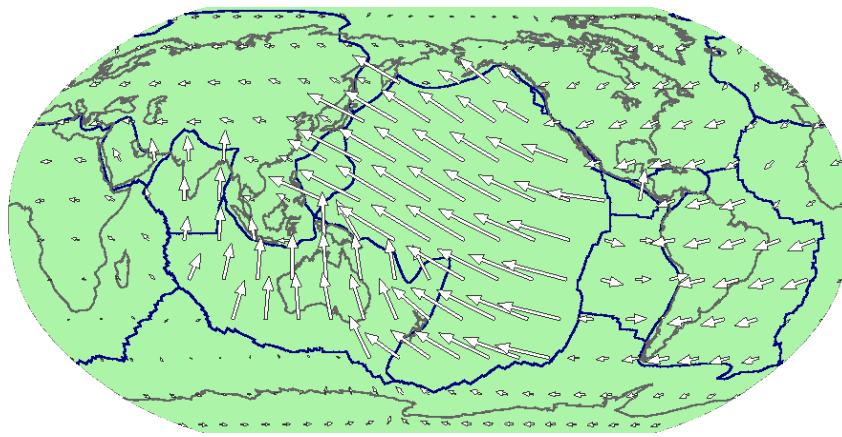
- Ridge push (lithospheric thickening) $\sim 10^{12}$ N/m
- Slab pull $\sim 10^{13}$ N/m

Geophysical constraints on slab dynamics

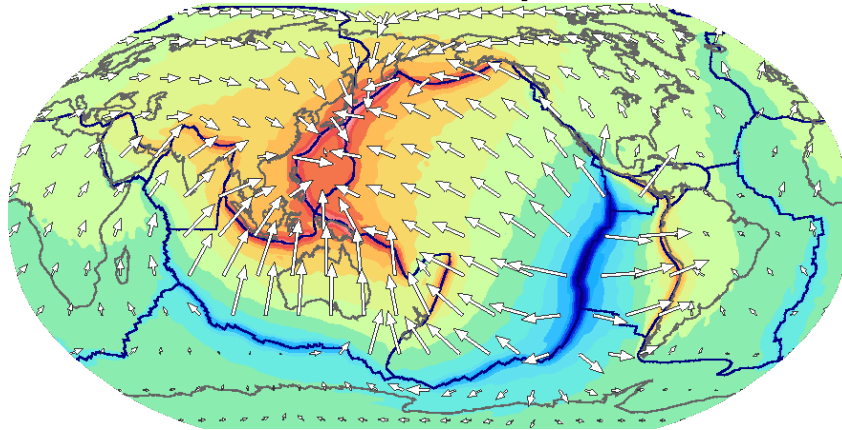
Crustal velocities in plate model



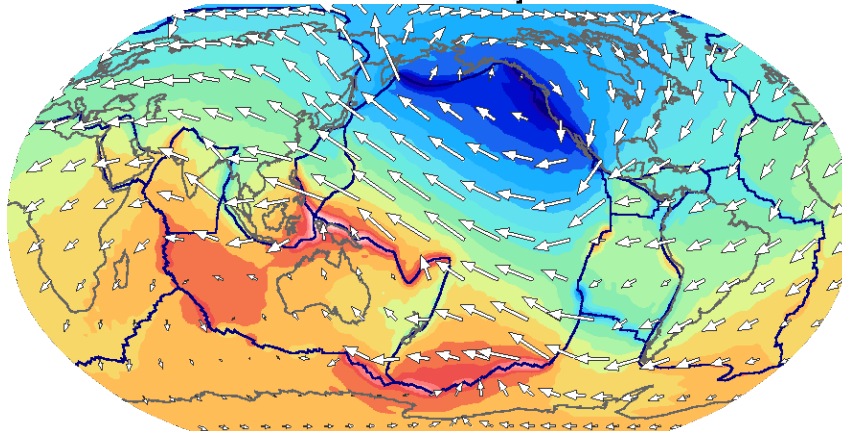
Kinematic characteristics of present-day plate tectonics



Poloidal component



Toroidal component



Sources and sinks

Strike slip motion,
Transform faults

Spin for $l = 1$
(net rotation, NR)

Time variability in plate motions (and heat transport)

Plate reconstruction A:

Mueller et al. (2008)

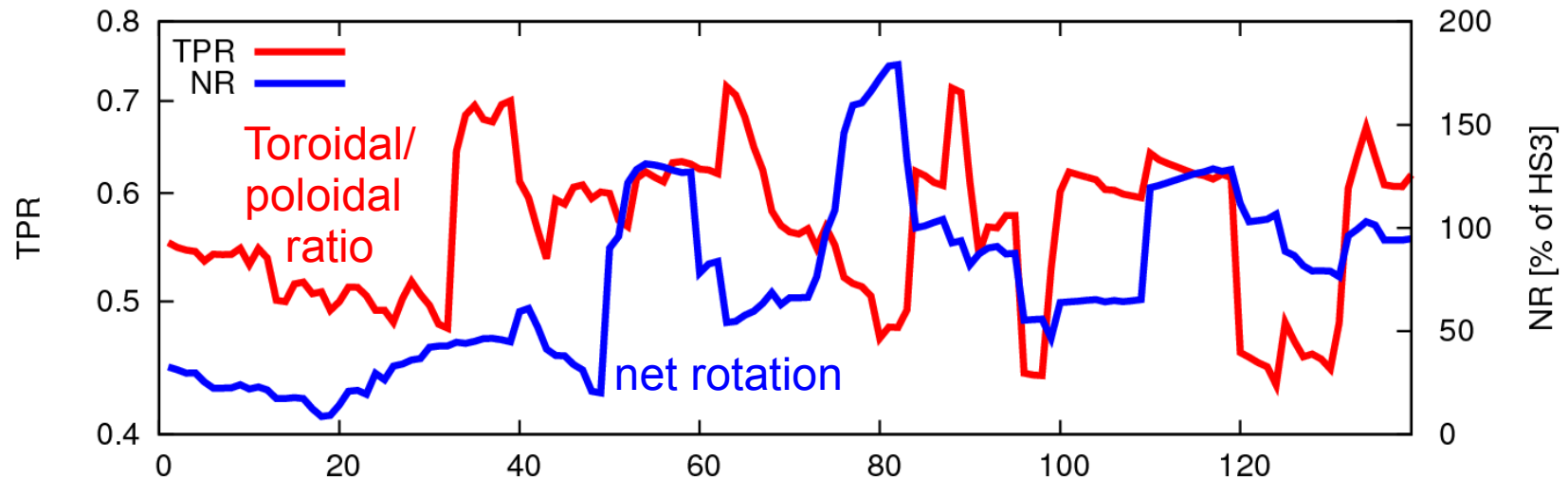
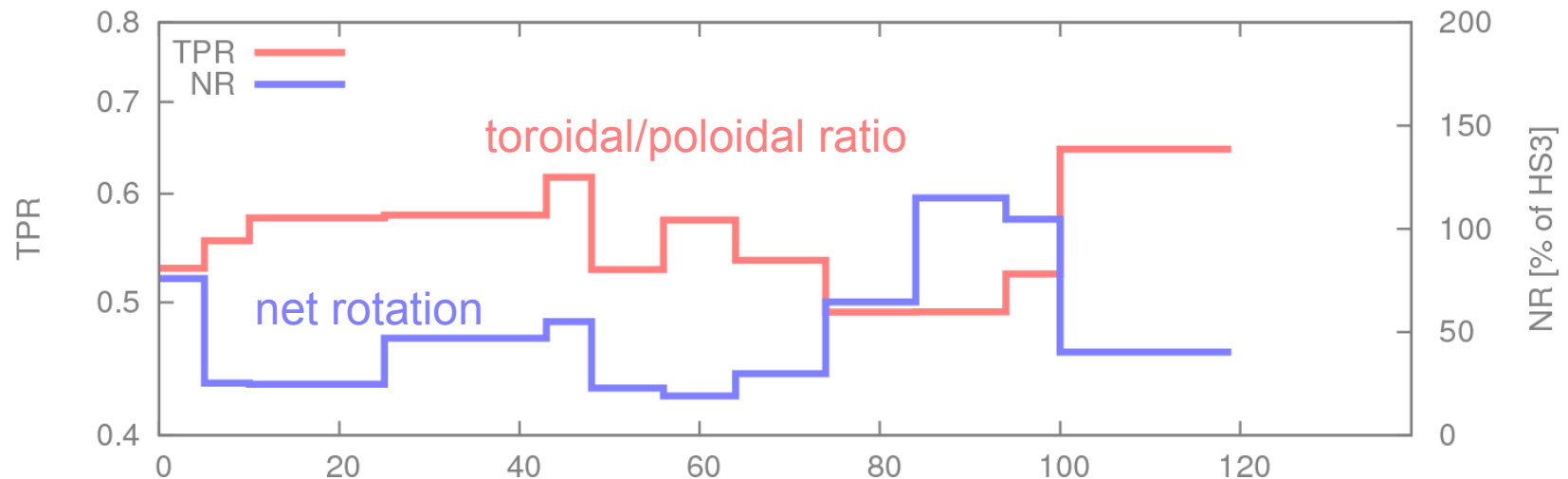
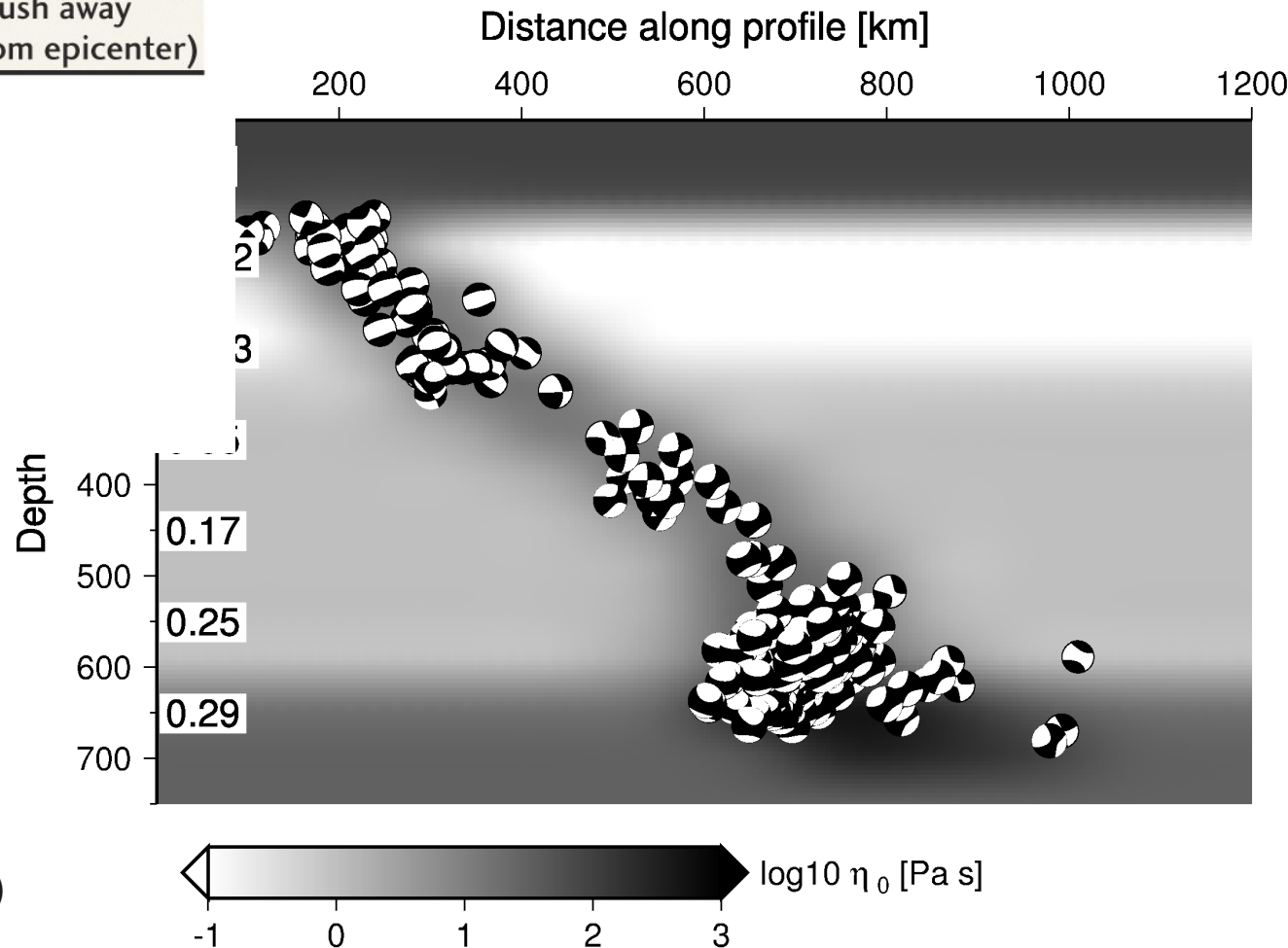
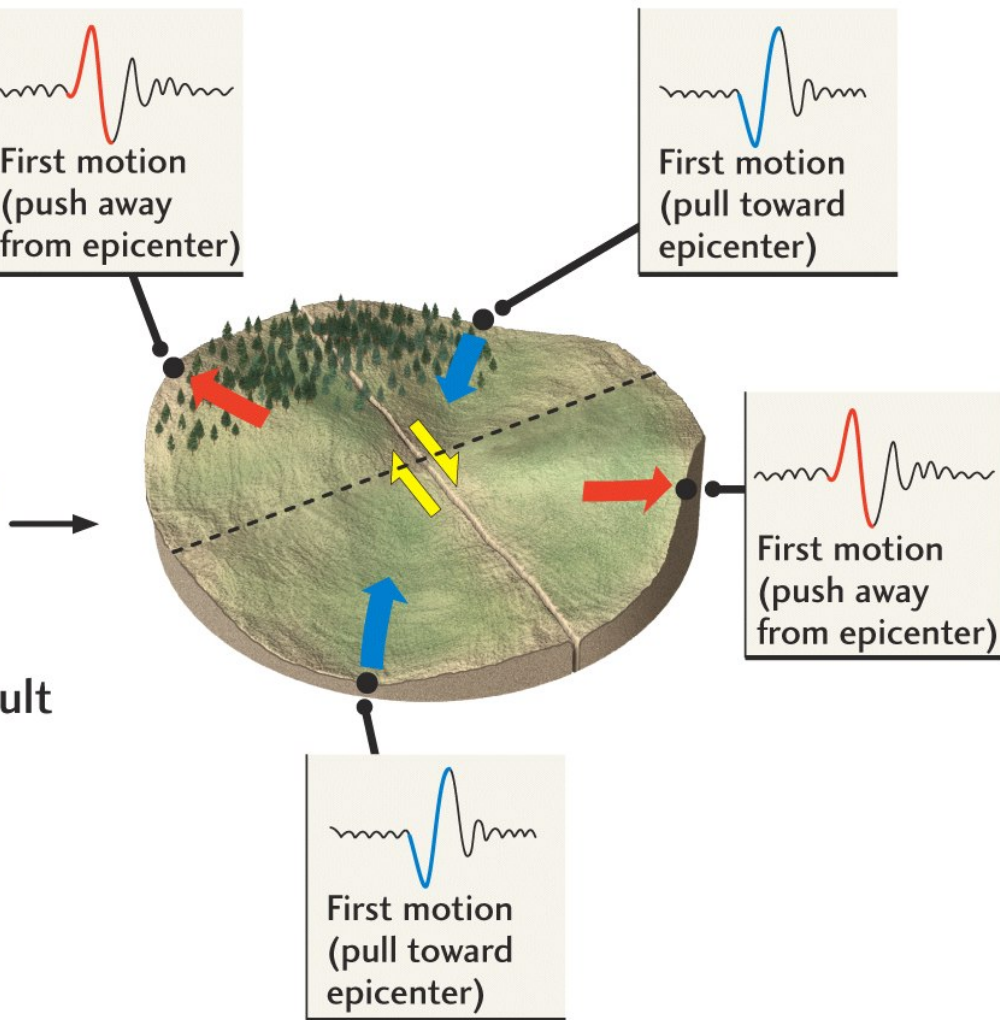


Plate reconstruction B:

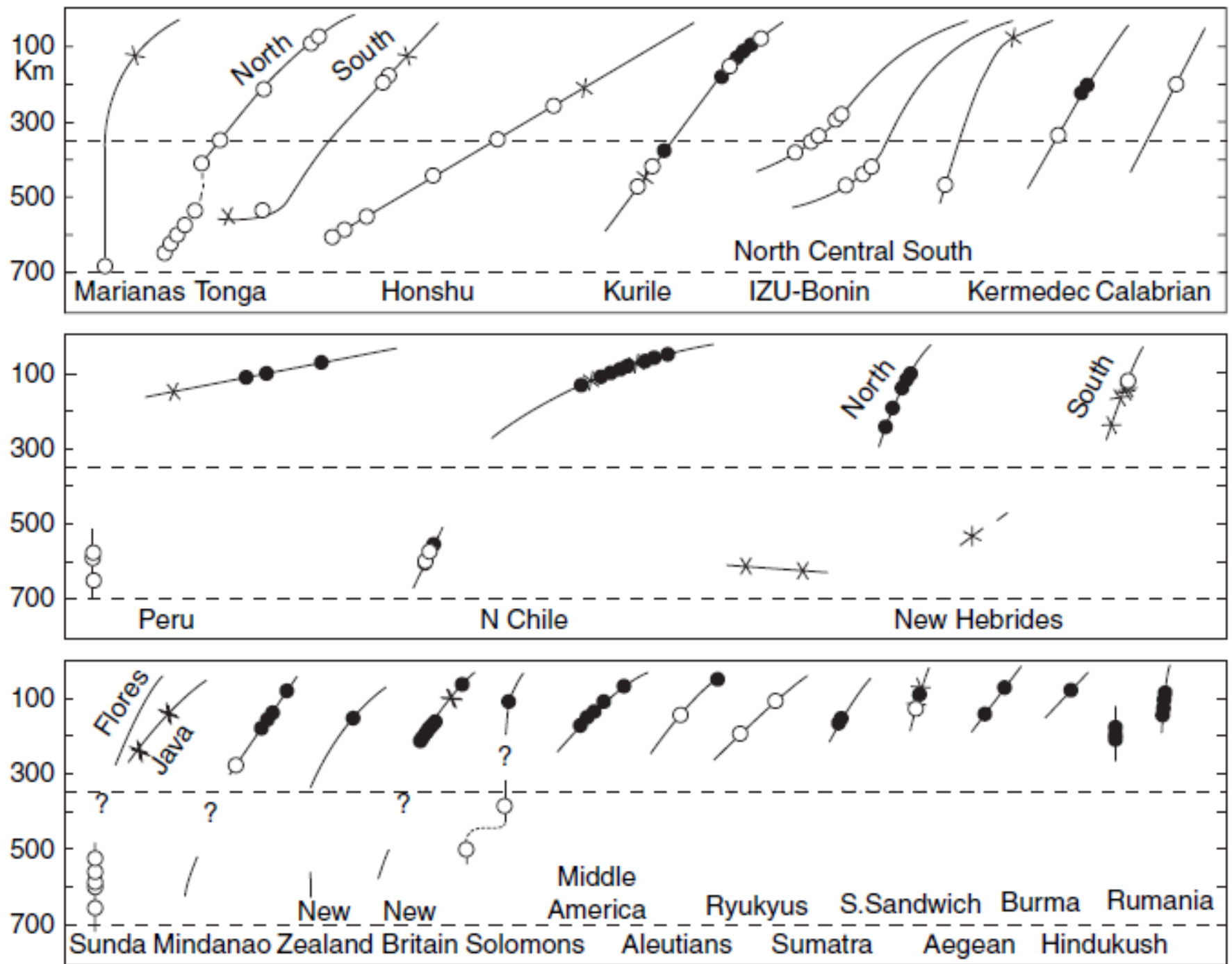
Lithgow-Bertelloni et al. (1993)

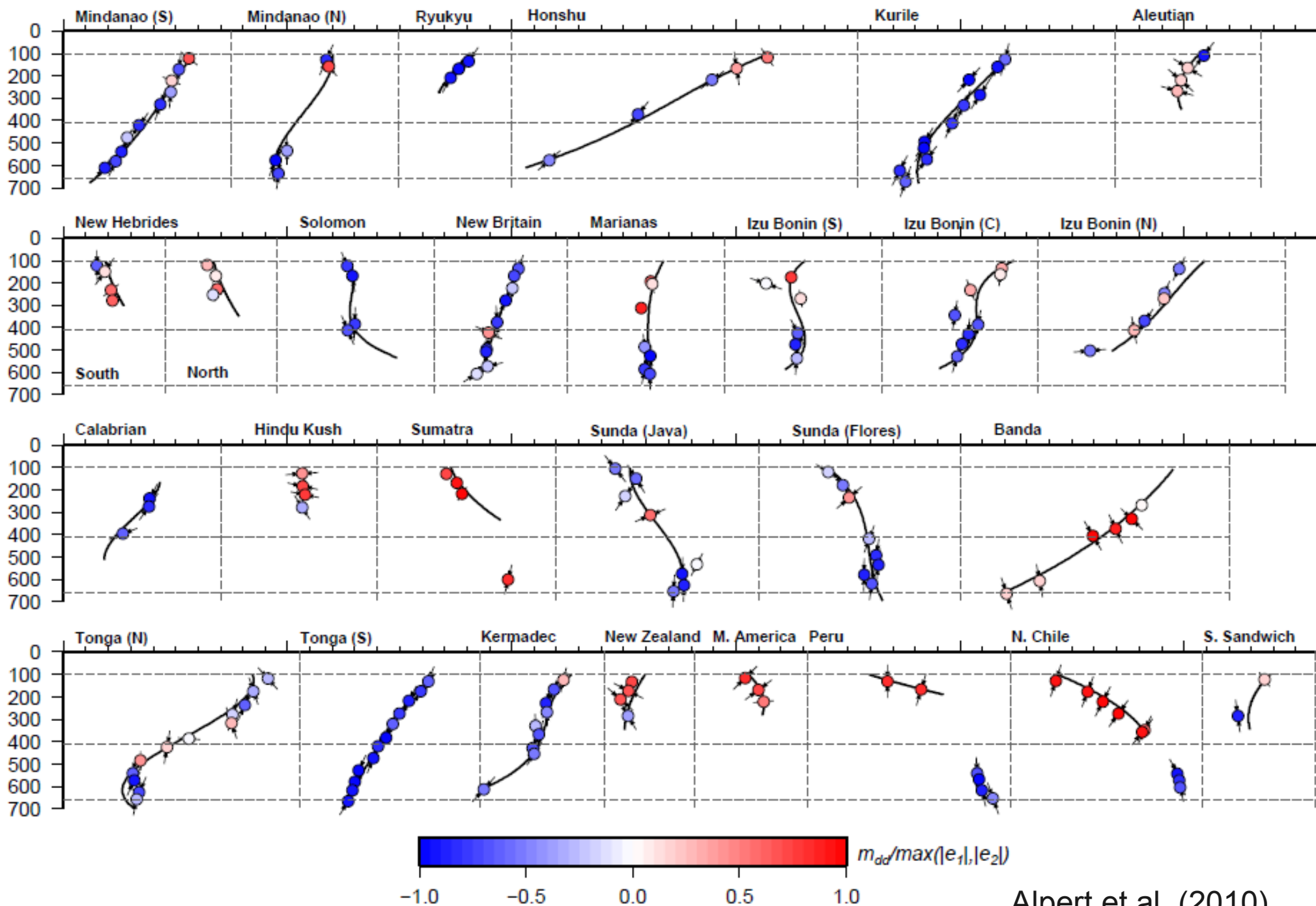


Seismicity constraints

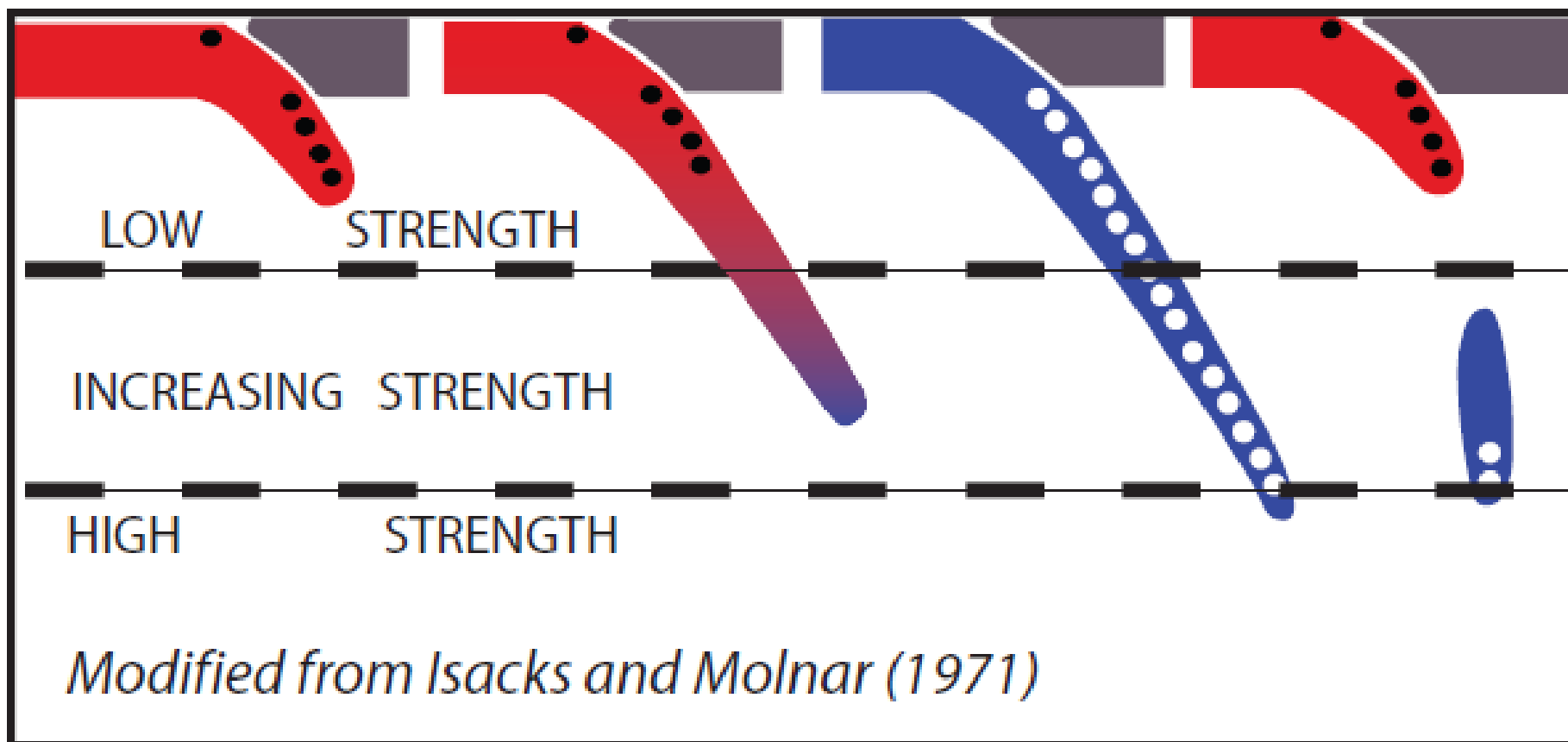


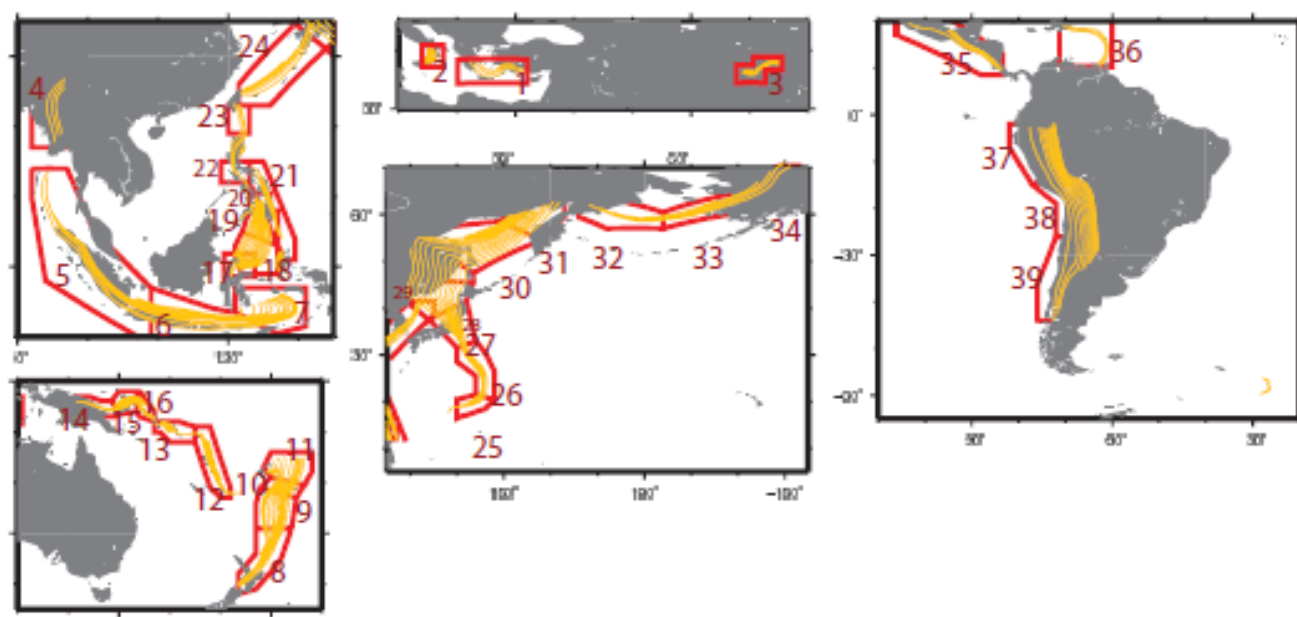
Alpert et al. (2010)



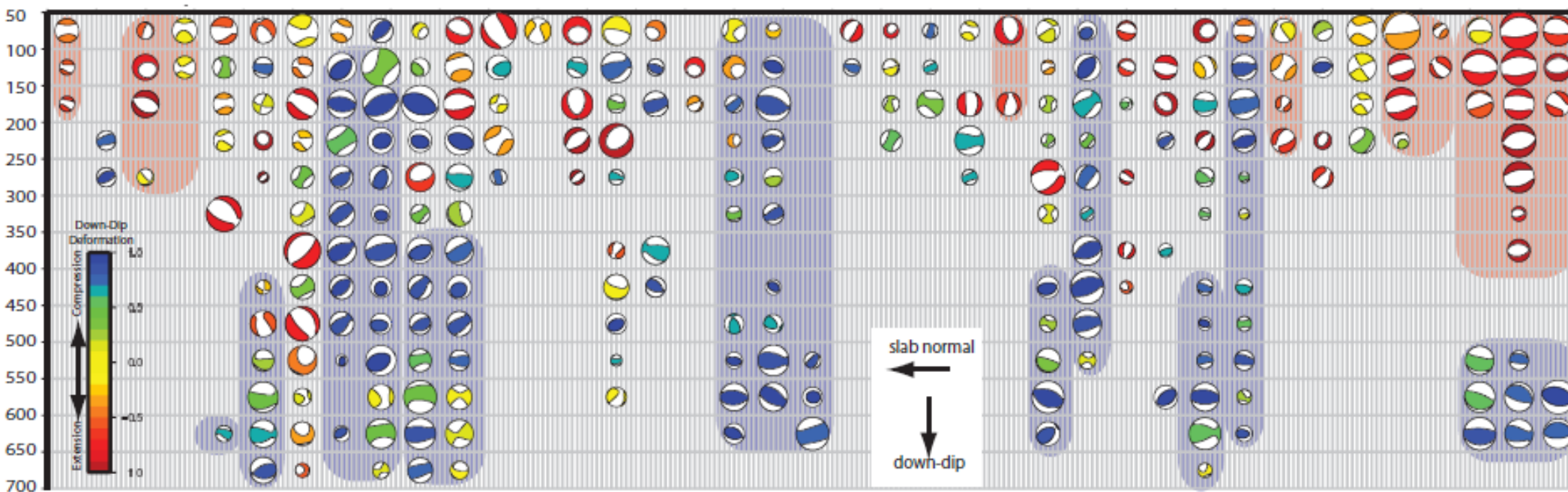


Alpert et al. (2010)



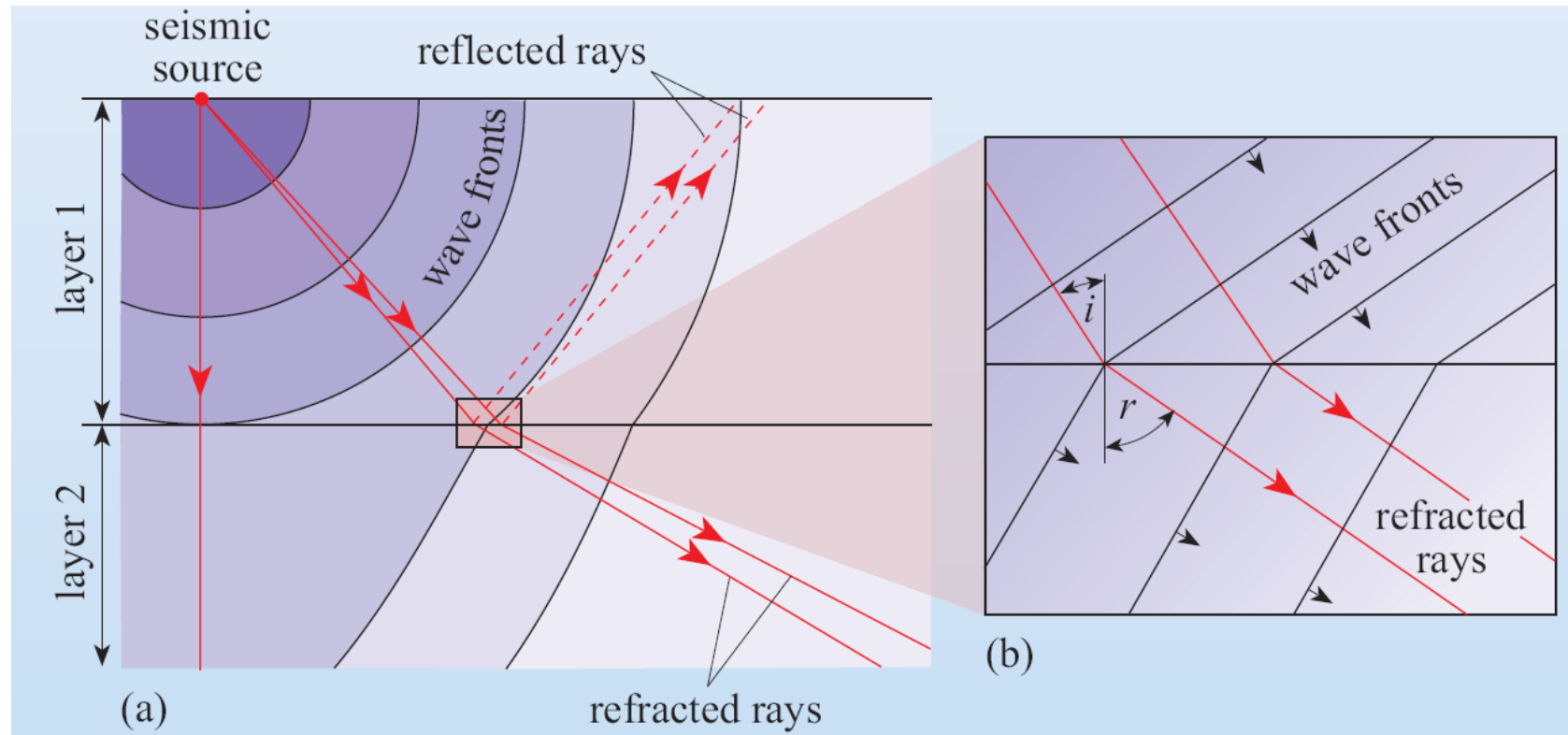


1. Hellenic Arc
2. Italia
3. Hindu Kush
4. Burma
5. Indonesia (W)
6. Indonesia (C)
7. Indonesia (E)
8. Tonga (S)
9. Tonga (C-1)
10. Tonga (C-2)
11. Tonga (N)
12. New Hebrides (S)
13. New Hebrides (N)
14. N. Britain (W)
15. N. Britain (C)
16. N. Britain (E)
17. Sulawesi
18. Mindanao (S)
19. Mindanao (C)
20. Mindanao (N)
21. E. Philippines
22. Luzon (S)
23. Luzon (N)
24. Ryukyu
25. Marianas (S)
26. Marianas (C)
27. Marianas (N)
28. Japan (S)
29. Japan (N)
30. Kurile (S)
31. Kurile (N)
32. Aleutians (W)
33. Aleutians (C)
34. Aleutians (E)
35. C. America
36. Caribbean
37. S. America (N)
38. S. America (C)
39. S. America (S)

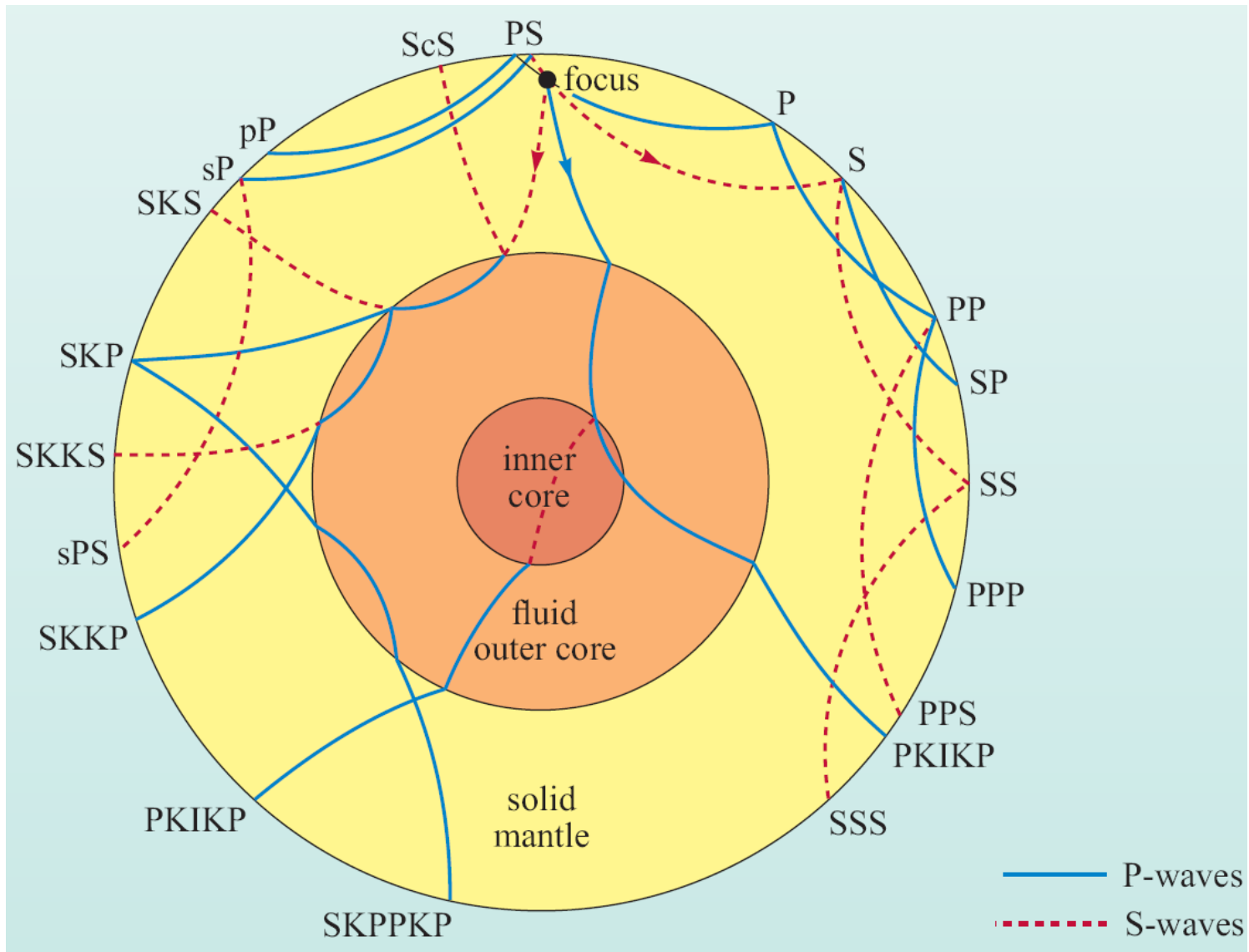


Structural seismology

Seismic waves

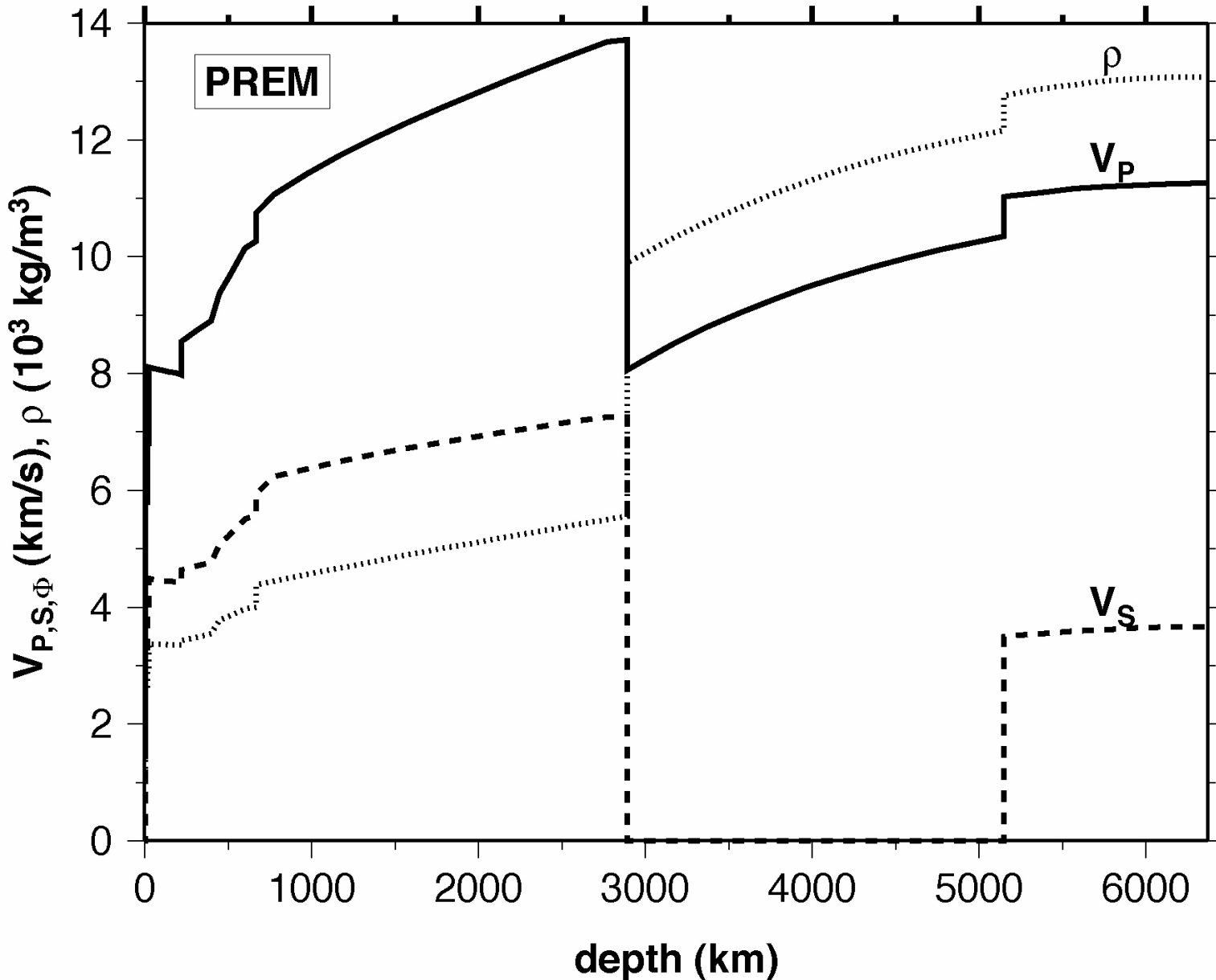


Seismic phases



PREM - Preliminary Earth Reference model (1981)

based on travel times, surface wave dispersion curves, normal mode frequencies, constraints on mass and moment of inertia + some constraints on V_P - ρ relation.



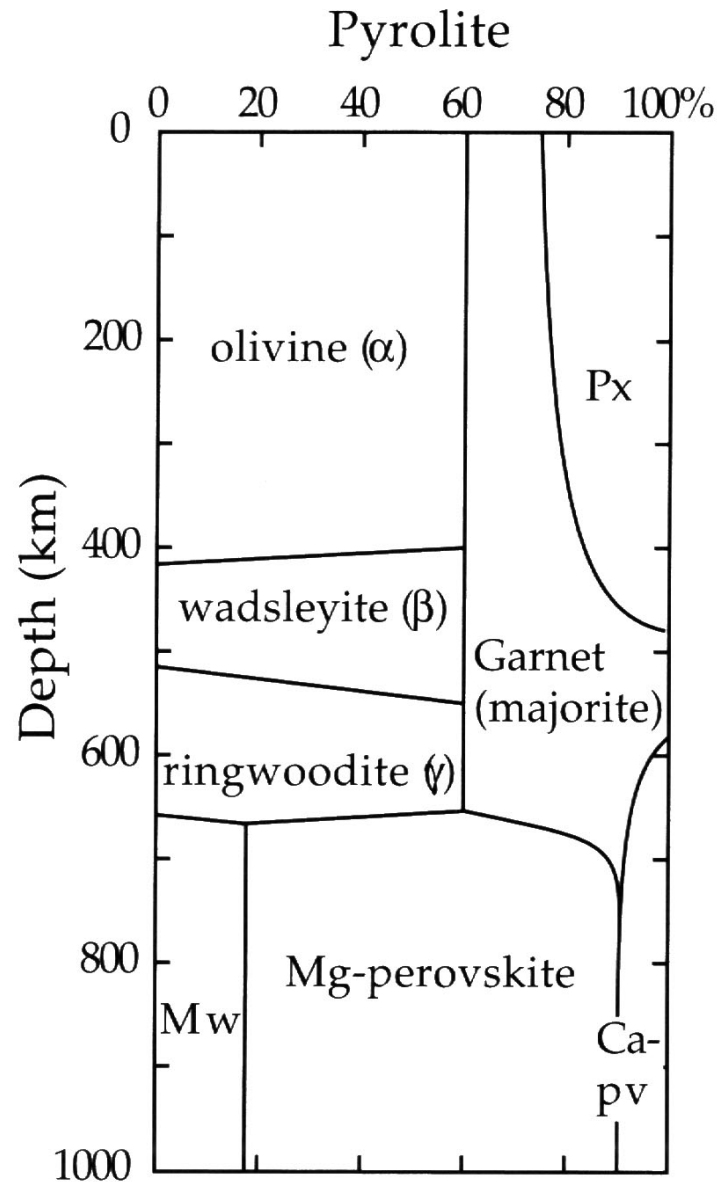
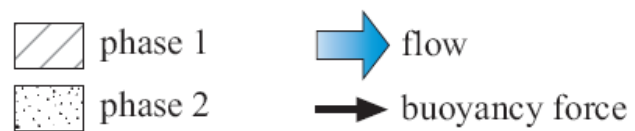
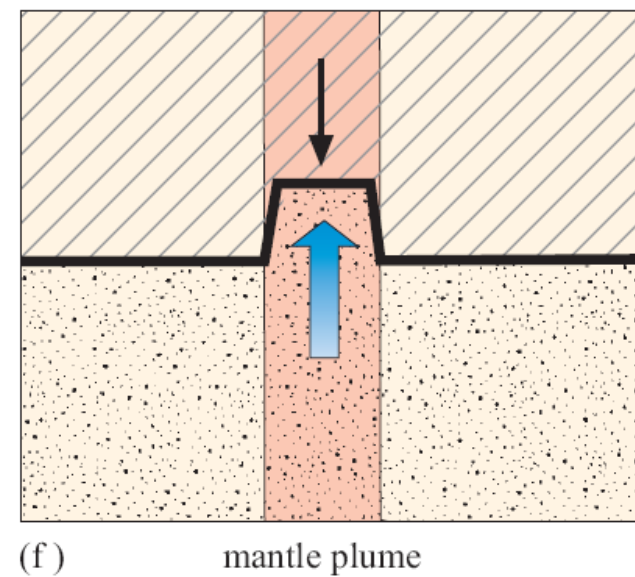
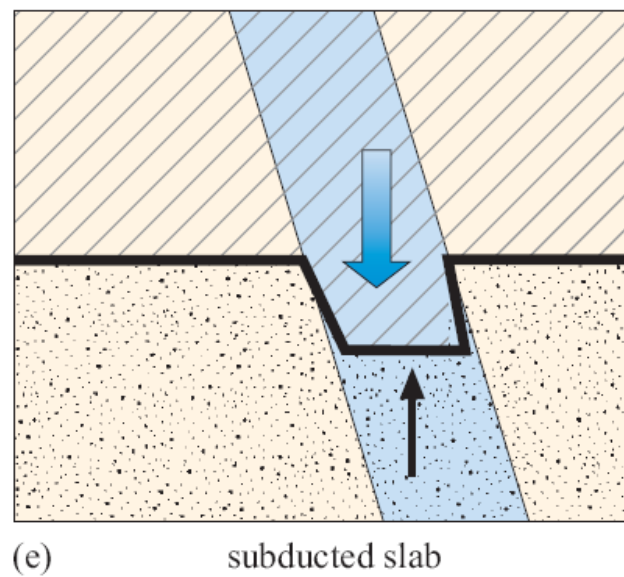
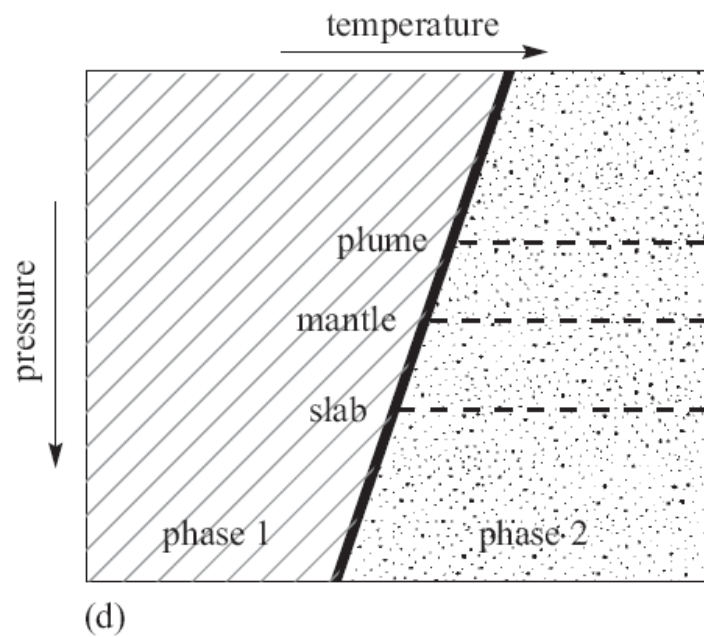
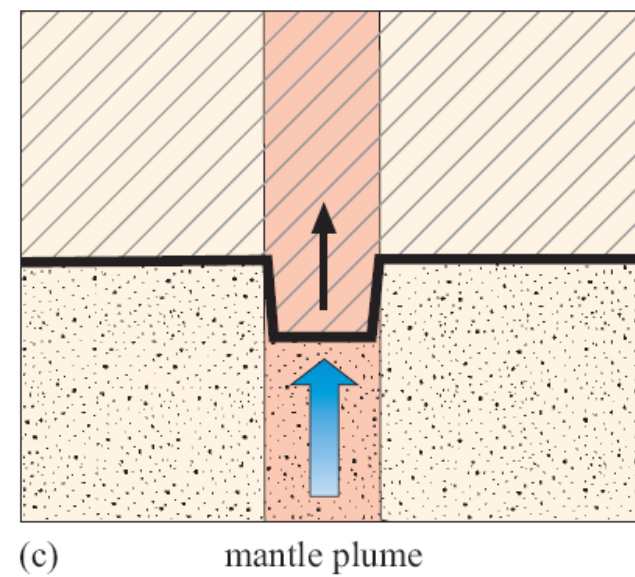
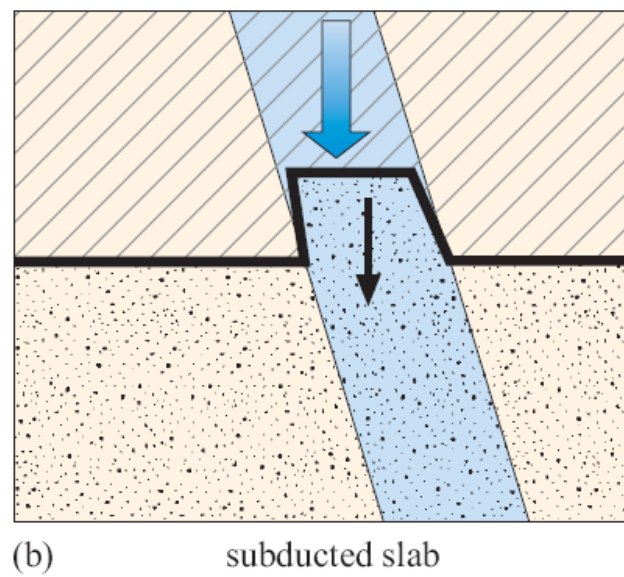
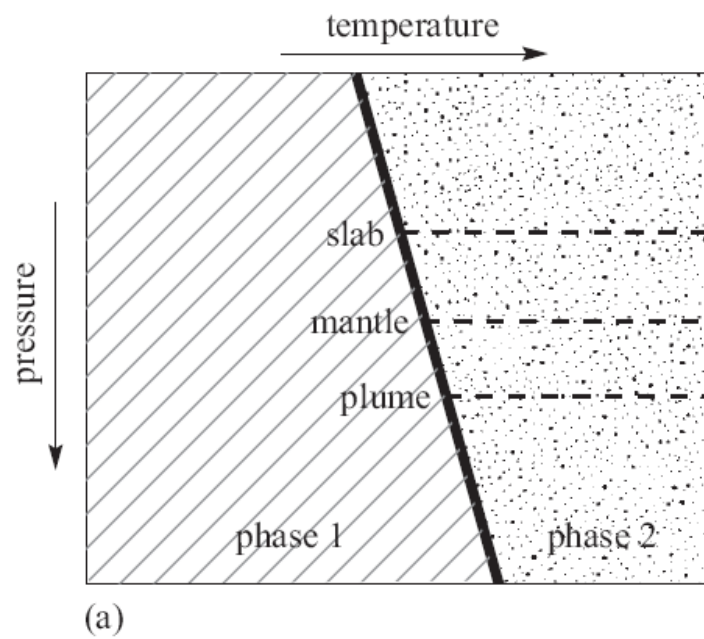
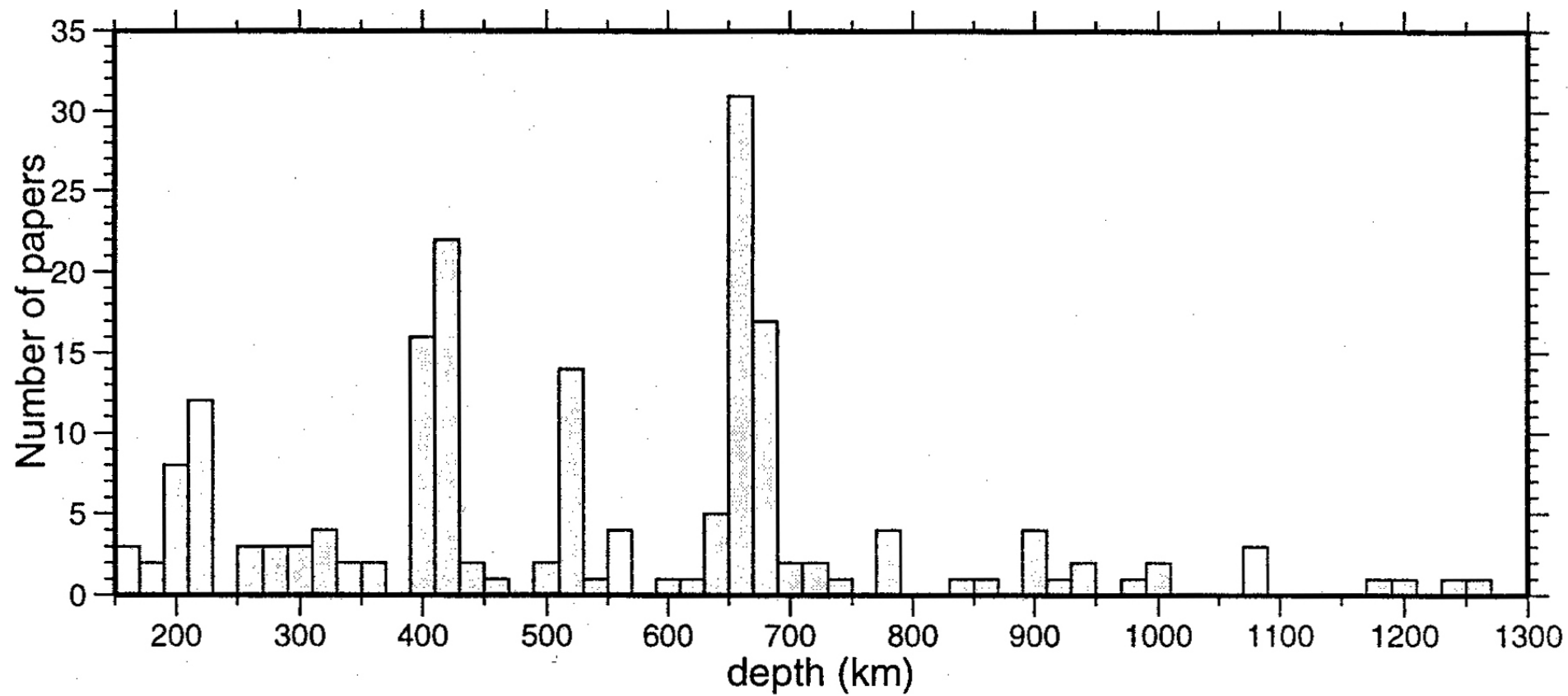


Figure 5.8. Sequence of pressure-induced transformations and reactions as a function of depth in a mantle of pyrolite composition. Px: pyroxene, Mw: magnesio-wüstite, pv: perovskite. After Irifune [44].

from Davies (1999)





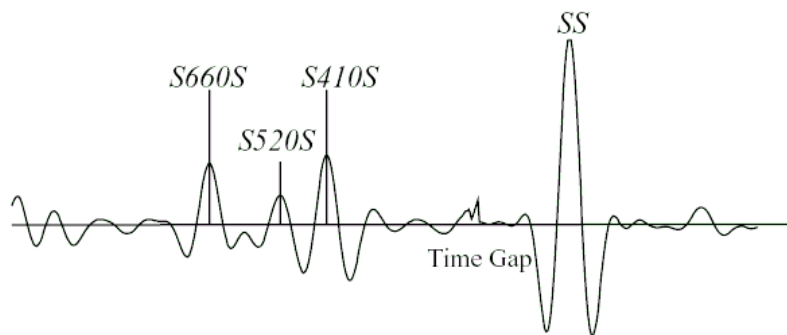
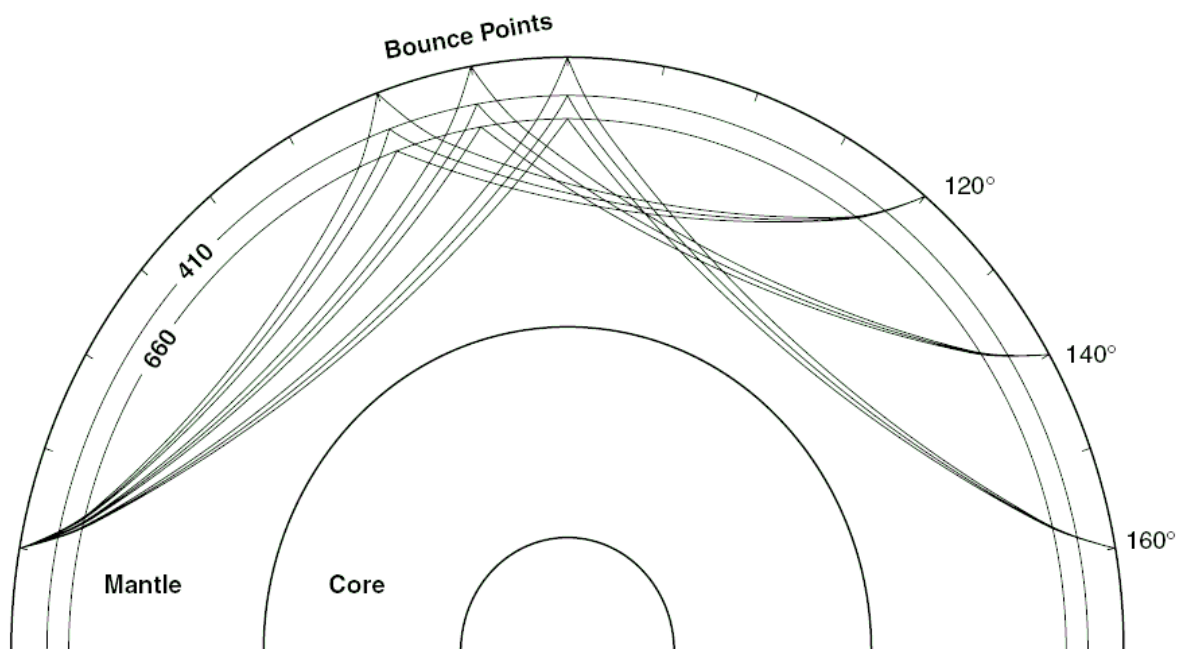
Deuss (2002)

Mineralogical zones	Seismological layers and Boundaries	Bullen's names	Dynamical layers
7 km	Crust	A	Lithosphere
Peridotite zone	Mohorovicic ~35 km	B	10-100 km
~350 km	(Upper mantle)		
Transition zone	410 km discontinuity	C	Upper Mantle
~750 km	(Transition zone)		
Perovskite zone	660 km discontinuity	D	Lower Mantle
	(Lower mantle)	D'	
	~2750 km		TBL
	(D'') 2889 km	D''	CBL
	Core-mantle boundary	E	
Outer Core		F	
Inner core boundary 5154 km		G	
Inner Core			

Figure 5.2. A partial terminology of mantle layers that distinguishes the different concerns and usages of mineral physics, seismology and dynamical layers. TBL: thermal boundary layer; CBL: chemical boundary layer.

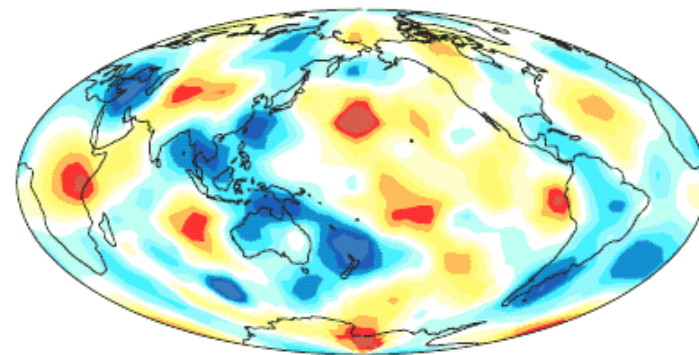
from Davies (1999)

SS precursors

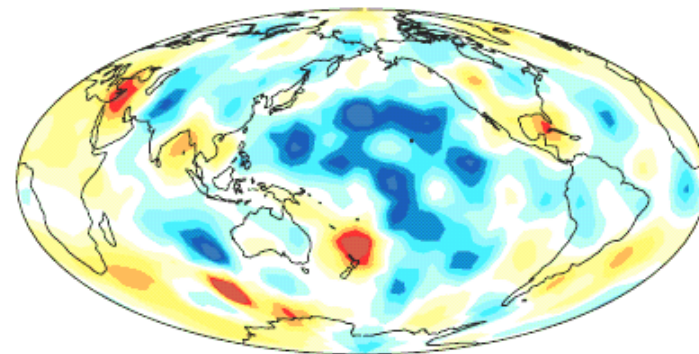


Crust and Velocity Corrected Discontinuities

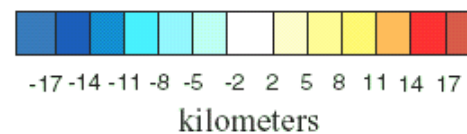
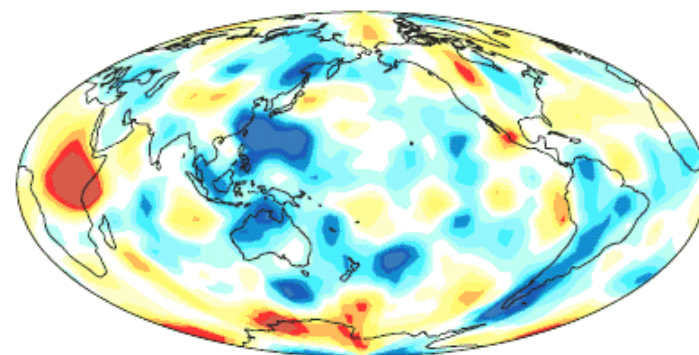
Transition Zone Thickness



410 Topography

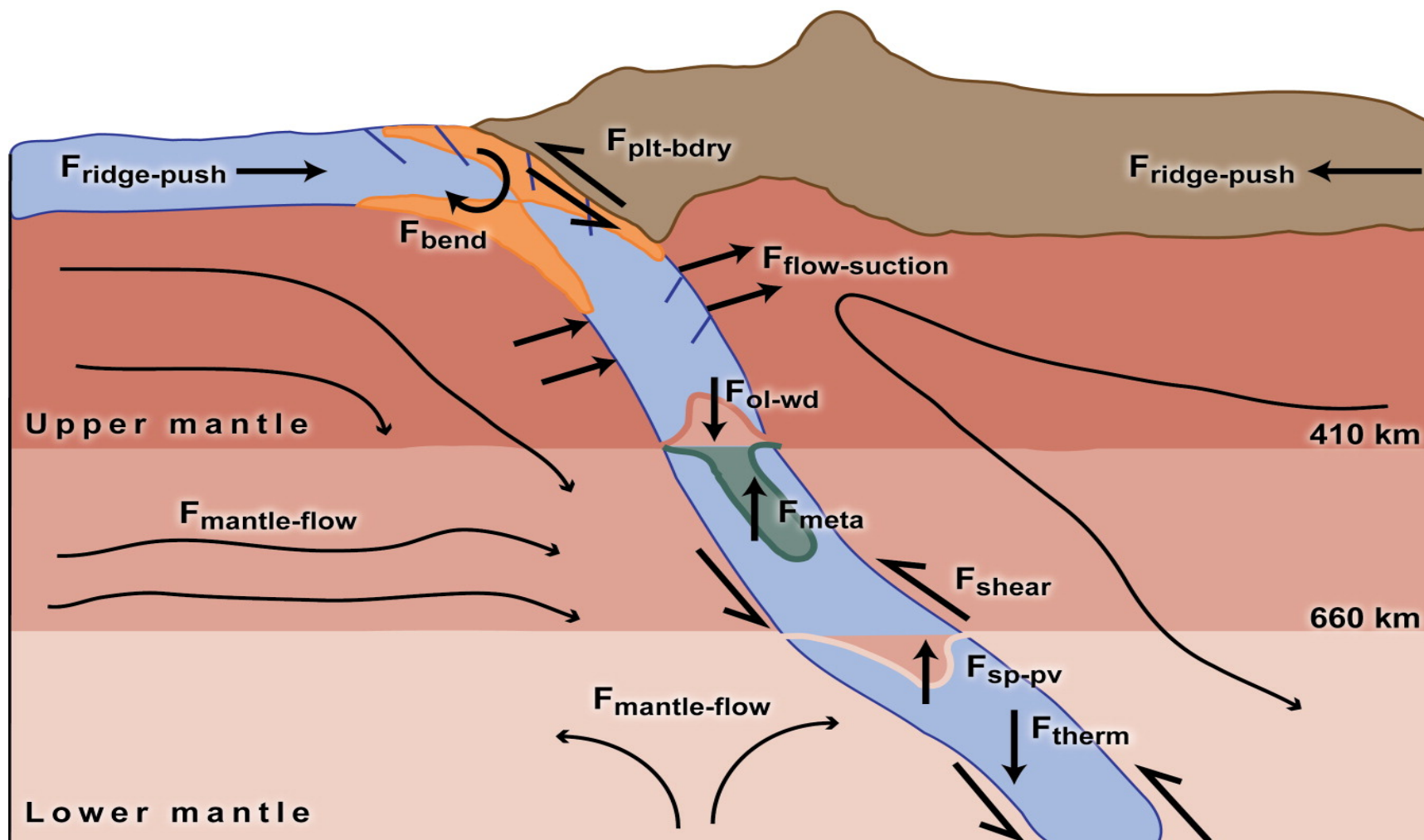


660 Topography



Houser et al. (2008)

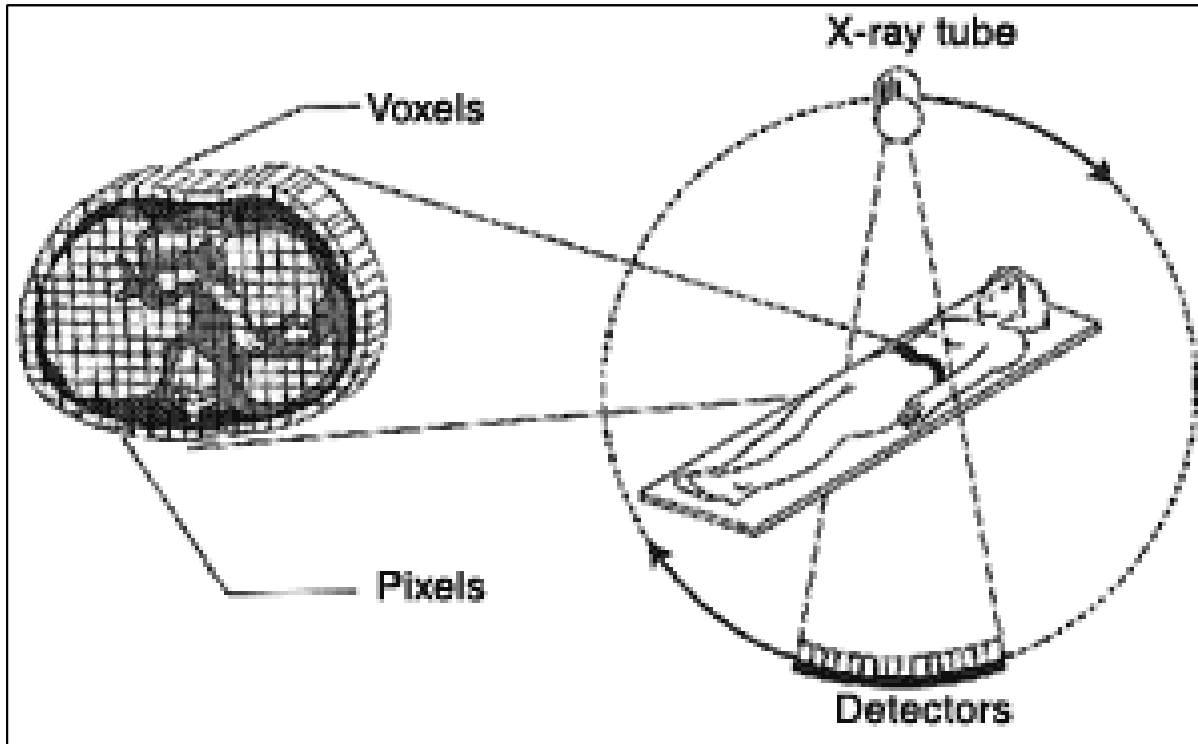
A modified set of forces



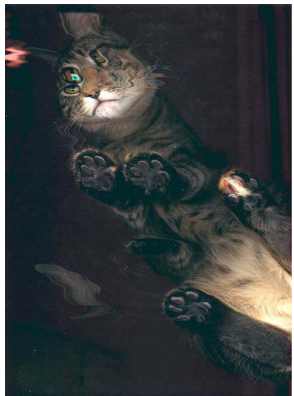
Billen MI. 2008.

Annu. Rev. Earth Planet. Sci. 36:325–56.

Seismic tomography



- CAT-scan like technique
- Earthquakes : sources
Seismometers : receivers
- Measure travel times to invert for 3-D velocity structure



Theory of linear, ray tomography

Travel time t of a ray given by integral over slowness $p = 1/v$ along path s

$$t = \int_{\text{path}} p[r(s), \theta(s), \phi(s)] ds$$

Travel time anomaly δt from reference model given by slowness perturbations δp , same path

$$\delta t = \int_{\text{path}} \delta p[r(s), \theta(s), \phi(s)] ds$$

Choose n basis functions for parameterization

$$\delta p(r, \theta, \phi) = \sum_{i=1}^n \delta p_i f_i(r, \theta, \phi)$$

Evaluate path integral for every δt observation for m number of observations

$$\delta t = \sum_{i=1}^n \delta p_i \int_{\text{path}} f_i[r(s), \theta(s), \phi(s)] ds$$

$$\delta t_j = \sum_{i=1}^n \delta p_i \int_{\text{path}_j} f_i[r(s), \theta(s), \phi(s)] ds \quad (j = 1, \dots, m)$$

Write in matrix form, A is $m \times n$, solve

$$A_{ji} = \int_{\text{path}_j} f_i[r(s), \theta(s), \phi(s)] ds \quad A \cdot x = d$$

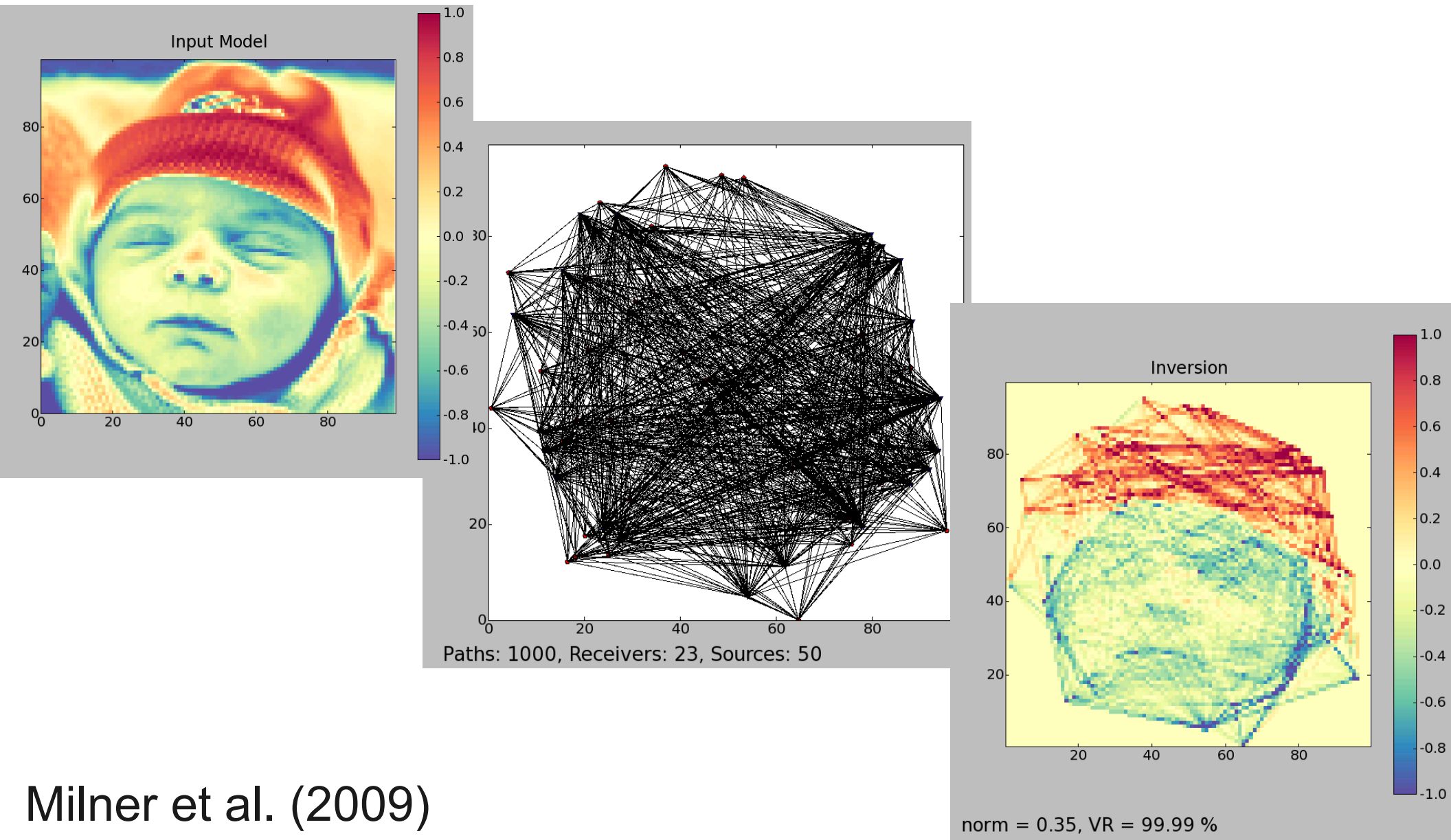
For $m > n$, system is overdetermined (and δt has errors), solve in least squares sense, $(A^T A)^{-1}$ is the “general inverse”

$$\|A \cdot x - d\| = \min \iff x = x_{LS} \quad A^T \cdot A \cdot x = A^T \cdot d \quad x_{LS} = (A^T \cdot A)^{-1} \cdot A^T \cdot d$$

Tomographic problems typically ill-conditioned, introduce regularization (“damping”):

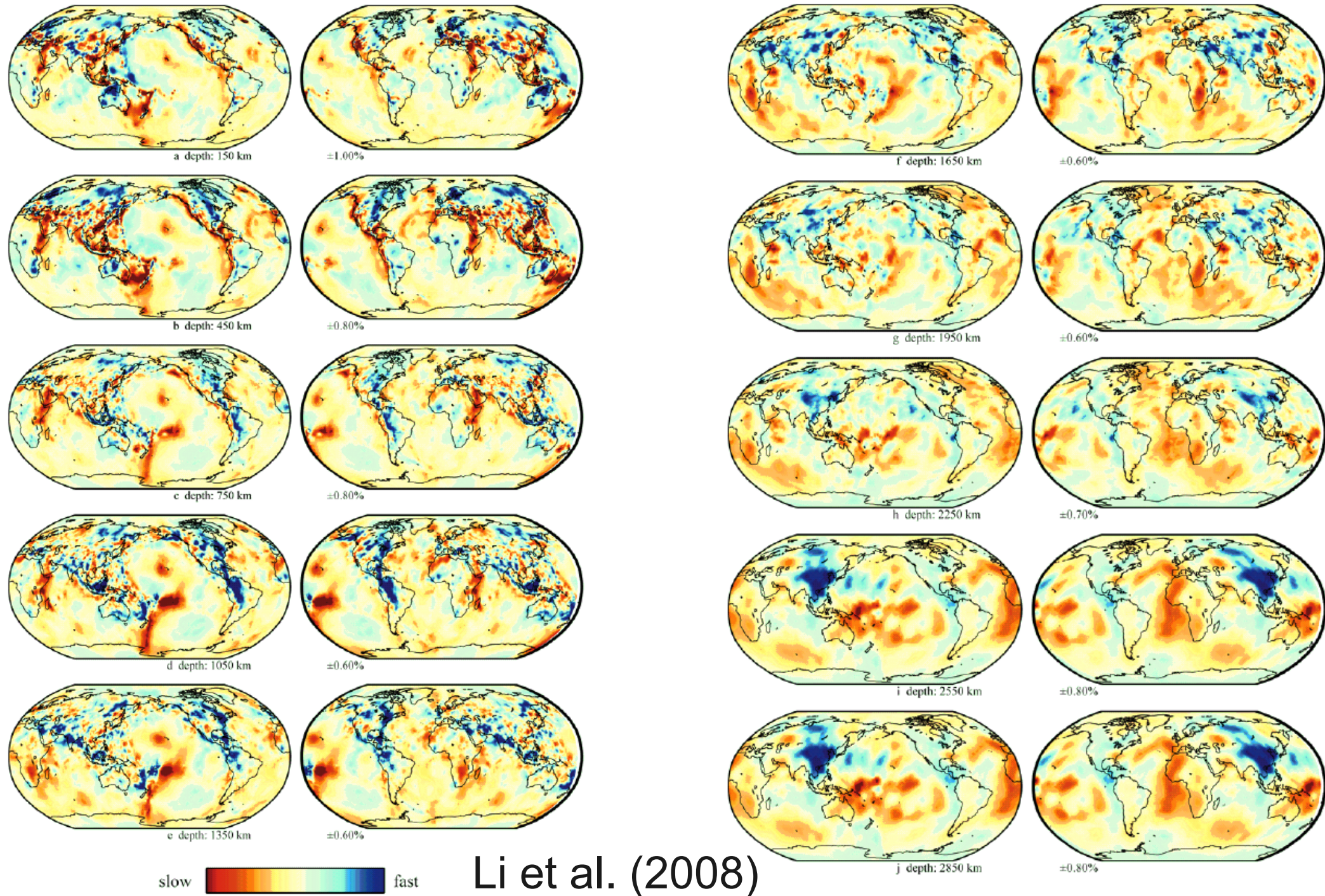
$$B \cdot x = c \quad \begin{bmatrix} A \\ \lambda B \end{bmatrix} \cdot x = \begin{bmatrix} d \\ \lambda c \end{bmatrix} \quad x_{LS} = [A^T \cdot A + \lambda^2 B^T \cdot B]^{-1} \cdot [A^T \cdot d + \lambda^2 B^T \cdot c]$$

Tomography example

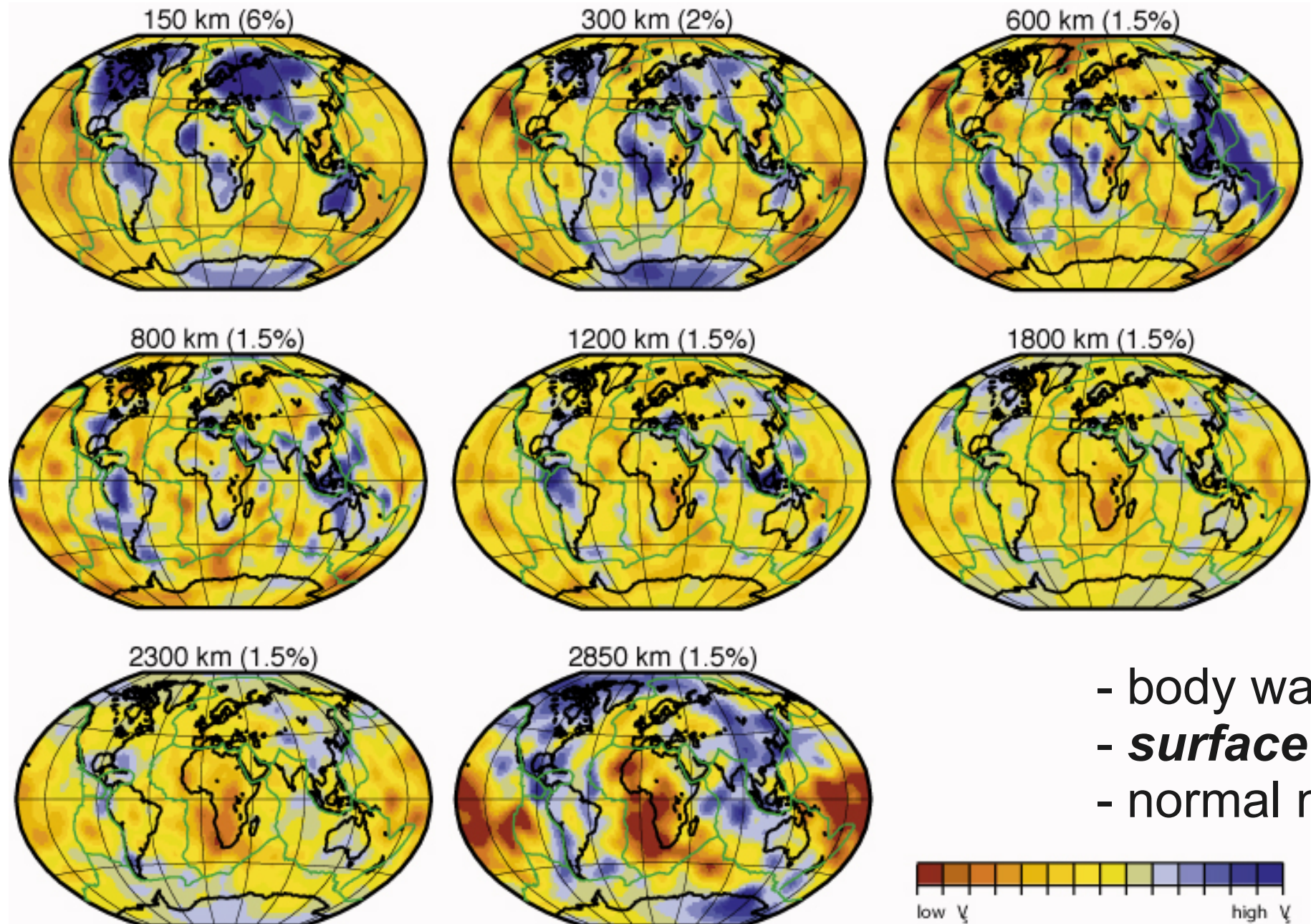


Milner et al. (2009)

P wave tomography



S wave tomography

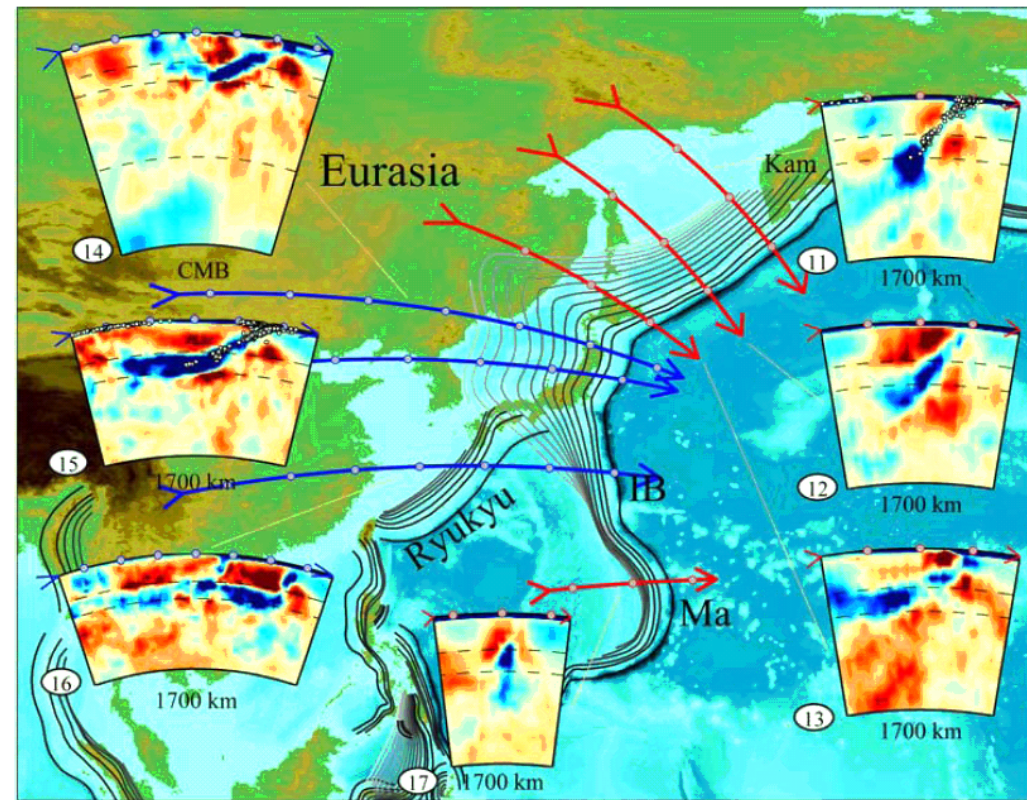
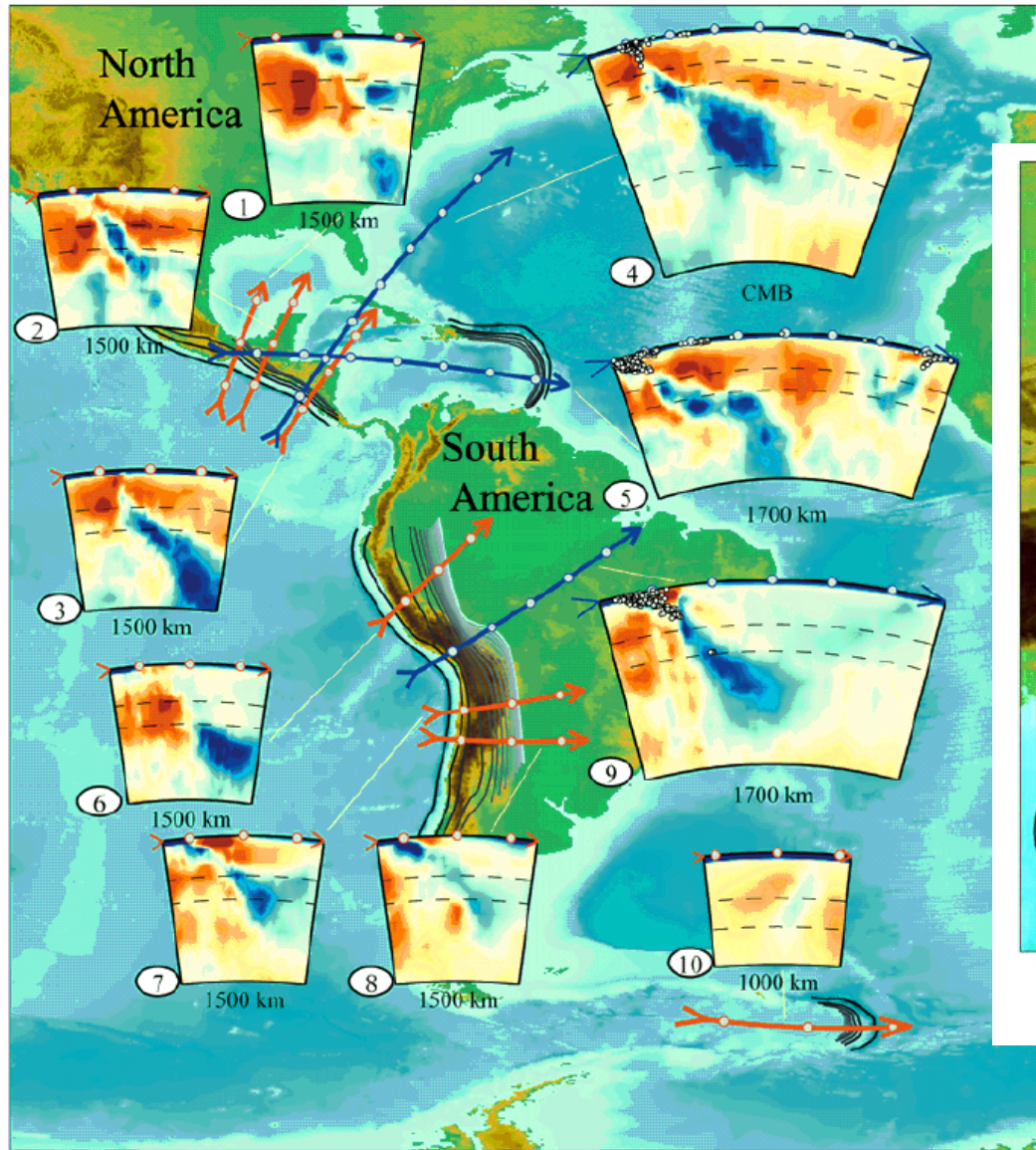



- body waves
- **surface waves**
- normal modes


Figure 8: Depth slices through S20R TS. High (relatively to PREM) shear velocity regions are indicated by blue and red colours, respectively, with an intensity that is proportional to the amplitude of the shear velocity perturbations. The range of shear velocity variation (in %) is given above each map. Green lines represent plate boundaries.

Ritsema et al. (2000)

Slabs in the lower mantle

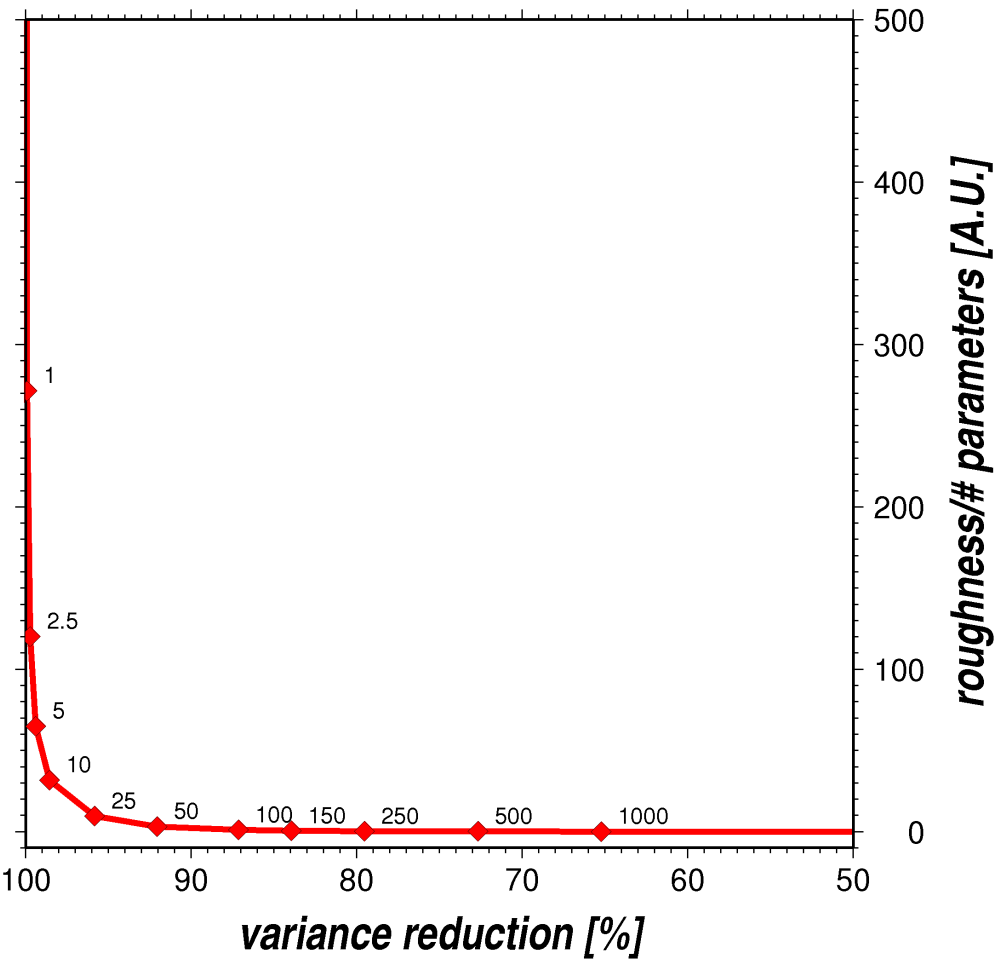
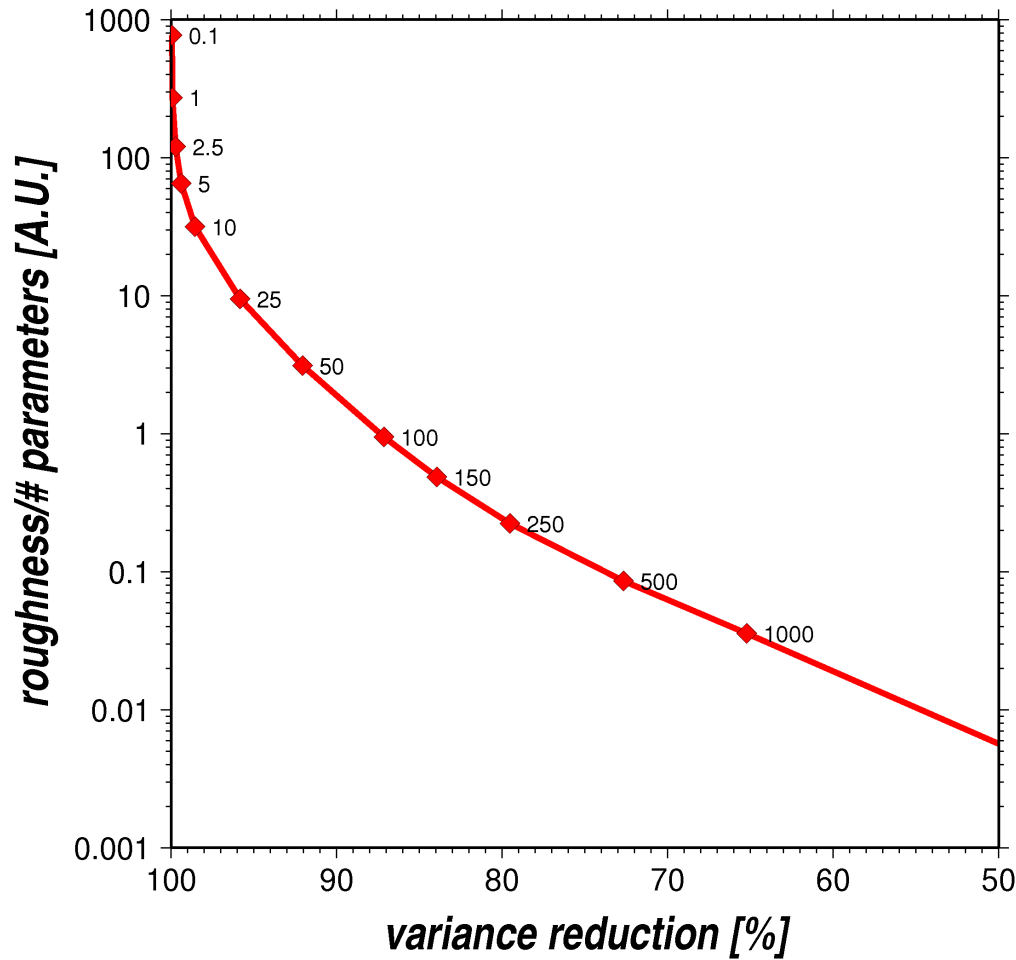


-0.8%  +0.8%
P-wavespeed perturbation

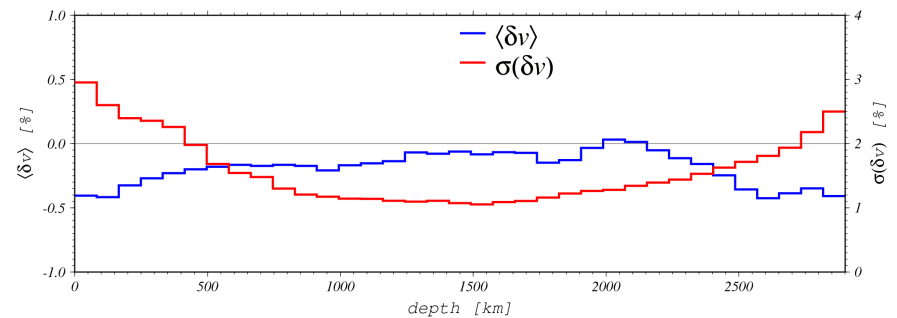
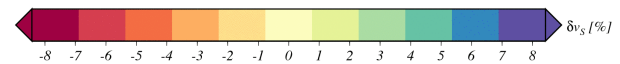
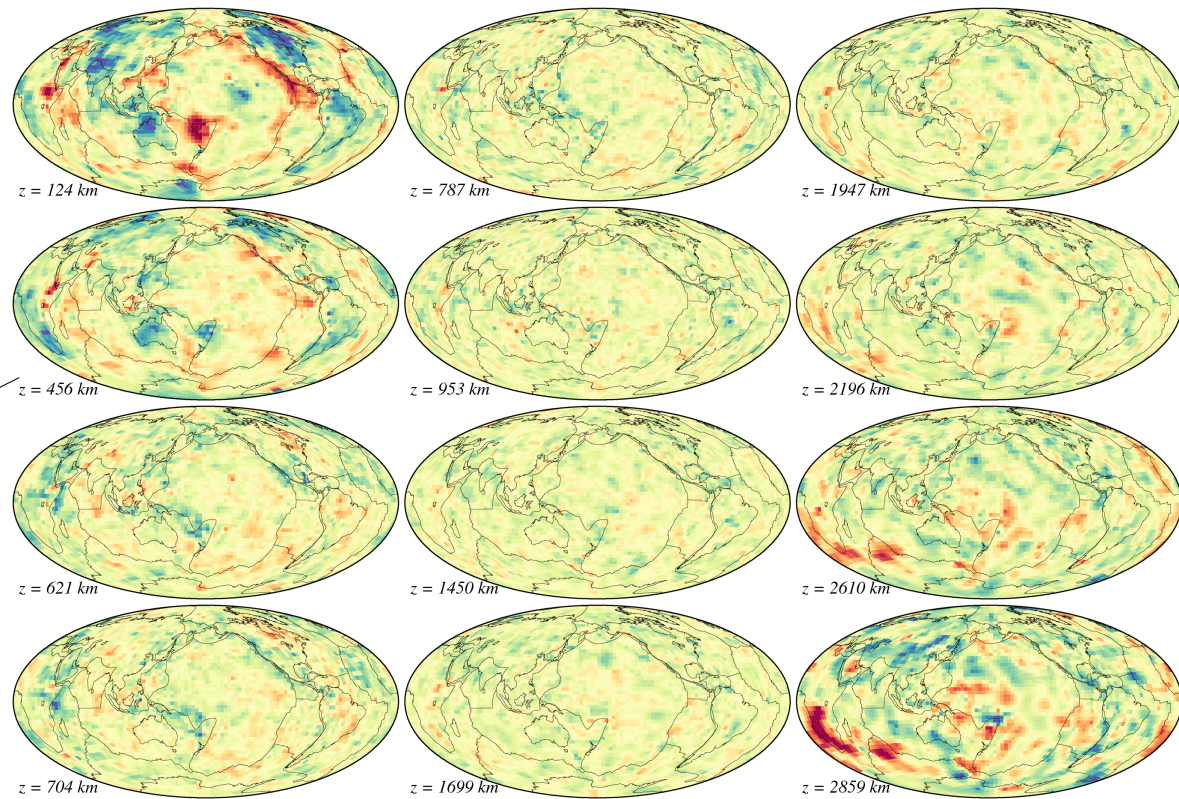
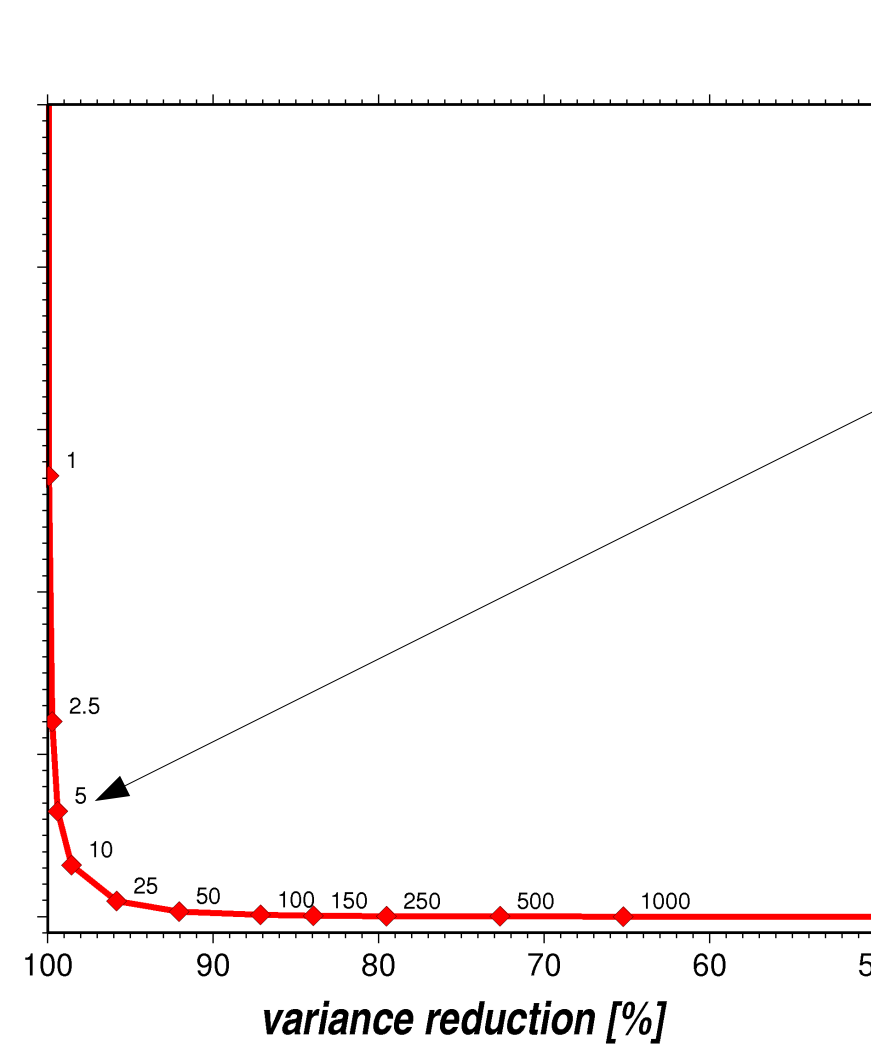
-0.8%  +0.8%
P-wavespeed perturbation

Li et al. (2008)

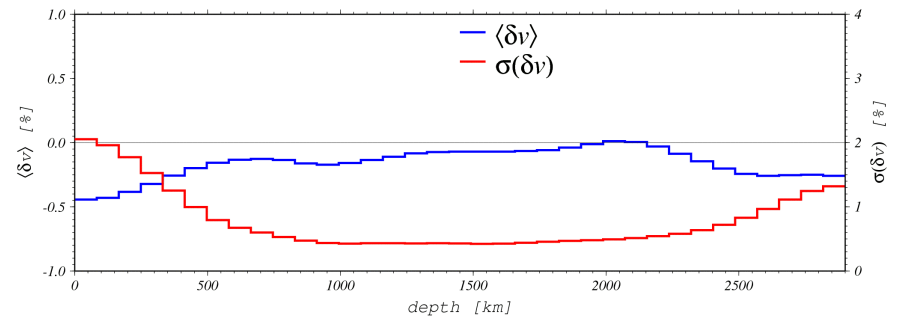
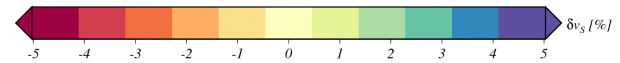
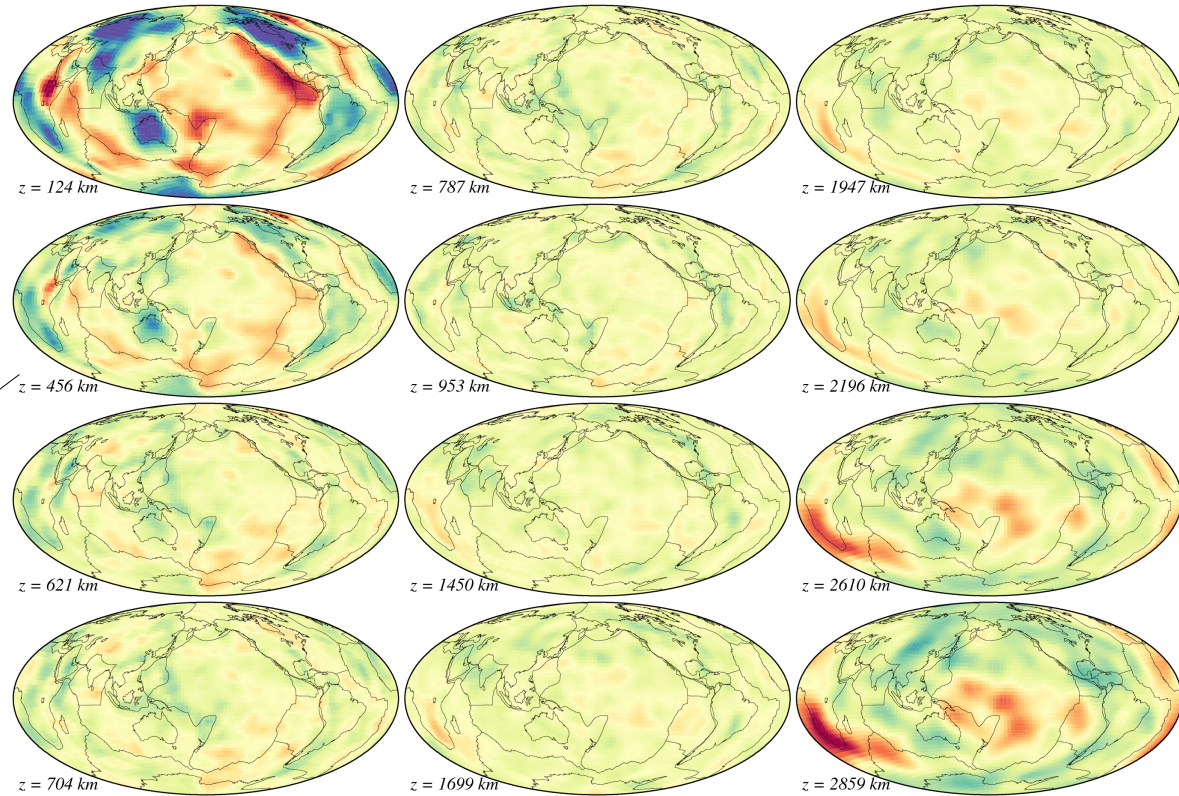
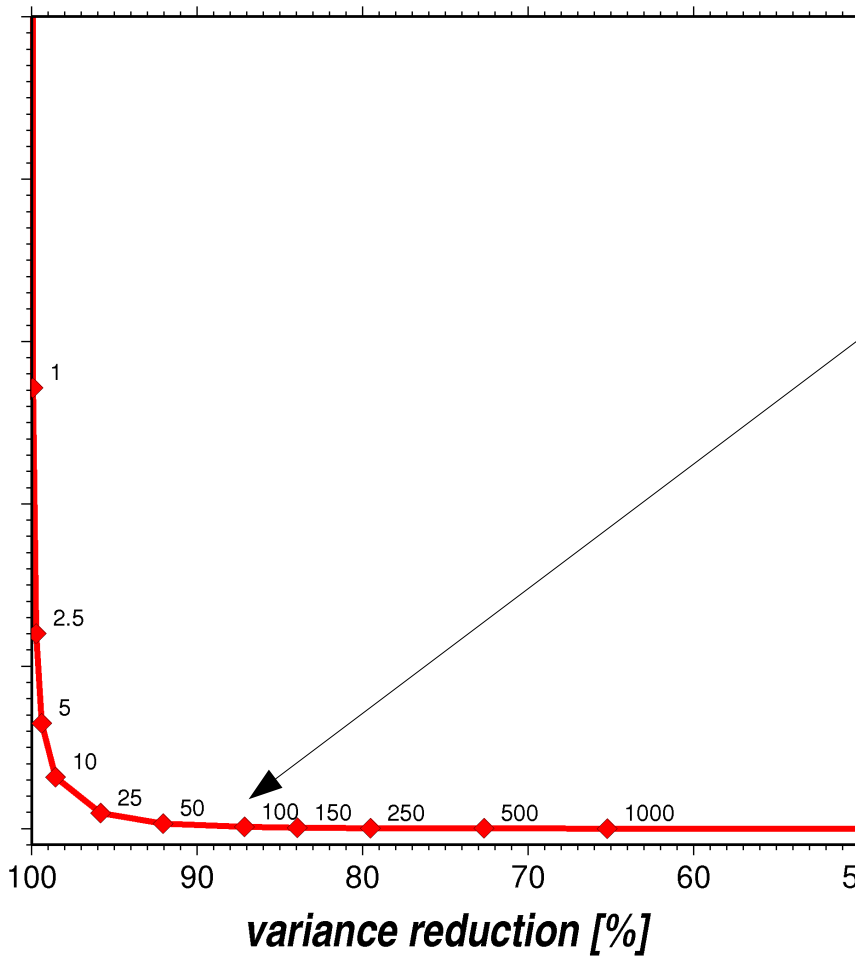
Damping and L curves



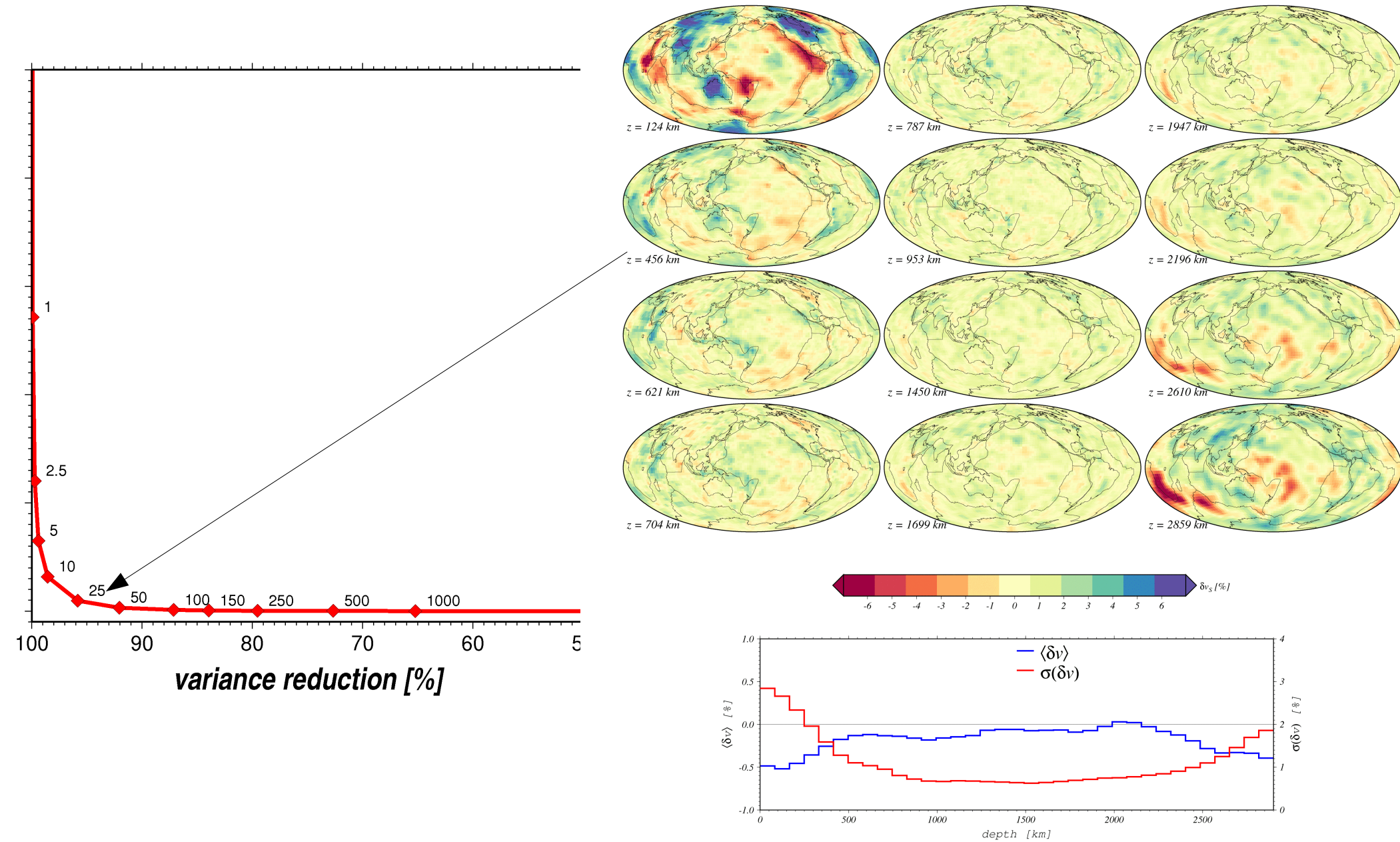
Damping and L curves



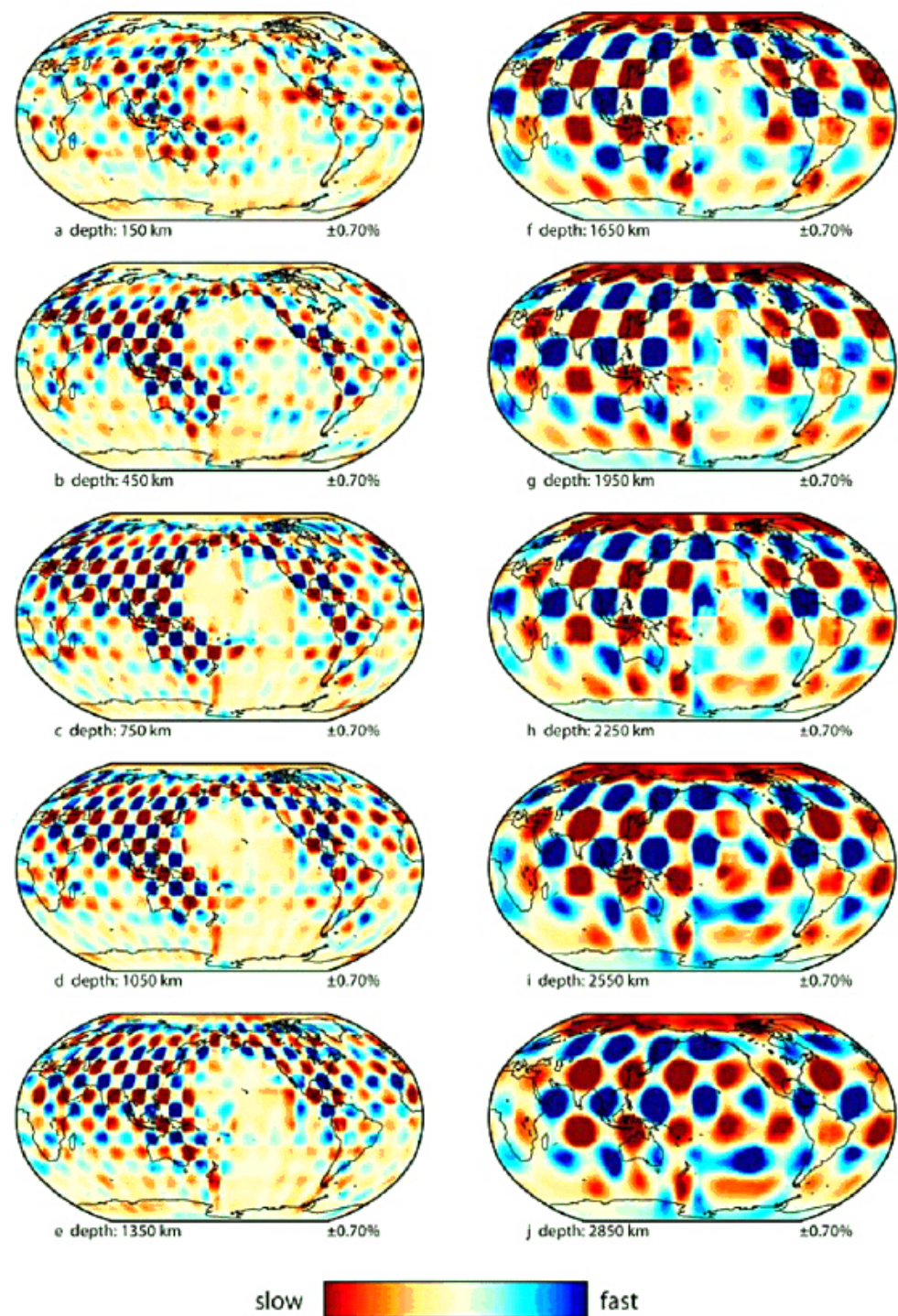
Damping and L curves



Damping and L curves



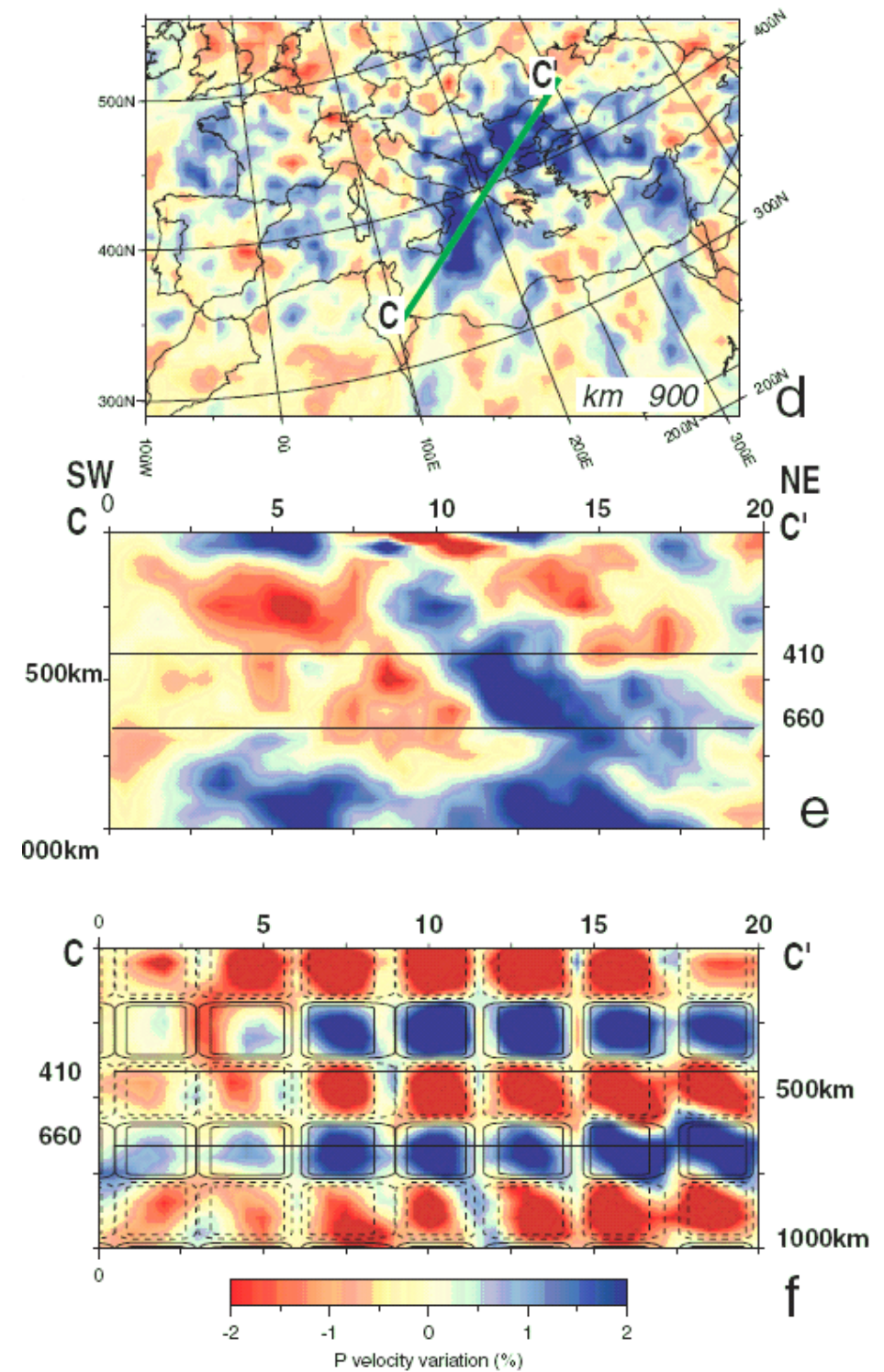
P wave tomography checkerboards



Li et al. (2008)

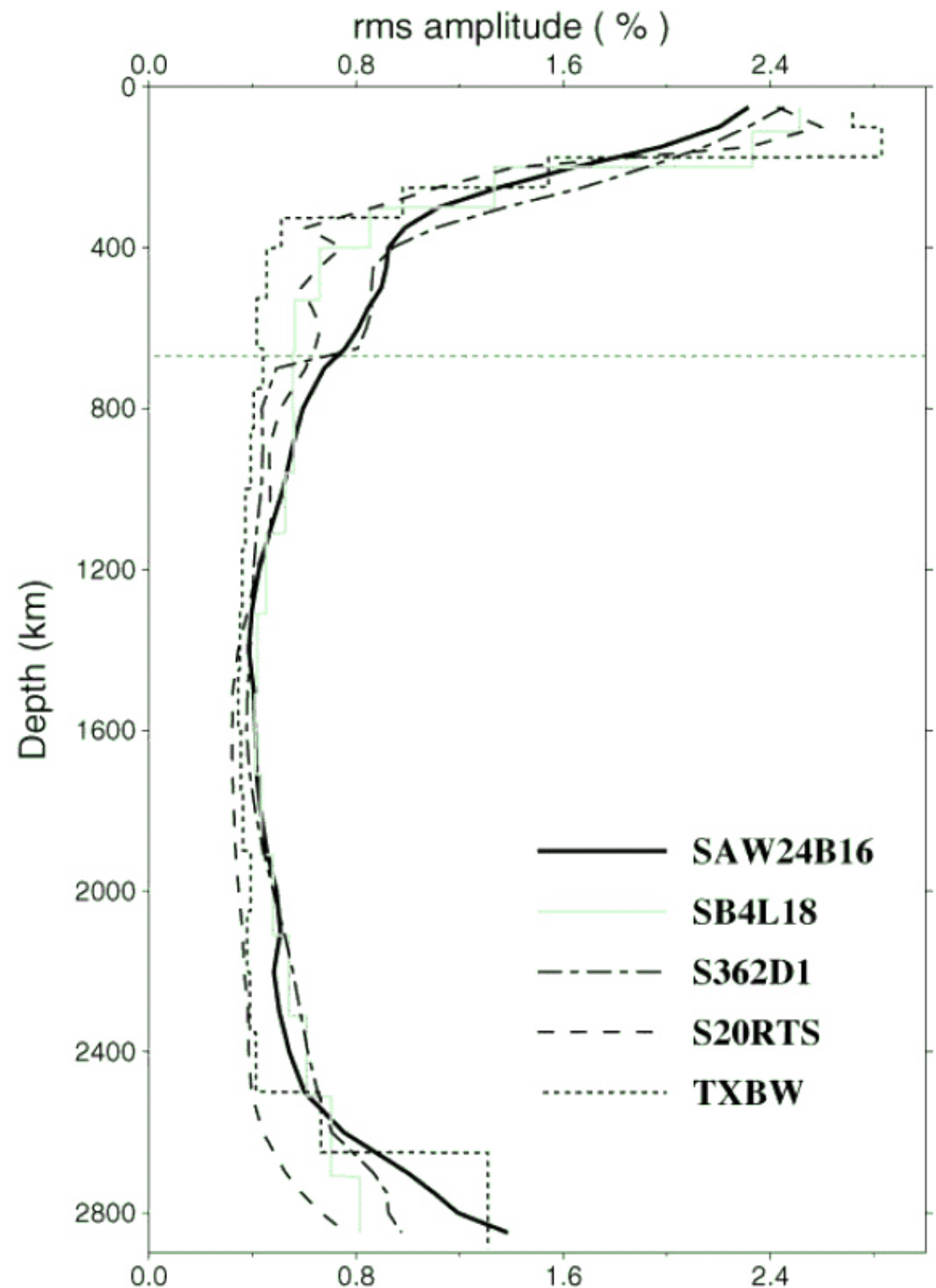
Figure 5. Recovery fields of global resolution tests, using harmonic input patterns with constant amplitude $\pm 1.5\%$ in wave speed throughout the mantle. (a–e) Half-wavelength of $\sim 5^\circ$ (spatial wavelength of ~ 550 km at the surface). (f–j) Half-wavelength of $\sim 10^\circ$ (~ 600 km at the CMB).

Cross-section tests



Faccenna et al. (2003)

Quantitative analysis: RMS power



Spherical harmonics

Definition

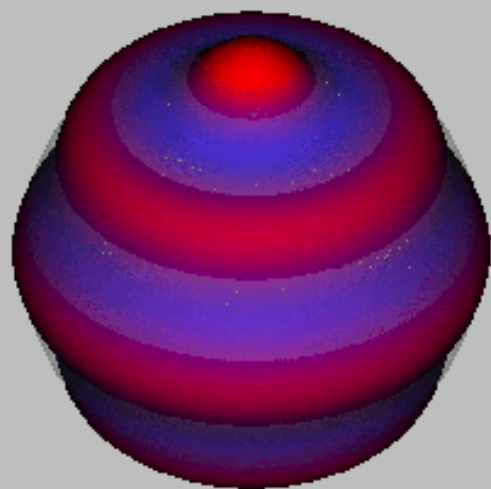
$$\delta v(\theta, \phi) \approx \sum_{l=0}^{\ell_{\max}} \left[a_{\ell 0} X_{\ell 0}(\theta) + \sqrt{2} \sum_{m=1}^{\ell} X_{\ell m}(\theta) \times (a_{\ell m} \cos m\phi + b_{\ell m} \sin m\phi) \right]$$

Power per degree
and unit area

$$\sigma_{\ell}^2 = \frac{1}{2\ell + 1} \sum_{m=0}^{\ell} (a_{\ell m}^2 + b_{\ell m}^2)$$

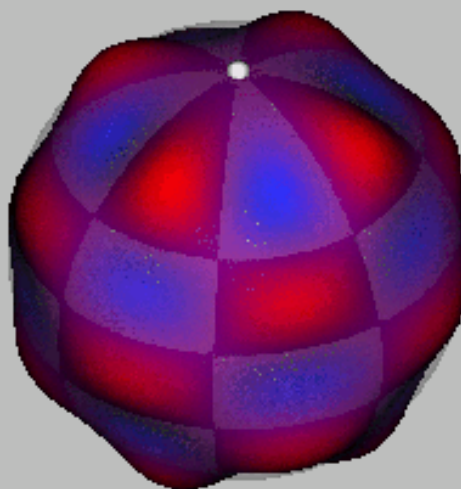
Correlation $r^{\ell} = \frac{\sum_{m=0}^{\ell} (a_{\ell m} c_{\ell m} + b_{\ell m} d_{\ell m})}{\sqrt{\sum_{m=0}^{\ell} (a_{\ell m}^2 + b_{\ell m}^2)} \sqrt{\sum_{m=0}^{\ell} (c_{\ell m}^2 + d_{\ell m}^2)}}$

$l = 8, m = 0$



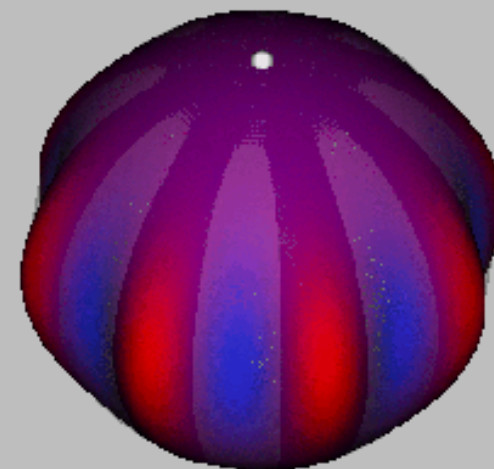
zonal

$l = 8, m = 4$



tesseral

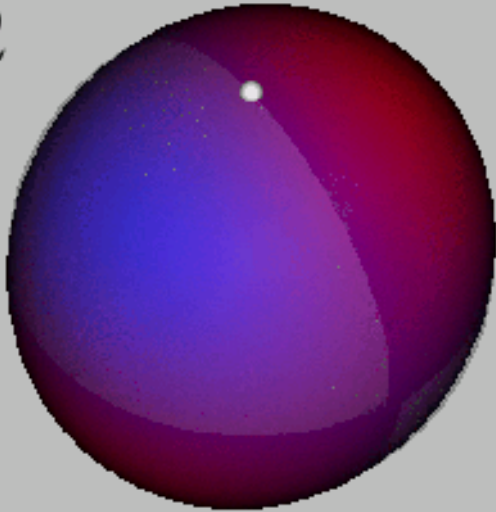
$l = 8, m = 8$



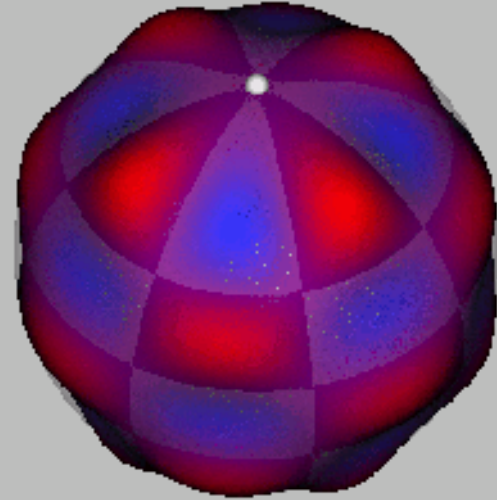
sectoral

Spherical harmonics: tesseral wavelengths

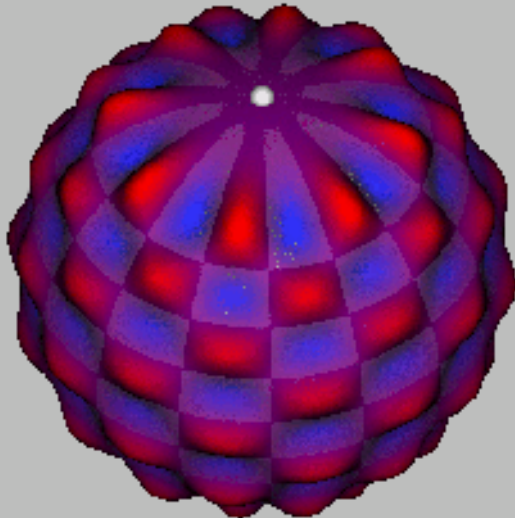
$l=2$



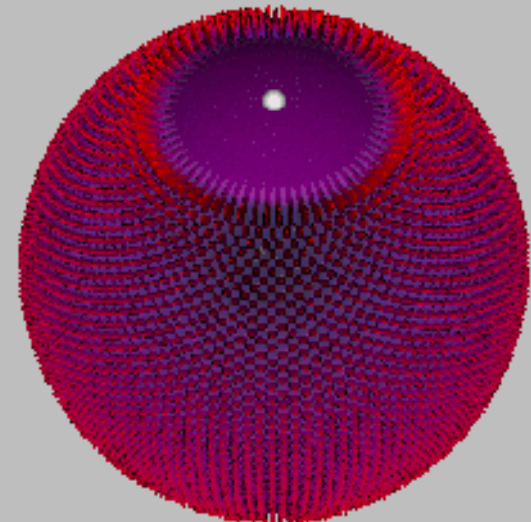
$l=8$



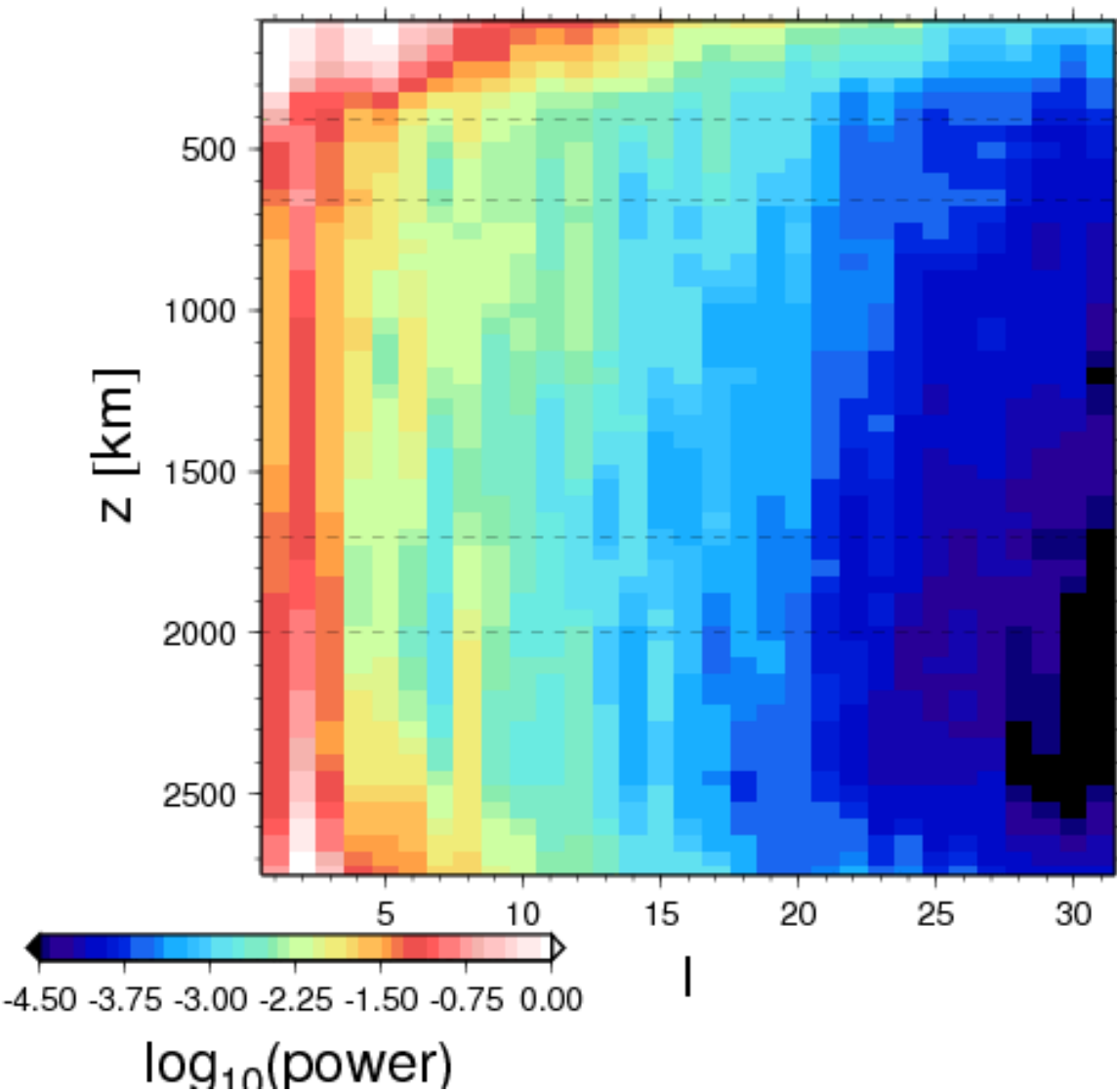
$l=16$



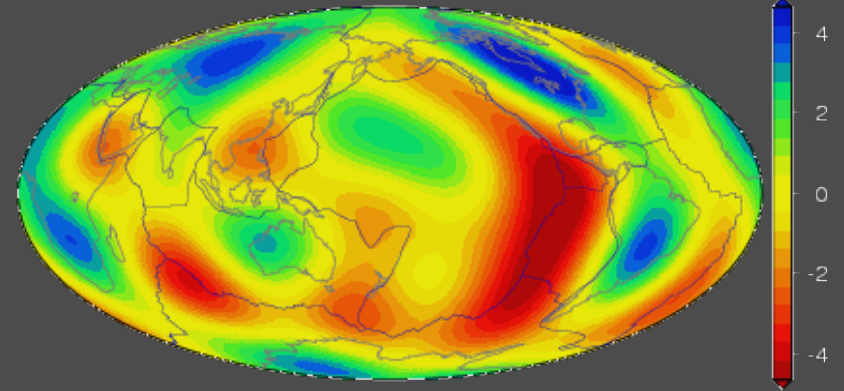
$l=128$



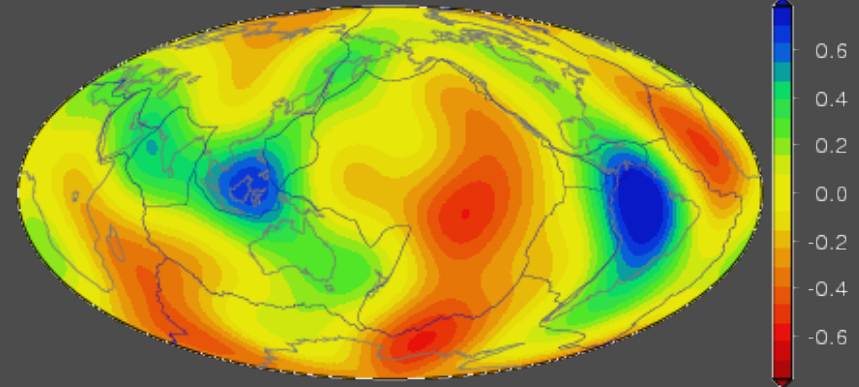
Power spectra



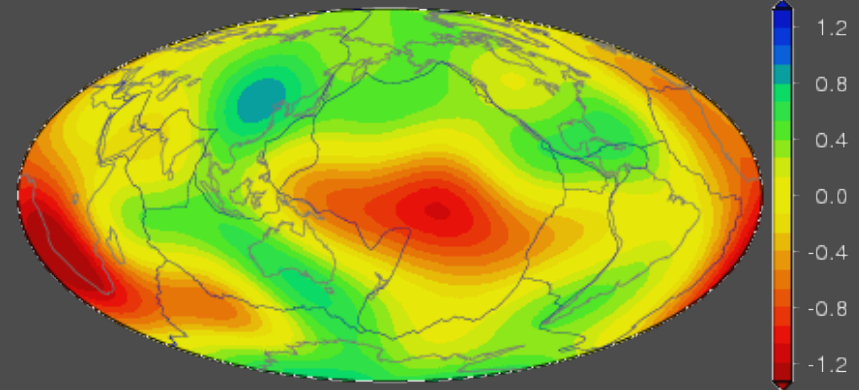
$z = 100 \text{ km}$



$z = 1000 \text{ km}$



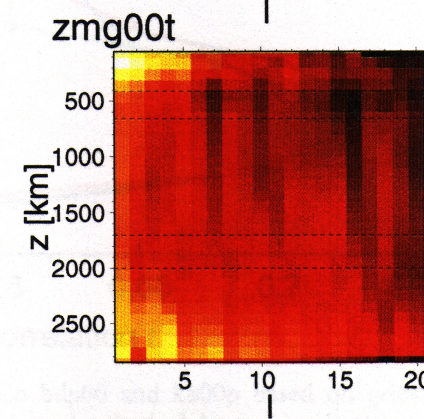
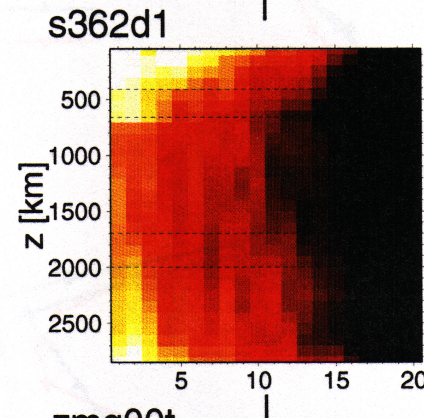
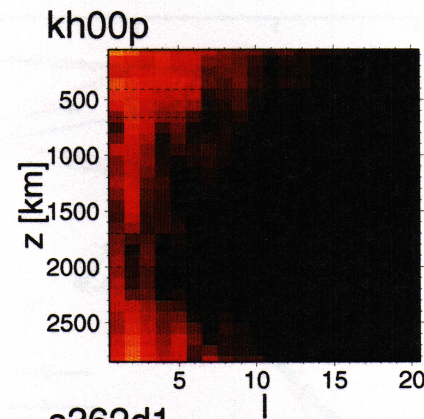
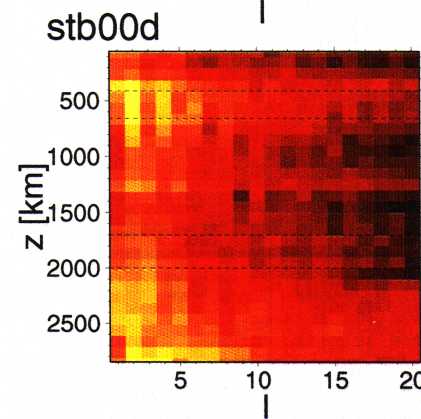
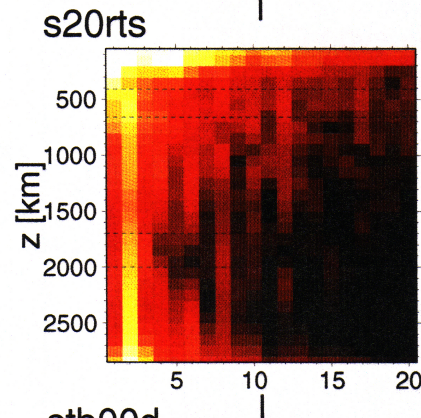
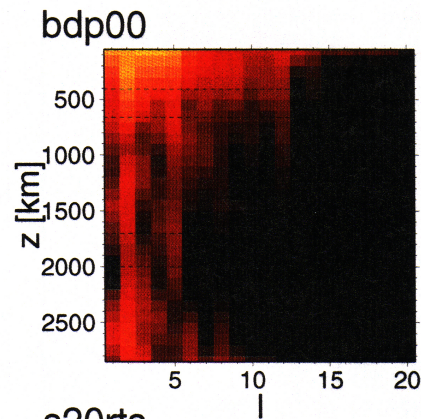
$z = 2500 \text{ km}$



P-HVD
(ISC - P, pP,...)
(Boschi et al.)

S-(S, surface
waves, modes)
(Ritsema et al.)

dyn-slablet
(Steinberger)



P-(ISC P,pP,...)
(Karason et al.)

S- (S, Love,
Rayleigh)
(Gu et al.)

dyn-convection
(Zhong et al.)

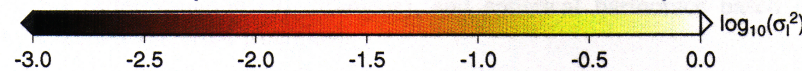


Figure C2. Absolute power per degree and unit area on a logarithmic scale, $\log_{10}(\sigma_l^2)$, for P , S , and geodynamic models (compare Figure 1 and additional online material).